



# NHSC/PACS Web Tutorials

## Running the PACS Spectrometer pipeline for CHOP/NOD Mode

### PACS-301

#### *Pipeline Level 0 to 1 processing*

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Updated by Babar Ali, February 2013

Updated by Steve Lord, Oct 2013

Updated for HIPE 12.0.0 by David Shupe, May 2014

Updated for HIPE 13.0.0 by Roberta Paladini, June 2015



# Introduction

This tutorial will guide you through the interactive spectrometer pipeline from loading raw data into HIPE to obtain calibrated data with astrometry in the case of chop/nod mode.

## Pre-requisites

The following tutorials should be read before this one:

- **PACS-101:** *How to use these tutorials.*
- **PACS-102:** *Accessing and storing data from the Herschel Science Archive*
- **PACS-103:** *Loading scripts*

## Sequel: **PACS-302 – Level 1-2 processing**



# Overview

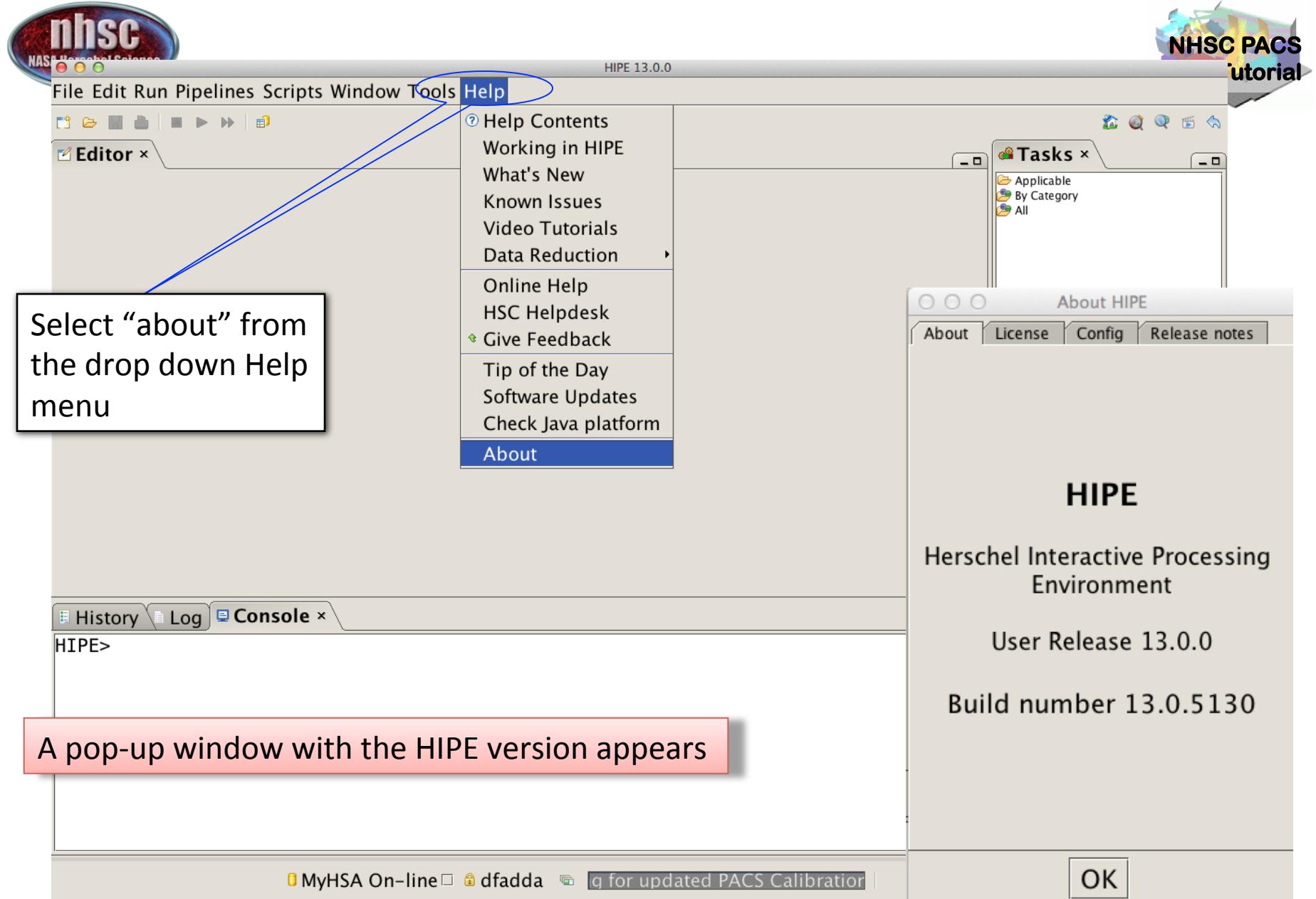
- Step 1** Check HIPE version and your local memory
- Step 2** Set up script for the particular OBSID
- Step 3** Run the  $0 \rightarrow 0.5$  pipeline
- Step 4** Run the  $0.5 \rightarrow 1$  pipeline

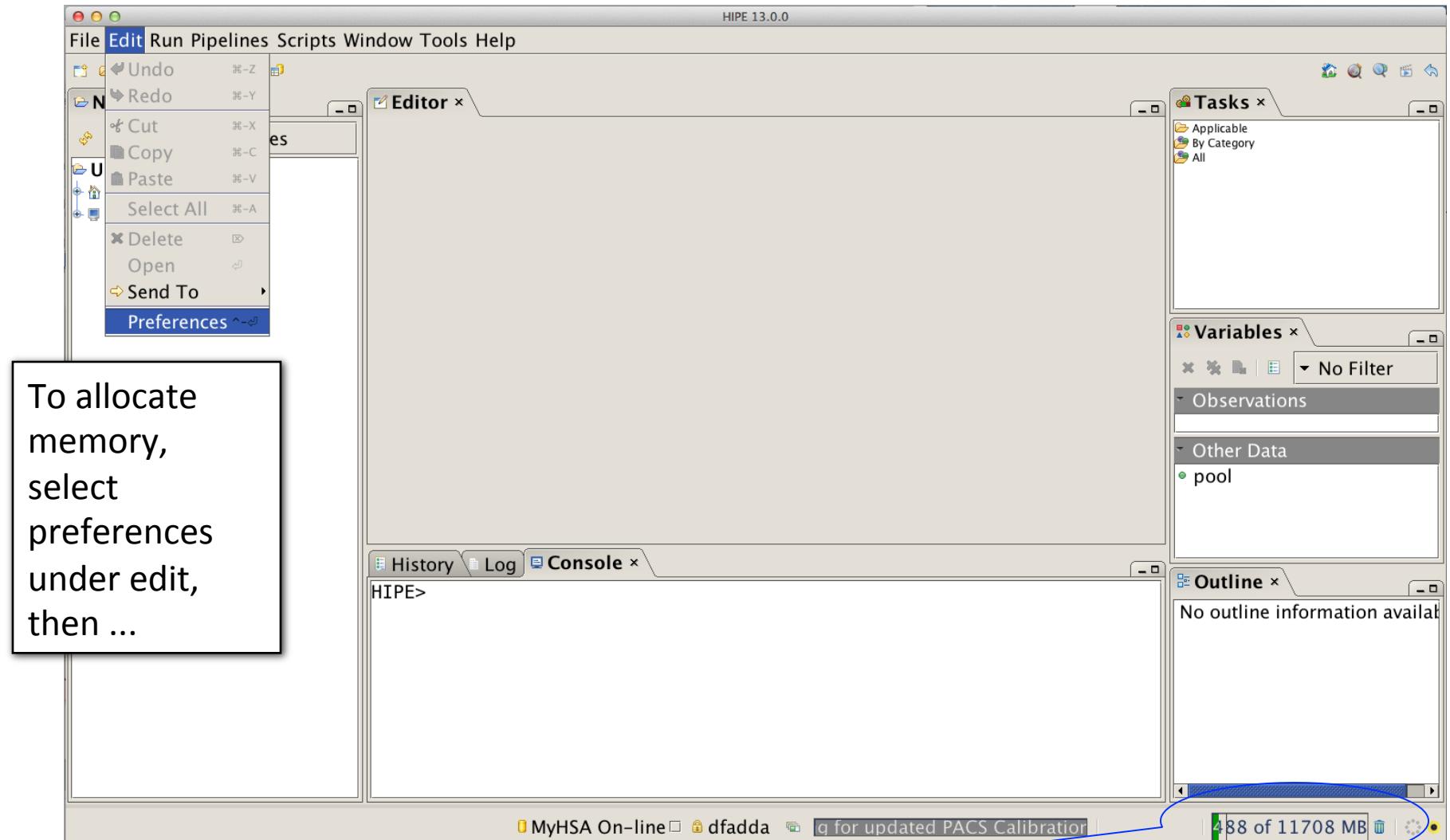


# Step 1

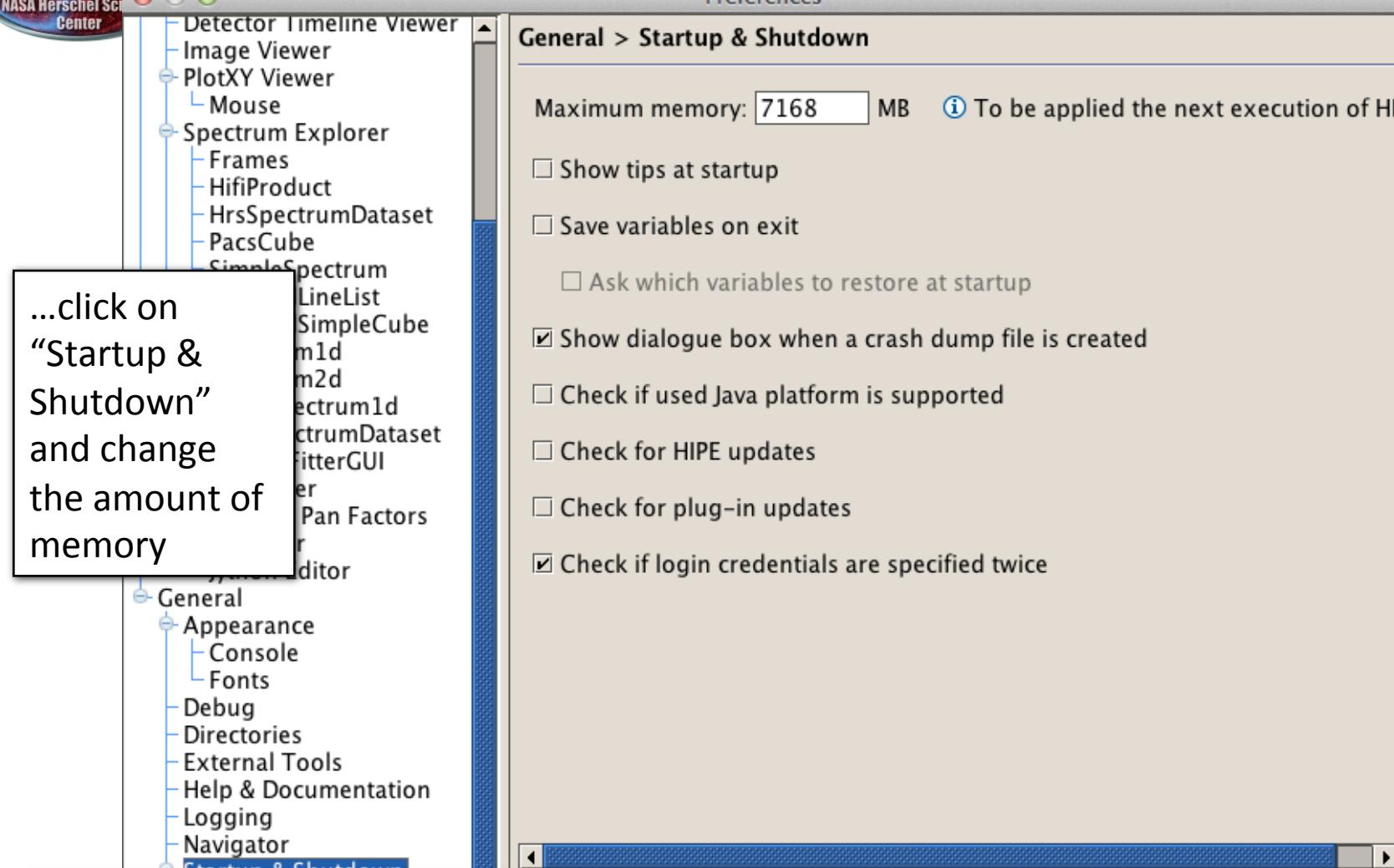
## Check HIPE version and memory allocation

The version used for the tutorial is User Release 13.0.0,  
also known as Build number 13.0.5130





N.b.: Memory used and available



The allocated memory should be a bit smaller than the total RAM of your computer.  
(e.g. 7.5 out of 8.0 Gbytes)

You must exit and restart HIPE to obtain the new amount of memory.



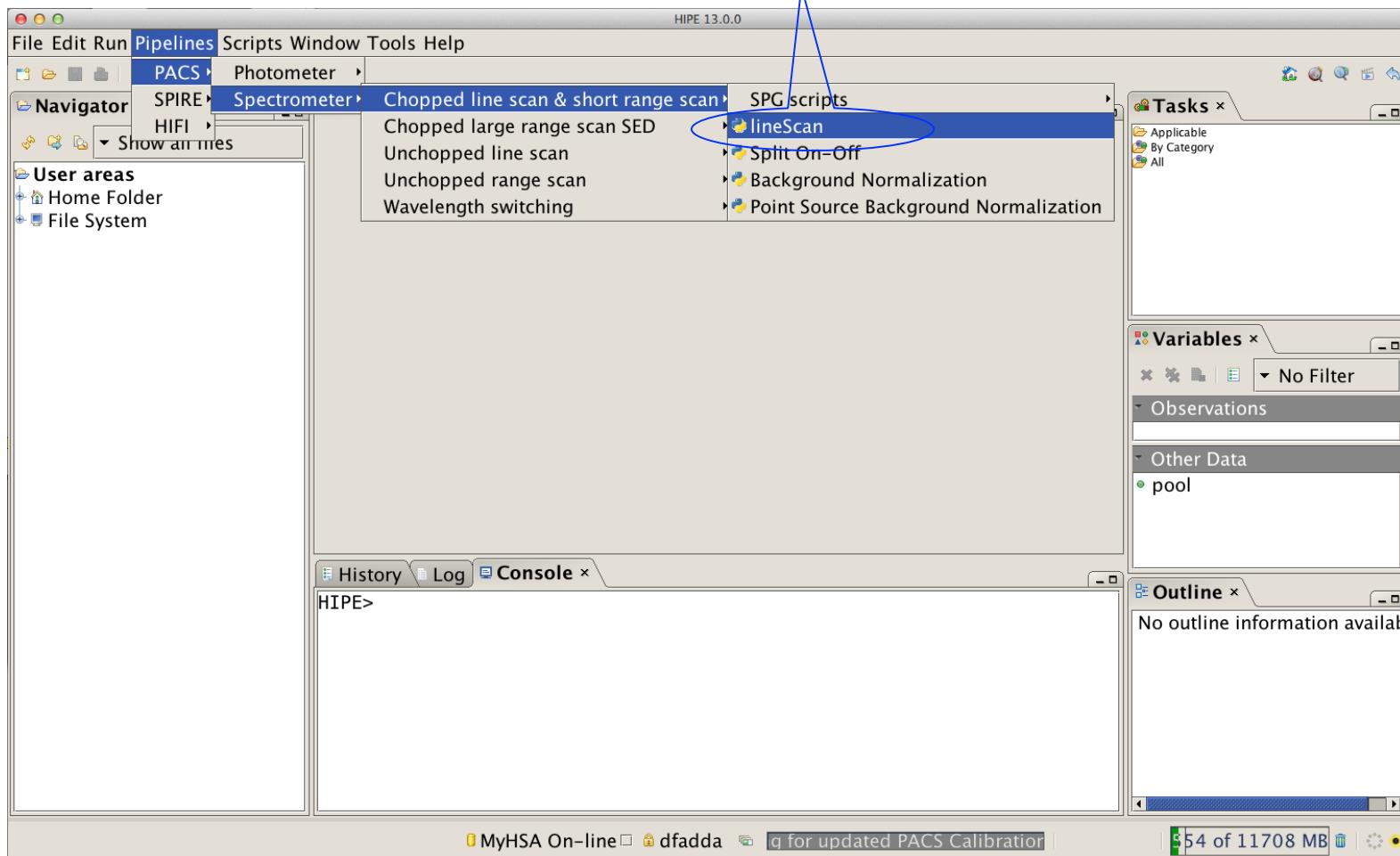
## Step 2

### Setup

Load pipeline script; load observation; check  
your data; and select the camera

# Loading the script

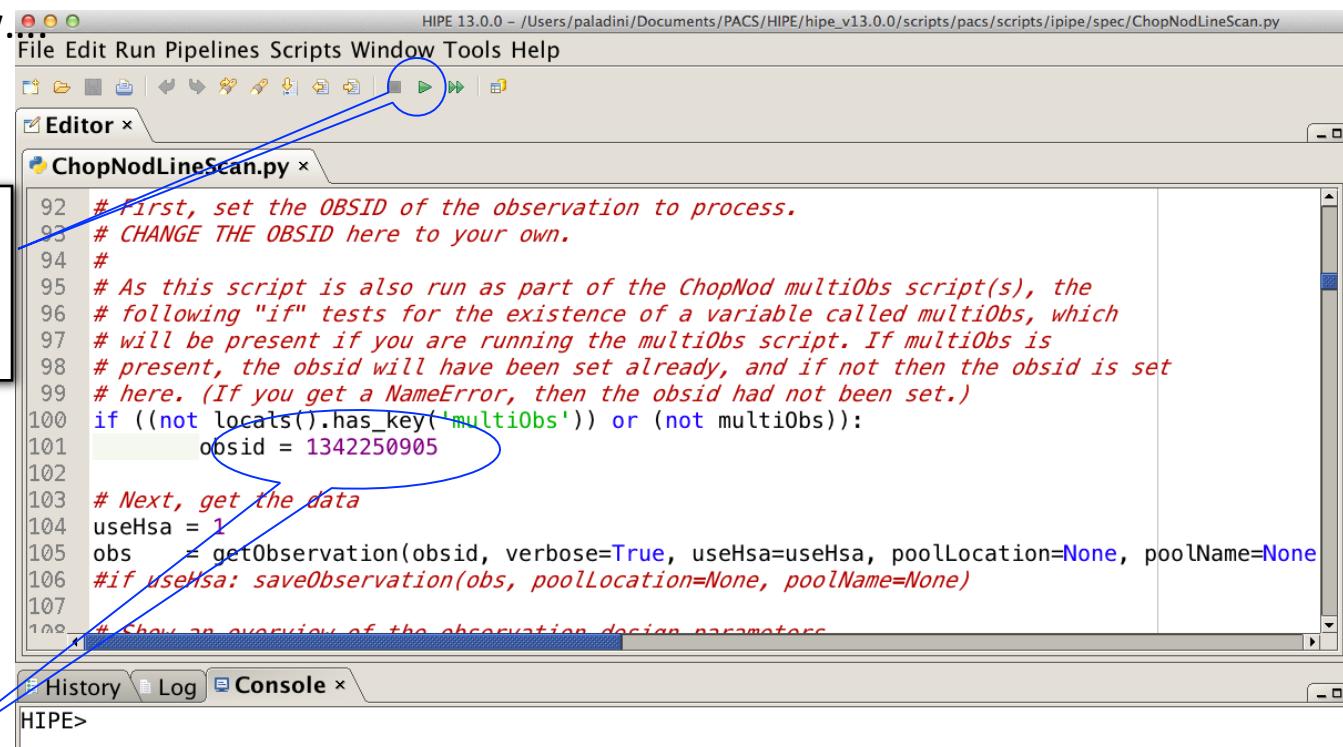
The “linescan” script used in this tutorial corresponds to the script available directly from the distribution.



# Loading the observation

Once the script is loaded, one simply steps through the lines to execute it. But first modify it for OBSID of the observation desired. Modify the obsid in the script and click through using the green arrow... 

Hit the green arrow to step through the entire script



Modify this line. The default  
is obsid 1342250905

# Loading the observation

If the data is not stored as a local pool, you may need to tell getObservation to acquire the data from HSA. In this case, make sure useHsa=1

```
"  
95 # As this script is also run as part of the ChopNod multi0bs script(s), the  
96 # following "if" tests for the existence of a variable called multi0bs, which  
97 # will be present if you are running the multi0bs script. If multi0bs is  
98 # present, the obsid will have been set already, and if not then the obsid is set  
99 # here. (If you get a NameError, then the obsid had not been set.)  
100 if ((not locals().has_key('multi0bs')) or (not multi0bs)):  
101     obsid = 1342250905  
102  
103 # Next, get the data  
104 useHsa = 1  
105 obs    = getObservation(obsid, verbose=True, useHsa=useHsa, poolLocation=None, poolName=None)  
106 #if useHsa: saveObservation(obs, poolLocation=None, poolName=None)  
107  
108 # Show an overview of the observation design parameters  
109 if verbose: obsSummary(obs)  
110
```

# Loading the observation

Next step, we load the observational context ( a structure containing all the observational data, information about them and calibration data).

```

File Edit Run Pipelines Scripts Window Tools Help
Editor x
ChopNodLineScan.py x
94 #
95 # As this script is also run as part of the ChopNod multiObs script(s), the
96 # following "if" tests for the existence of a variable called multiObs, which
97 # will be present if you are running the multiObs script. If multiObs is
98 # present, the obsid will have been set already, and if not then the obsid is set
99 # here. (If you get a NameError, then the obsid had not been set.)
100 if ((not locals().has_key('multiObs')) or (not multiObs)):
101     obsid = 1342250905
102
103 # Next, get the data
104 useHsa = 1
105 obs = getObservation(obsid, verbose=True, useHsa=useHsa, poolLocation=None, poolName=None)
106 #if useHsa: saveObservation(obs, poolLocation=None, poolName=None)
107
108 # Show an overview of the observation design parameters
109 if verbose: obsSummary(obs)
110
111 # Extract the level-0 products from the ObservationContext
112 pacsPropagateMetaKeywords(obs, '0', obs.level0)

```

History Log Console x

```

HYPE> import os,sys
HYPE> verbose = True
HYPE> if ((not locals().has_key('multiObs')) or (not multiObs)):

```

Click through this line using the green arrow.

# The observation summary

**Observation Summary:**

OBSID: 1342250905  
 Instrument: PACS  
 AOR label: RedRectangle-01  
 Proposal: OT1\_vbujarra\_4  
 Target: Red Rectangle  
 Actual RA: 6h 19m 58.27s  
 Actual Dec.: -10° 38' 14.68''  
 Redshift: 0.0 (rad. vel. km/s)  
 Purpose: ---  
 Concat.: ---  
 OD: 1217  
 Start: 2012-09-11T19:17:13.000000 TAI (1726082233000000)  
 Duration: 1315.0 seconds (incl. spacecraft on-target slew time)

You may see a warning from obsSummary – it's not a concern but you can rerun with obsSummary(obs, forceUpdate=True)

**AOT and instrument configuration:**

AOT: PacsLineSpec  
 Mode: Pointed, Chop/Nod  
 Bands: B3A R1 (prime diffraction orders selected)  
 Is bright: NO (default range mode)  
 Chopper: medium throw  
 Nod cycles: 3

Prime lines targeted when the observation was planned

**Observation block summary:**

Channel	Name(*)	Camera	ID	Band(*)	Wave(*)	WaveMin	WaveMax	Repetitions(*)	ActualRep	Capacitance	OutOfBand
0 I 3P1-3P2		blue	2	B3A	micrometer	micrometer	micrometer			pF	
prime		-	102	R1	63.180	62.936	63.441	1	1	0.140	No
parallel		red			189.543	188.776	190.310	1	1	0.140	No

(\*) = requested in HSPOT

**System configuration summary:**

SPG pipeline version: SPG v13.0.0  
 SPG pipeline products creation date: 2015-05-22T22:20:54.680000 TAI (1811024454680000)  
 Mission configuration: MC\_H102ASTR\_P70ASTR\_S66ASTR\_RP  
 Processed to level: RED: 2.0 BLUE: 2.0  
 Quality Control: PENDING  
 Action: NONE

Pipeline version used in making the HSA products

# Retrieving the calibration tree

Then, the calibration tree is loaded.

ChopNodLineScan.py x

```

130 # Set up the calibration tree. We take the most recent calibration files,
131 # for the specific time of your observation (obs=obs)
132 calTree = getCalTree(obs=obs)
133 if verbose:
134     print calTree
135     print calTree.common
136     print calTree.spectrometer
137
138

```

This reads the time stamp of our obs and applies the calibration from the appropriate calibration tree.

History Log Console x

PACS Calibration Tree

Model : FM  
 Scope : BASE  
 Version : 69  
 Branches: [common, photometer, spect

The Cal trees can be accessed and updated from Preferences > Data Access > Pacs Calibration.

PacsCalCommon Calibration Products:

chopperAngle	: FM	print obs.meta["calVersion"] shows the
chopperAngleRedundant	: FM	calibration used in current observation.

# Setting the camera

```
137  
138 # -----  
139 # SELECT DATA FROM ONE CAMERA  
140 # -----  
141  
142 # Red or blue camera ?  
143 if ((not locals().has_key('multi0bs')) or (not multi0bs)):  
144     camera = 'blue'  
145
```

We select camera = 'blue'

After selecting the camera, we can check what camera we selected by simply printing:  
“print camera”

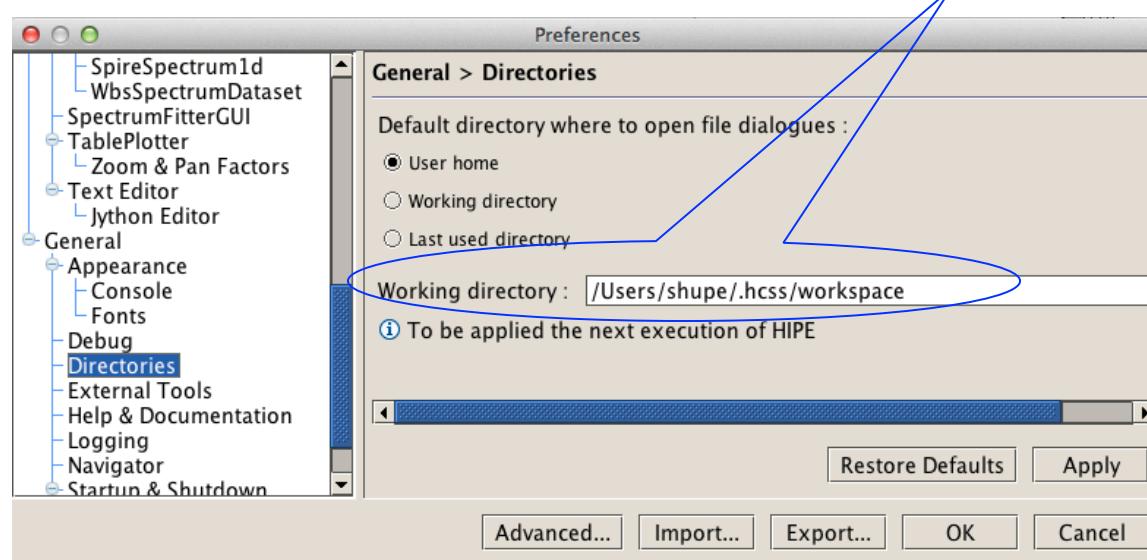
# Setting the output directory

```

169 # saveOutput: False - nothing is saved
170 #           True - the output directory 'outputDir' will be used to store the
171 #                   products of this pipeline (intermediate and final).
172 # When saveOutput is True, nameBasis will be used as basis for the filenames of all outputs
173 saveOutput = True
174
175 # To save to your favourite dir, use : outputDir = "/home/me/myDir/"
176 outputDir = str(Configuration.getWorkDir())+"/pacsSpecOut/"
177 if (not os.path.exists(outputDir)): os.mkdir(outputDir)
178 if verbose and saveOutput: print "The products of this pipeline will be saved in ",outputDir
179
180 # nameBasis will be used as 'basis' for the names of all final fits files
181 nameBasis = str(obsid)+"_"+target+"_"+od+"_Hipe_"+hipeVersion+"_calSet_"+calSet+"_"+camera+"_"+rsrf"
182

```

By default, the script will save intermediate and final products in your HIPE working directory. You can change the HIPE working directory using Edit -> Preferences -> Directories.



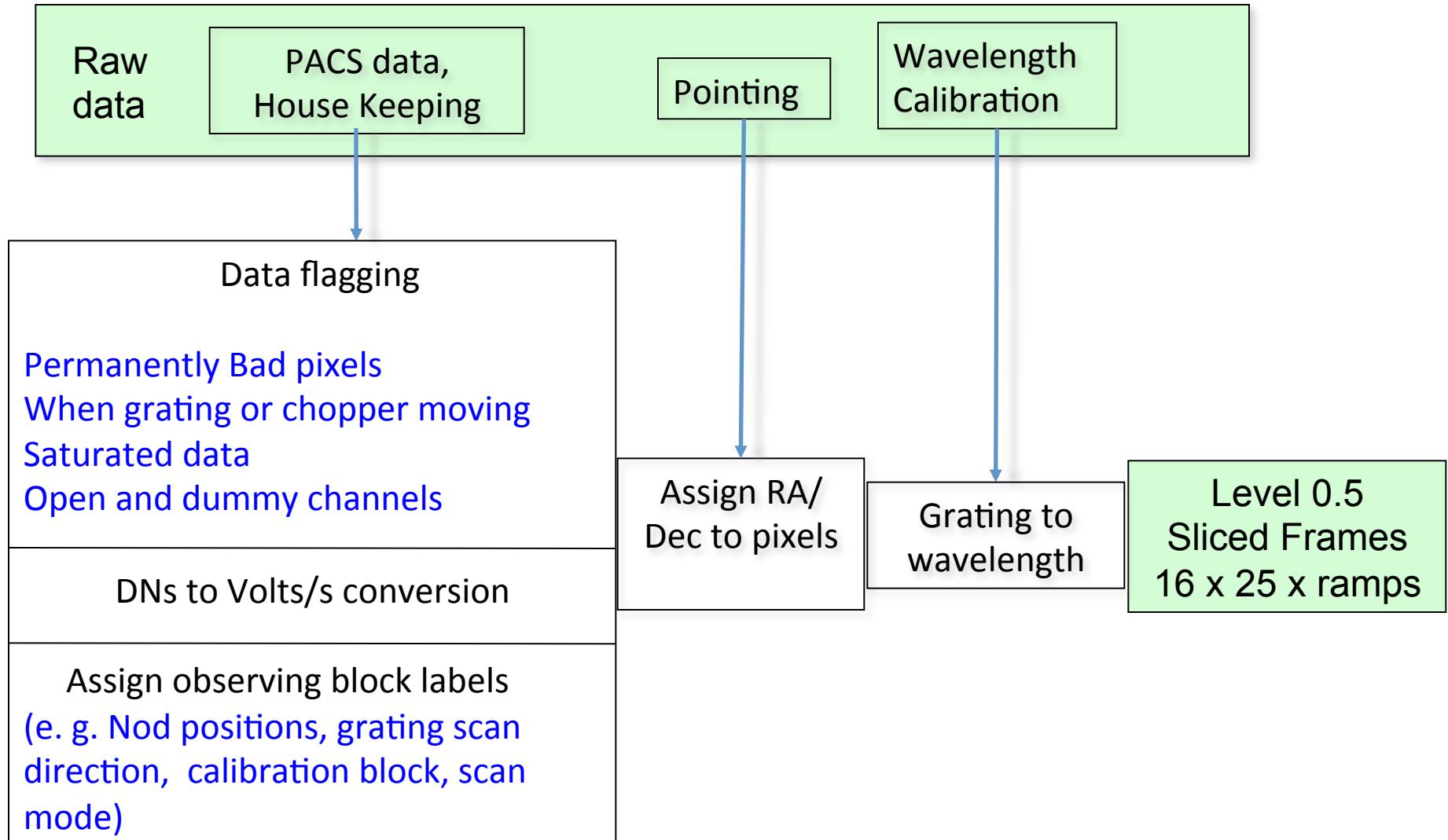


# Step 3

**Run the  $0 \rightarrow 0.5$  pipeline**

Basic calibration (pointing,  
wavelength calibration, slicing)

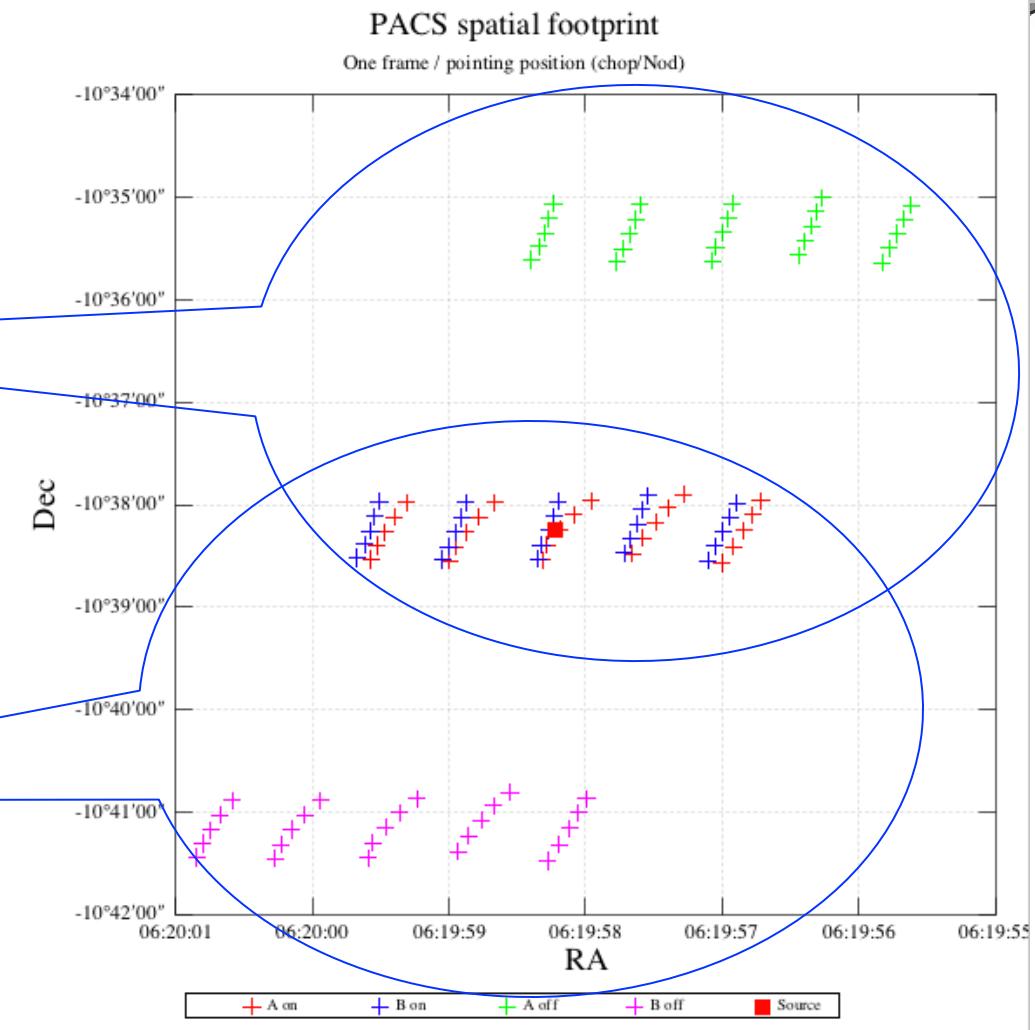
# Level 0 → 0.5



# Plot: footprint

Nod A

Nod B



```

221
222 # Show spatial footprint
223 if verbose: ppoint = slicedPlotPointing(slicedFrames, plotBoresight=False)
224

```

# Slicing into nods

The slicing of the data is performed according to rules made explicit in the pipeline. In our example, one line is observed in two nodding positions and with three repetitions. So, we expect 6 slices plus an initial slice containing the calibration block.

```

248
249 # The internal structure of your data has changed
250 if verbose: slicedSummary(slicedFrames)
251

```

noSlices: 7  
noCalSlices: 1  
noScienceSlices: 6

slice#	isScience	nodPosition	nodCycle	rasterId	lineId	band	dimensions	
							wavelengths	onSource
0	false	["B"]	0	0 0	[1]	["B3A"]	[18,25,679]	
1	true	["B"]	1	0 0	[2]	["B3A"]	[18,25,1631]	
2	true	["A"]	1	0 0	[2]	["B3A"]	[18,25,1631]	
3	true	["A"]	2	0 0	[2]	["B3A"]	[18,25,1631]	
4	true	["B"]	2	0 0	[2]	["B3A"]	[18,25,1631]	
5	true	["B"]	3	0 0	[2]	["B3A"]	[18,25,1631]	
6	true	["A"]	3	0 0	[2]	["B3A"]	[18,25,1631]	
7	true	["B"]	3	0 0	[2]	["B3A"]	[18,25,1631]	

# Check: after slicing

7 slices !

Line 1 – B & A  
nodes – cycle 1

Line 1 – B & A  
nodes – cycle 2

Line 1 – B & A  
nodes – cycle 3

# The internal structure of your data has changed  
if verbose: slicedSummary(slicedFrames)

Console							
slice#	isScience	wavelengths	nodPosition	nodCycle	rasterId	lineId	band
0	false	59.816 – 60.067	["B"]	0	0 0	[1]	["B3A"]
1	true	62.936 – 63.441	["B"]	1	0 0	[2]	["B3A"]
2	true	62.936 – 63.441	["A"]	1	0 0	[2]	["B3A"]
3	true	62.936 – 63.441	["A"]	2	0 0	[2]	["B3A"]
4	true	62.936 – 63.441	["B"]	2	0 0	[2]	["B3A"]
5	true	62.936 – 63.441	["B"]	3	0 0	[2]	["B3A"]
6	true	62.936 – 63.441	["A"]	3	0 0	[2]	["B3A"]

# Continue ...

With remaining Level 0 to 0.5  
processing steps as outlined in slide 18.  
Step through with the green arrow.

```
# -----  
#       Processing      Level 0.5 -> Level 1  
# -----
```

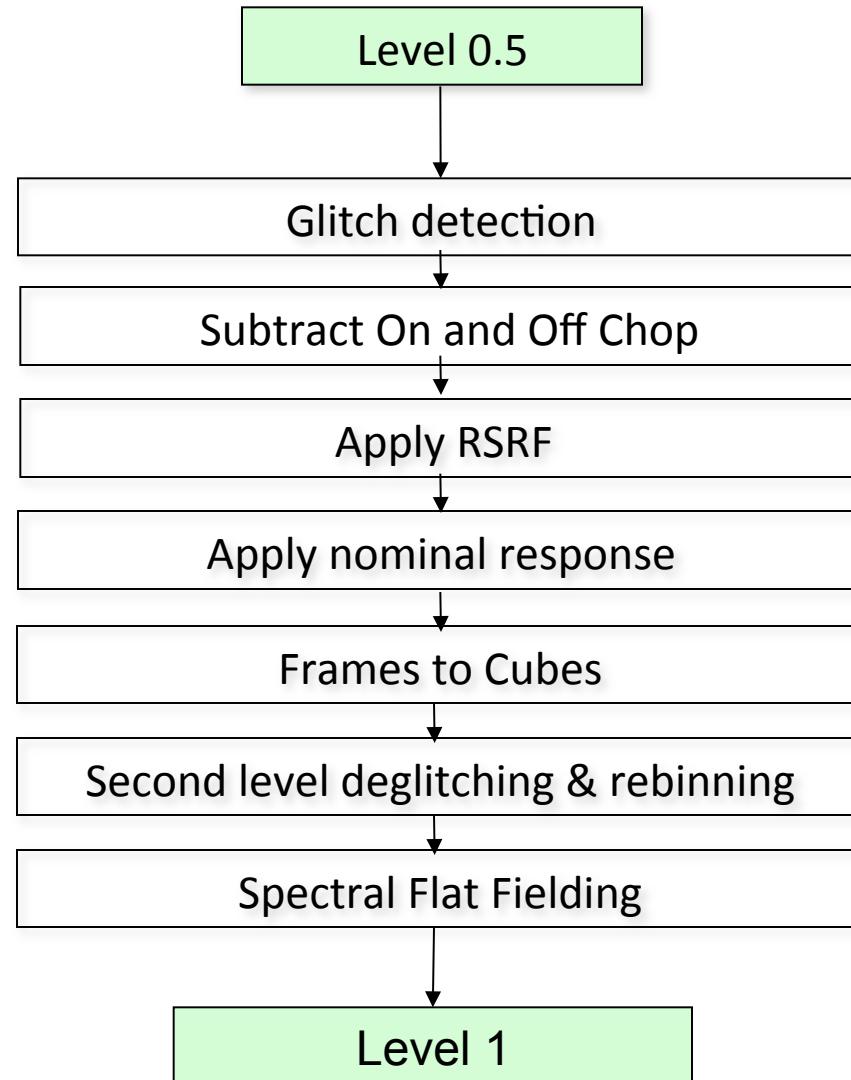


# Step 4

## Run the $0.5 \rightarrow 1$ pipeline

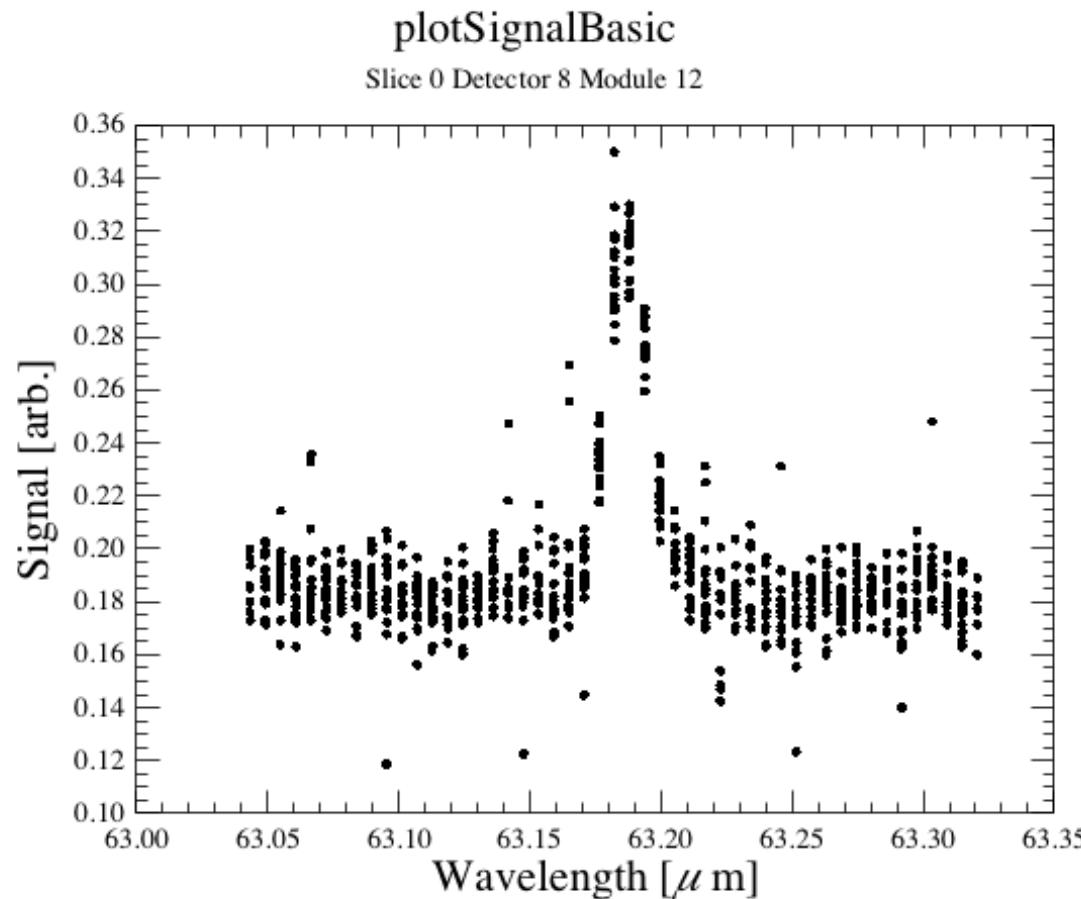
Glitch detection, chop differentiation, RSRF, flat

# Level 0.5 → 1

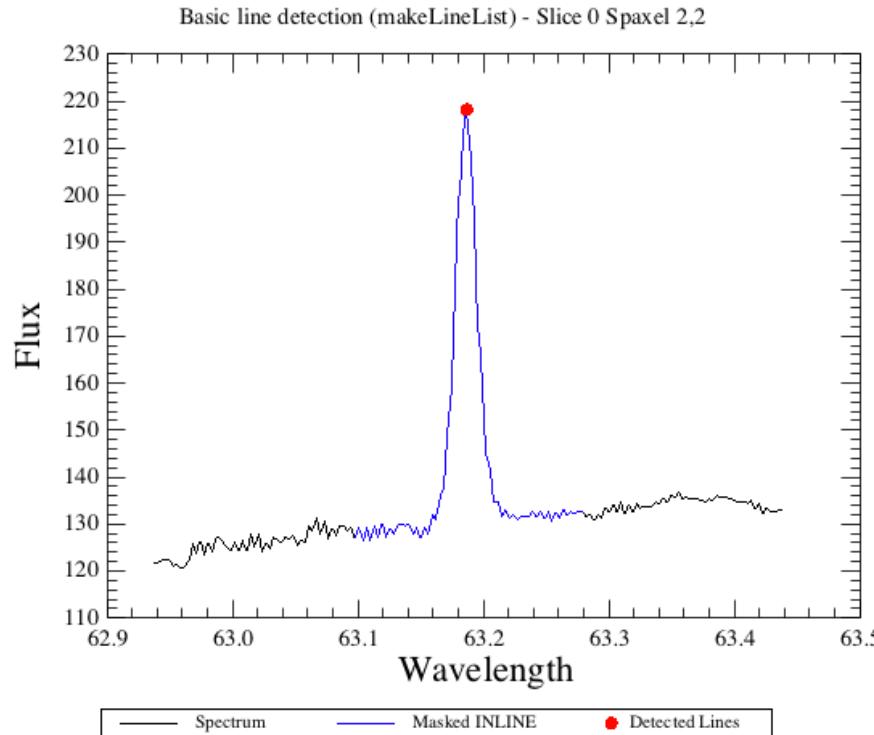


# Signal after chop subtraction

Verbose=1 shows  
The data are only on the ON position (OFF being subtracted)



# Check: Spectral FlatField

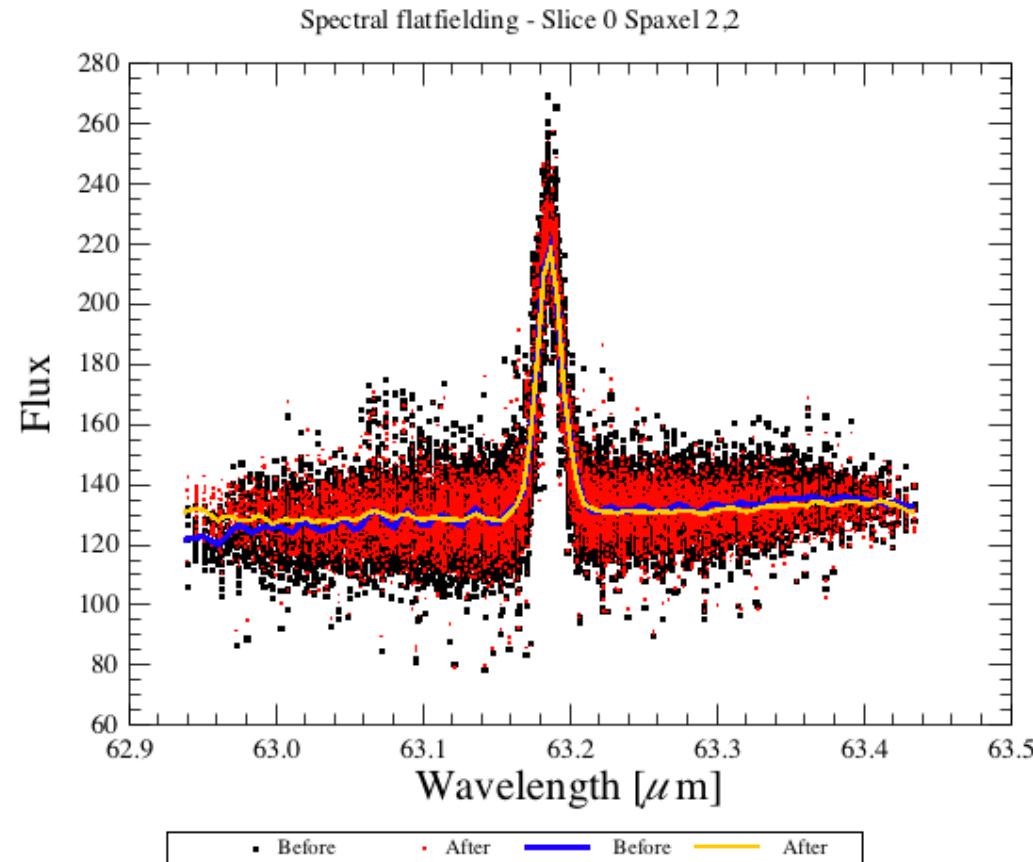


As a default, the code will search for lines in all the pixels and then mask them before computing the spectral flat field.

It is possible to give directly the list of lines to be masked via the parameter `lineList = [63.187]`, for instance.

This user-specified `lineList` is usually needed *only* for absorption lines.

# Check: Spectral FlatField



At this stage you will just want to check that the red “After” points have a tighter distribution (less scatter) than the black “Before” points.



# You are ready to continue with **PACS-302**

```
341 # 3. Actual spectral flatfielding
342 # slopeInContinuum is a boolean. Set it to true for lines existing on a continuum with a significant
343 slopeInContinuum = 1
344 slicedCubes = specFlatFieldLine(slicedCubesMask, scaling=1, copy=1, maxrange=[50.,230.], slopeInConti
345
346 # 4. Rename mask OUTLIERS to OUTLIERS_B4FF (specFlagOutliers would refuse to overwrite OUTLIERS) & de
347 slicedCubes.renameMask("OUTLIERS", "OUTLIERS_B4FF")
348 slicedCubes = deactivateMasks(slicedCubes, String1d(["INLINE", "OUTLIERS_B4FF"]))
349
350 # 5. Remove intermediate results
351 del waveGrid, slicedRebinnedCubes, slicedCubesMask
352
353 # --- End of Spectral Flat Fielding
354
355 #
356 # ----- Processing Level 1 -> Level 2
357 # -----
```