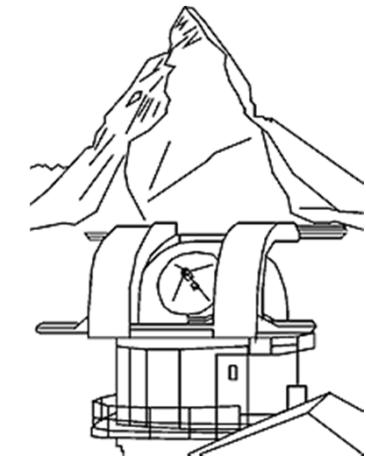


# *Cold Chemistry in Space and Laboratory*

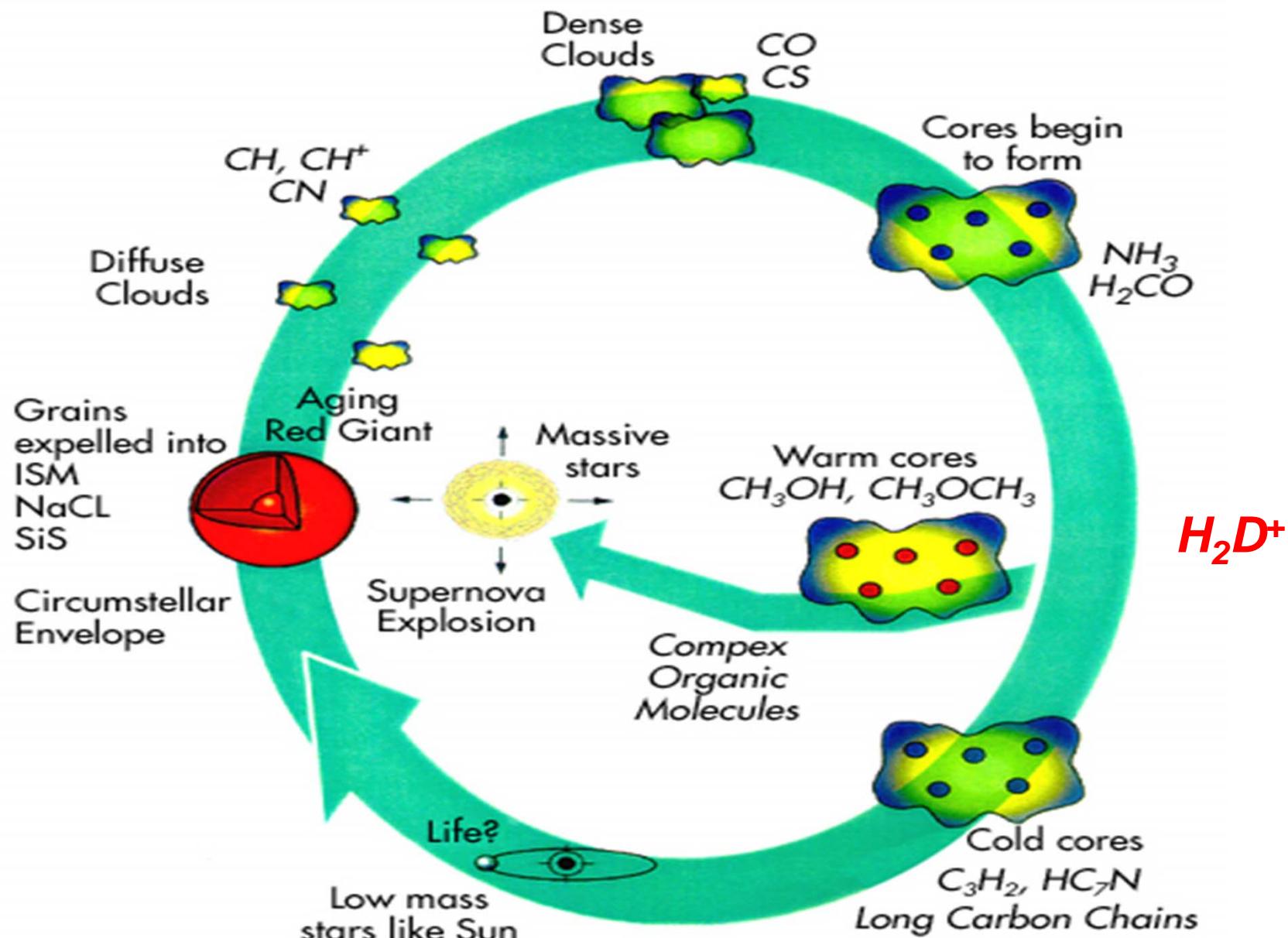
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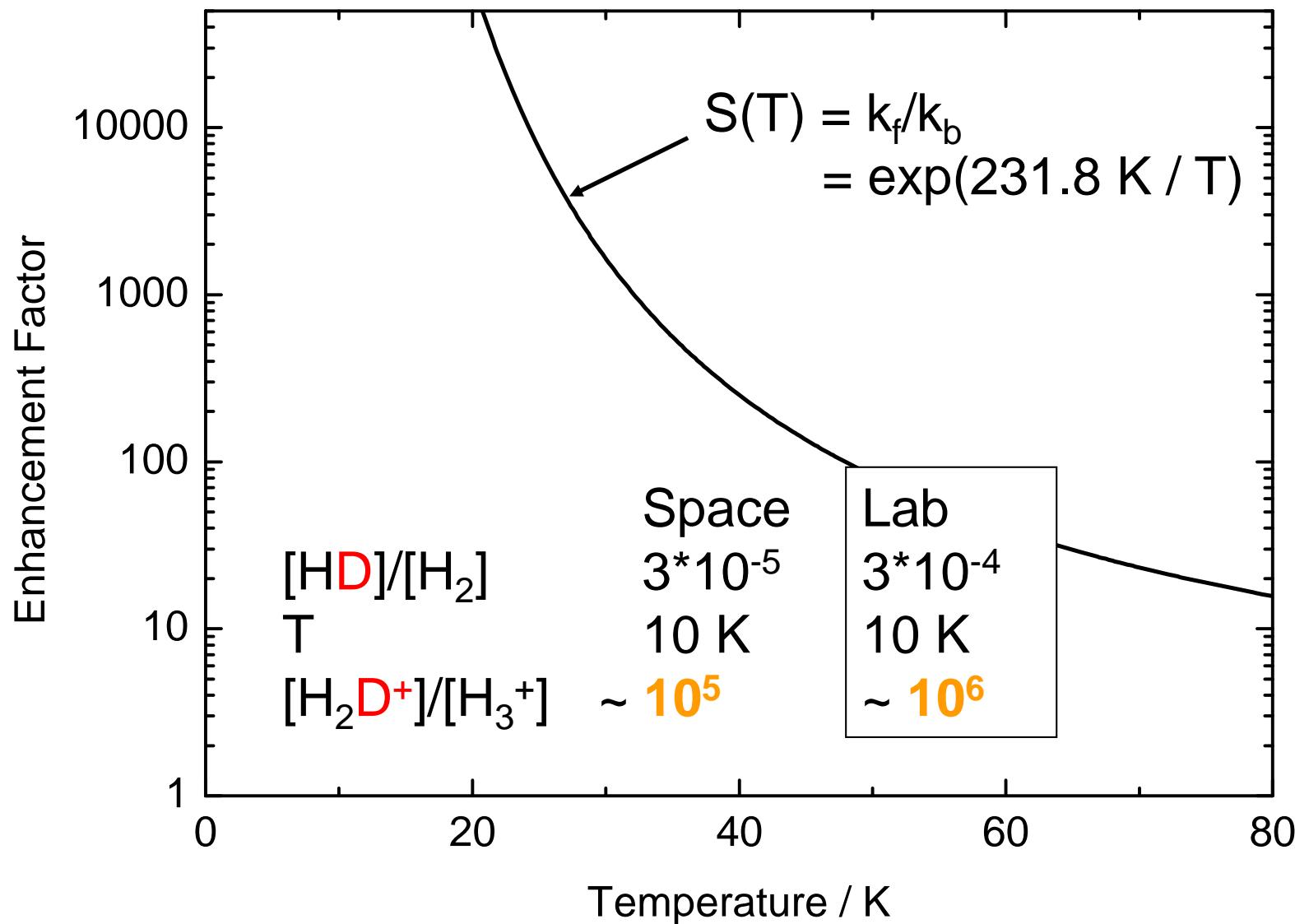


- H<sub>2</sub> Formation, OPR and Chemical Clocks
- H<sub>3</sub><sup>+</sup> / H<sub>2</sub>D<sup>+</sup> Isotopic Fractionation, H<sub>3</sub><sup>+</sup>/H<sub>2</sub>D<sup>+</sup>, OPR
- H<sub>2</sub>D<sup>+</sup> + H<sub>2</sub> THz Spectroscopy in Lab and Space

# Life cycle of Stars



# Isotopic Fractionation



# Experimental Method:

# Electrodynamical Trapping



Sandra Brünken

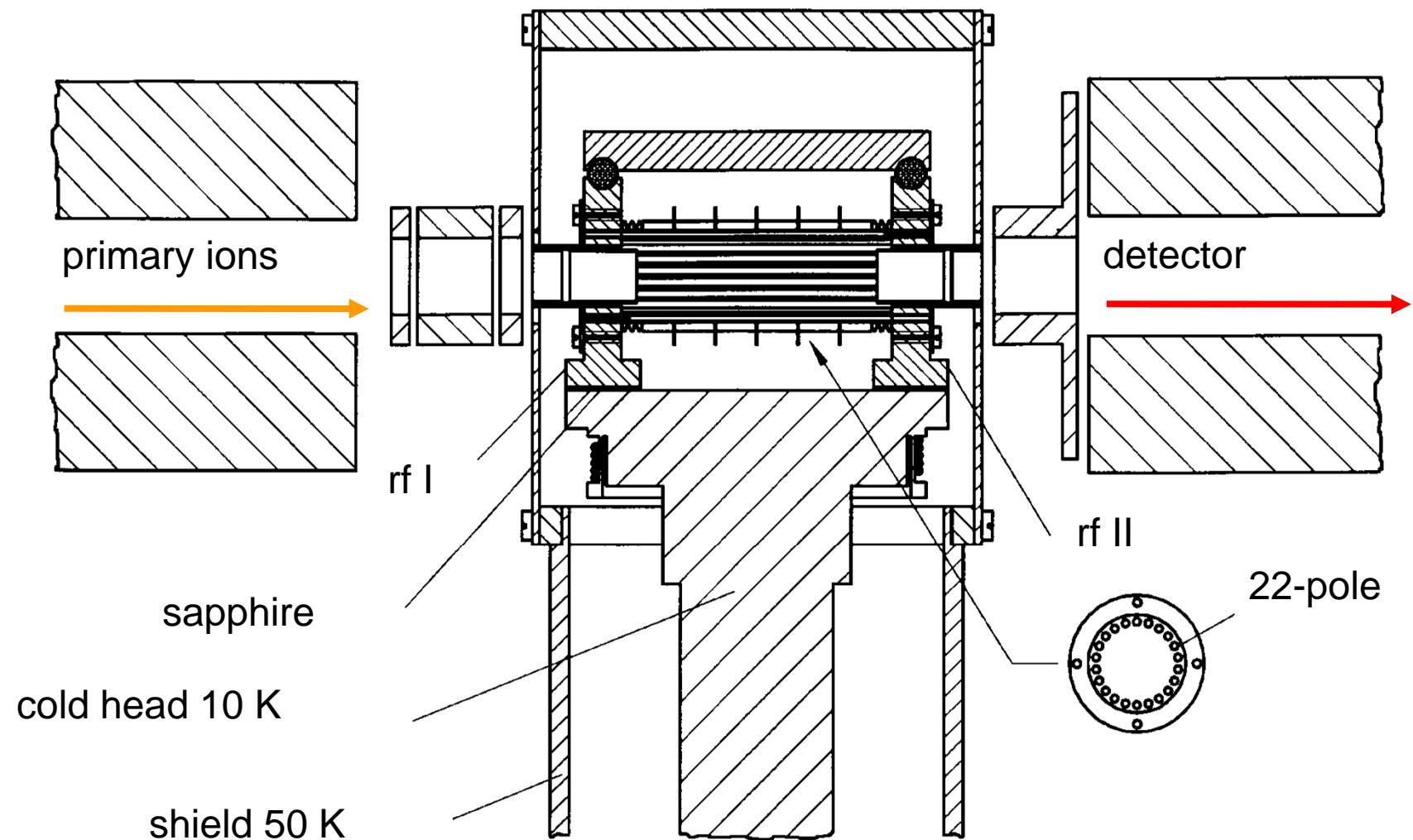


Dieter Gerlich

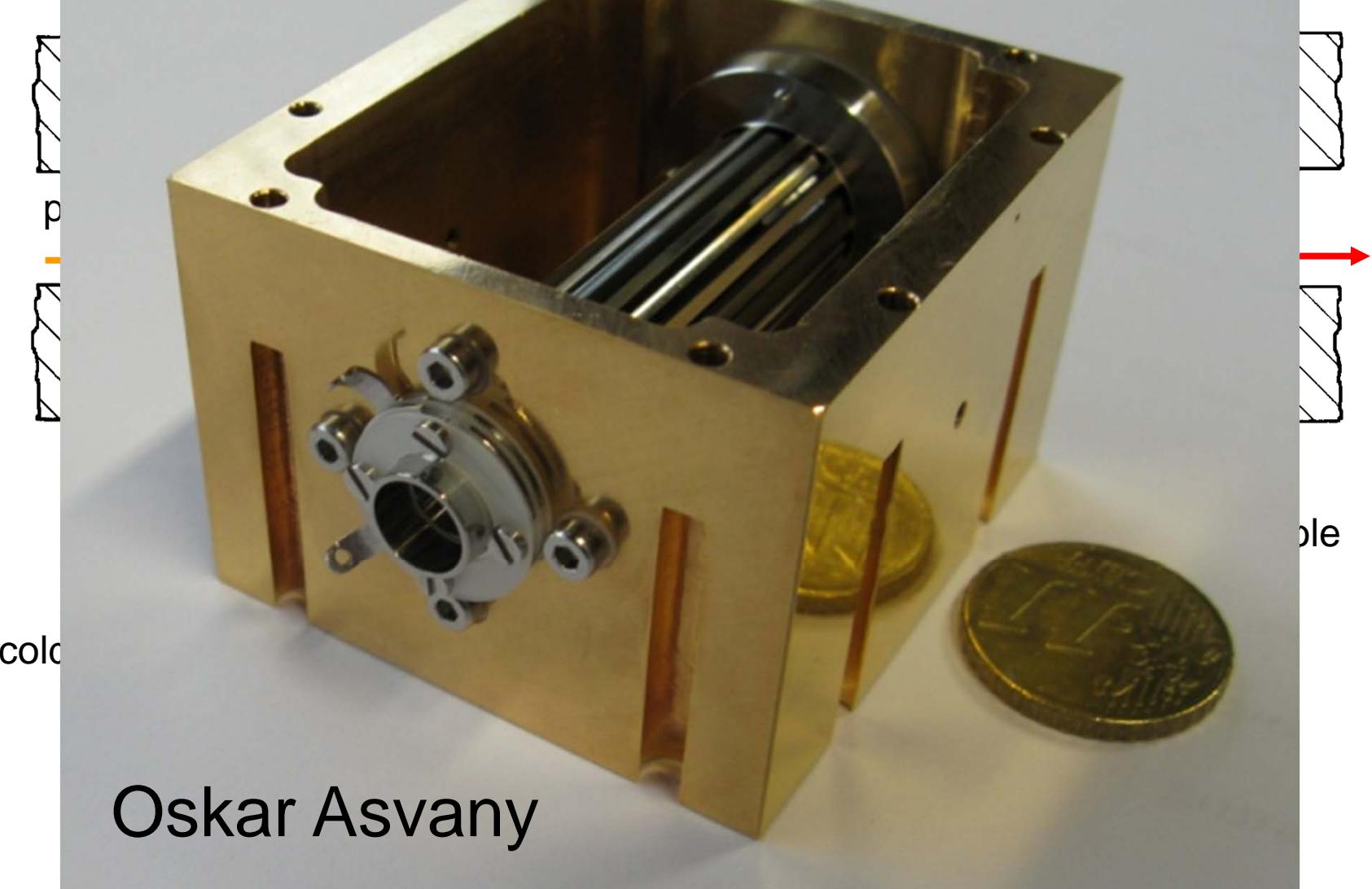


Oskar Asvany

# 22-Pole Low Temperature Ion Trap



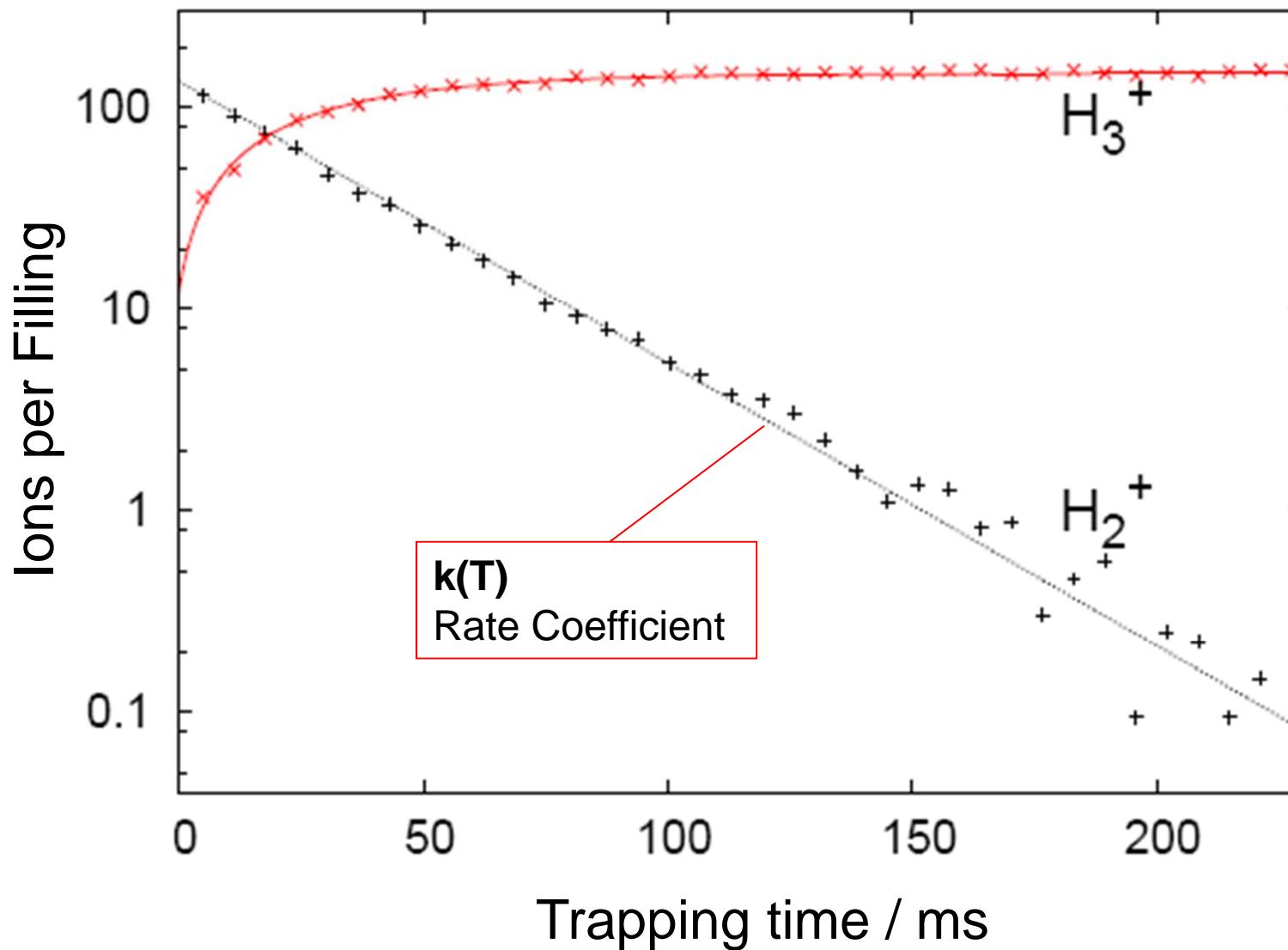
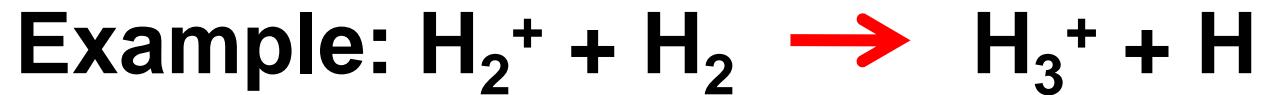
# 22-Pole Low Temperature Ion Trap

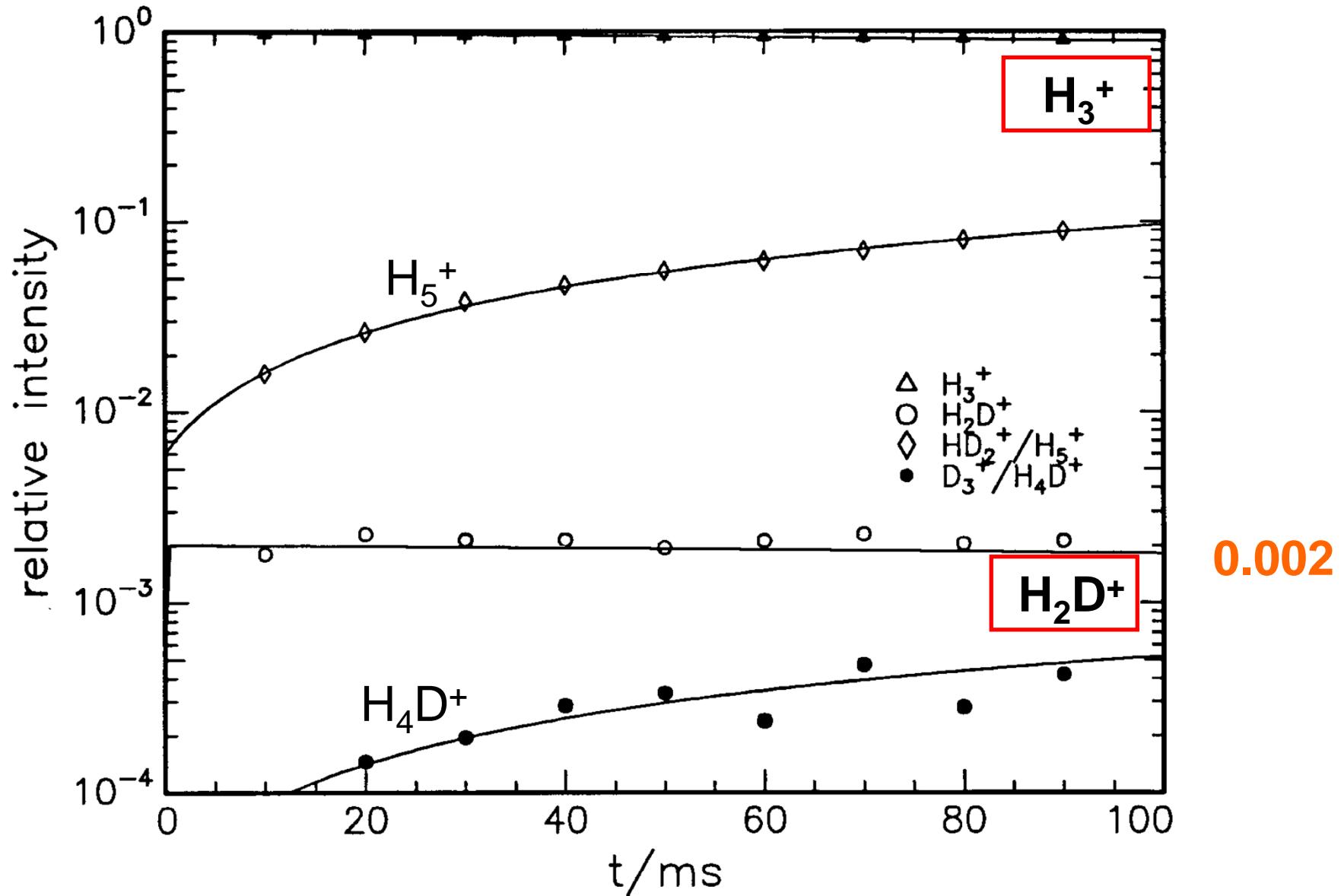
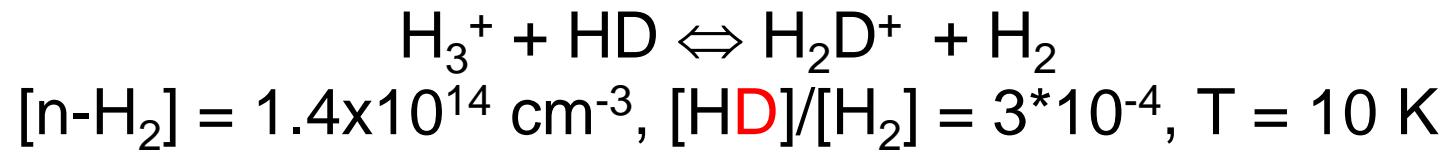


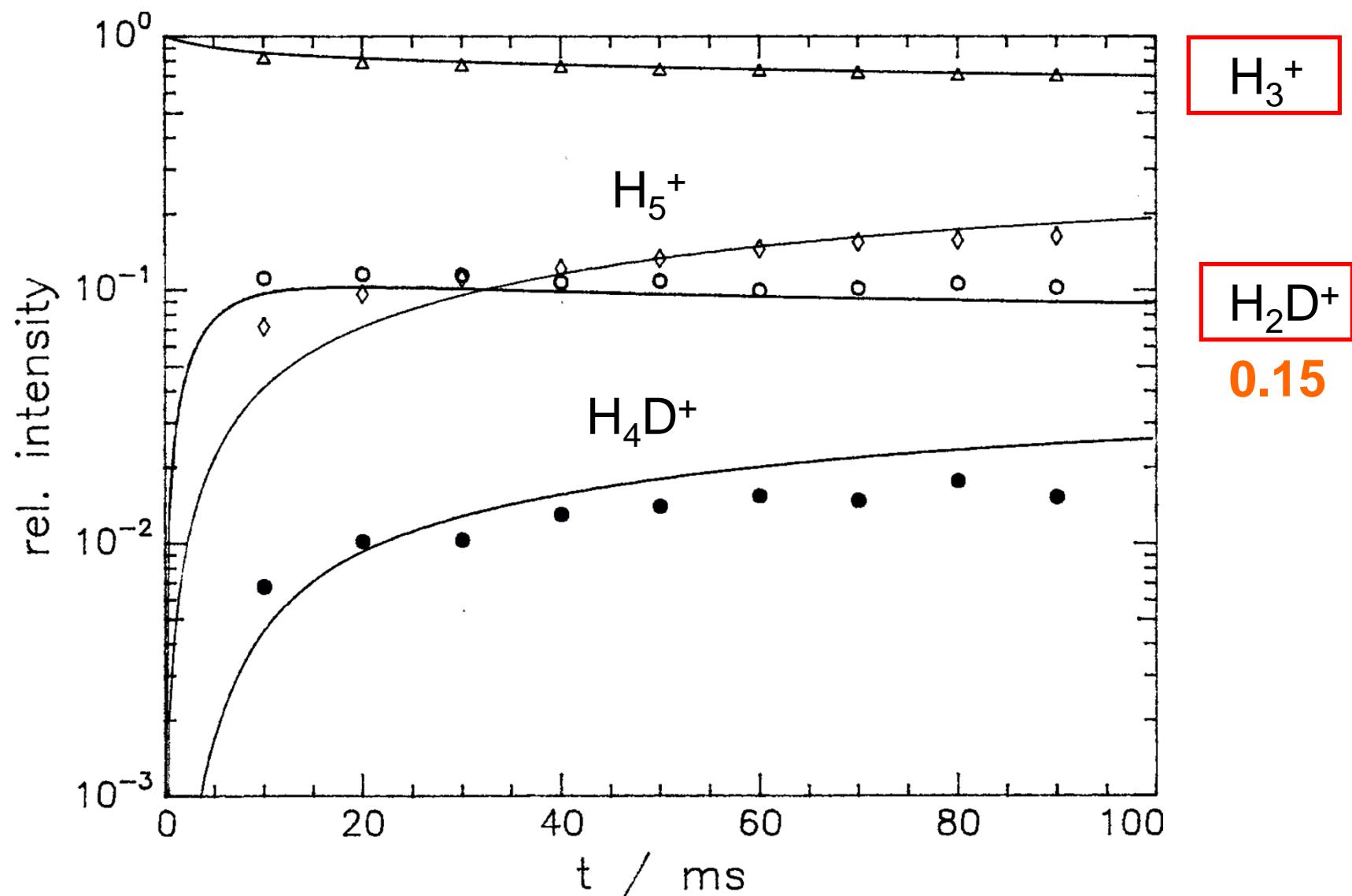
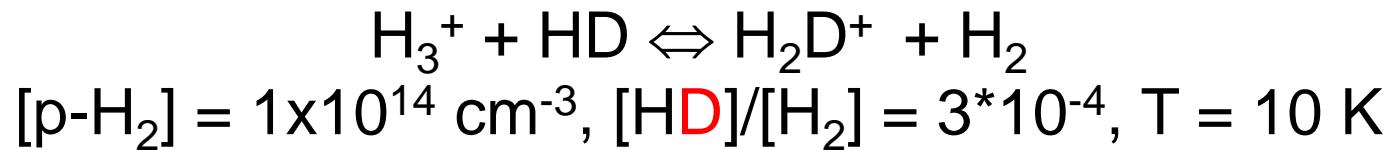
color

ole

Oskar Asvany

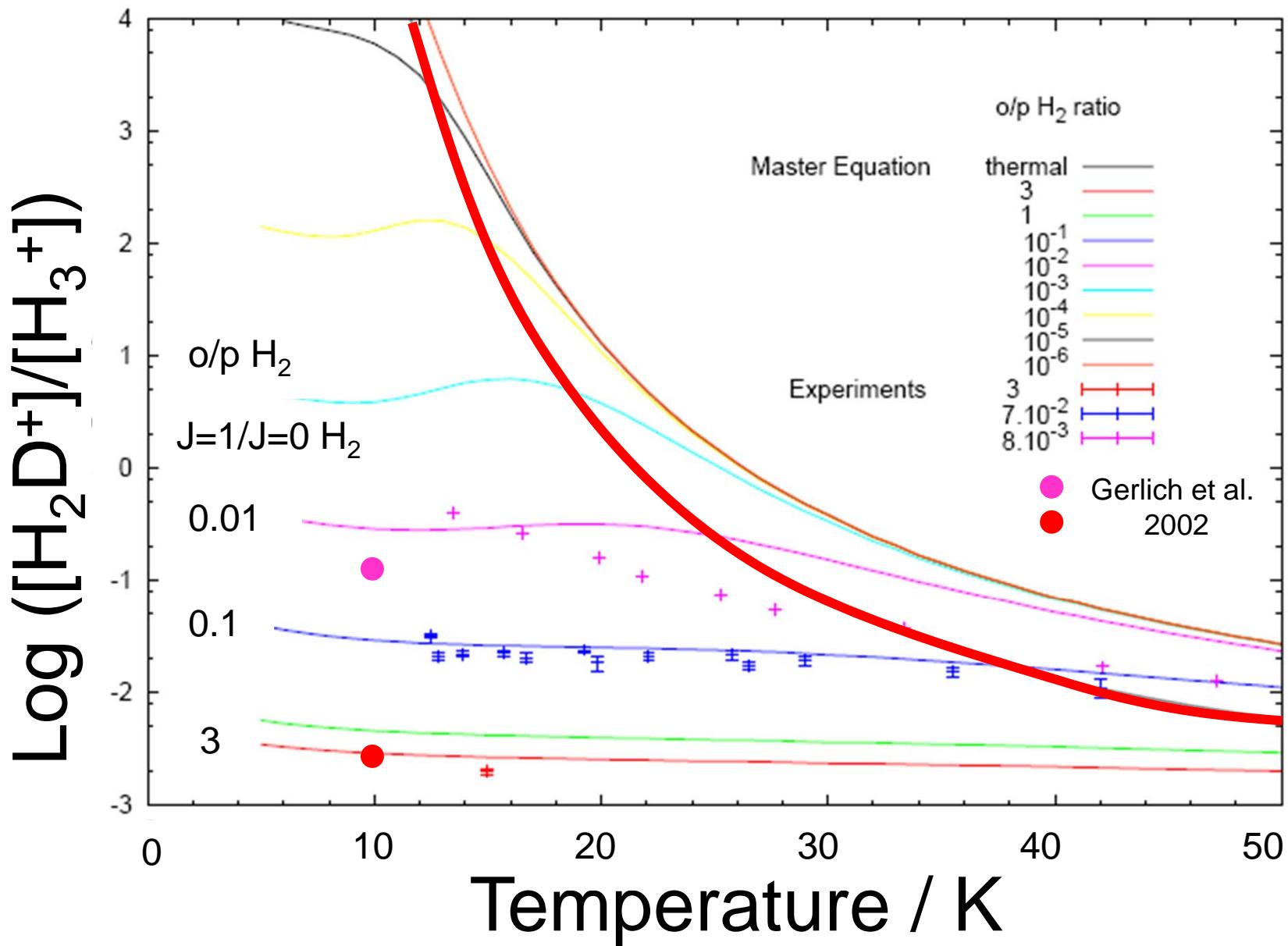




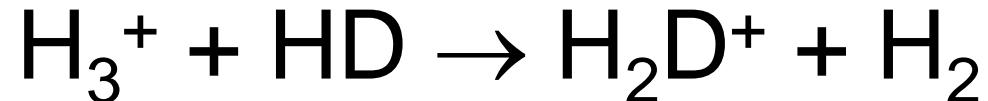


# Experimental Results & Modelling

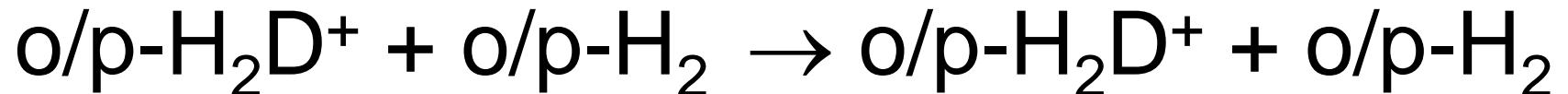
Hugo et al., J.Chem.Phys. 2009, **130**, 164302

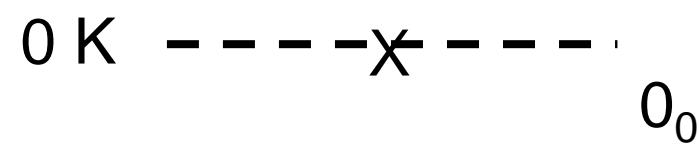
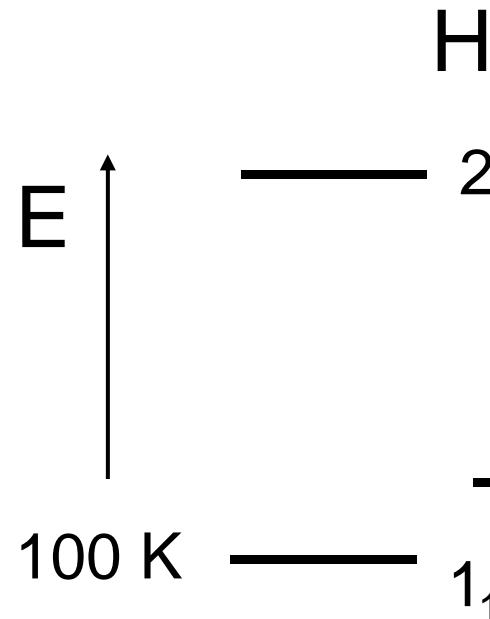


# Isotopic Fractionation



and the  $\text{H}_2 / \text{H}_2\text{D}^+$  OPR

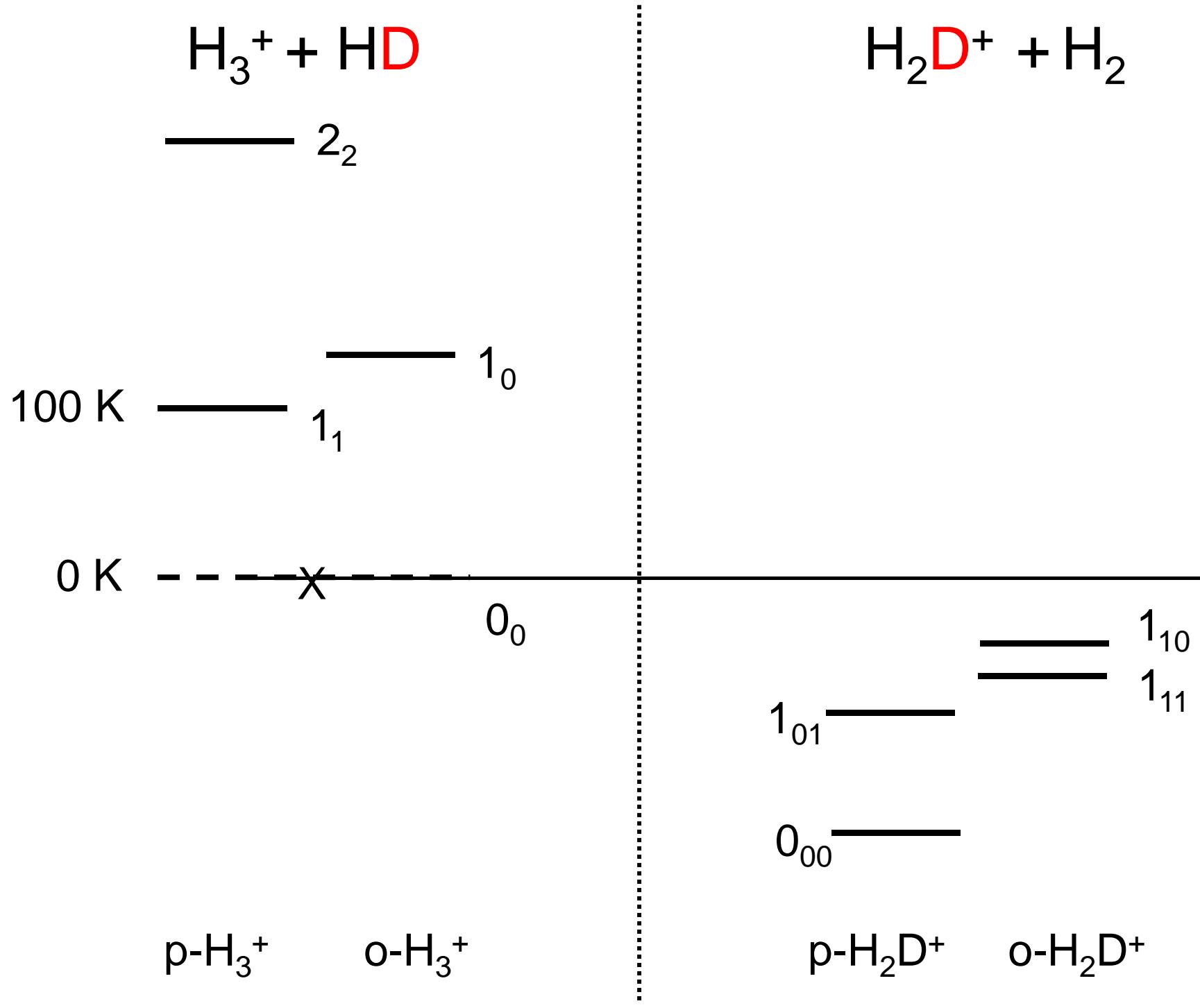


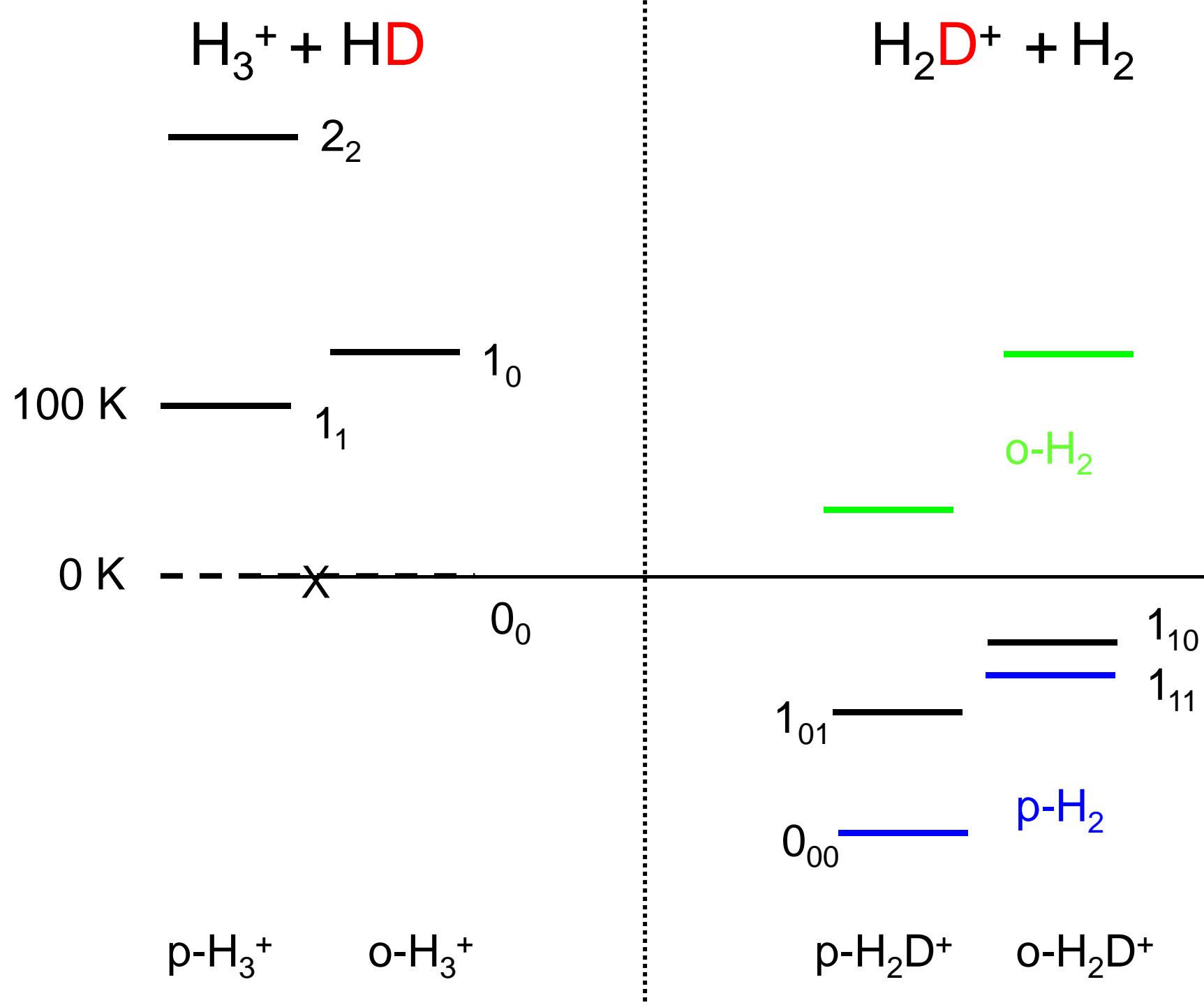


p- $\text{H}_3^+$

o- $\text{H}_3^+$

Lowest energy  
levels of  $\text{H}_3^+$

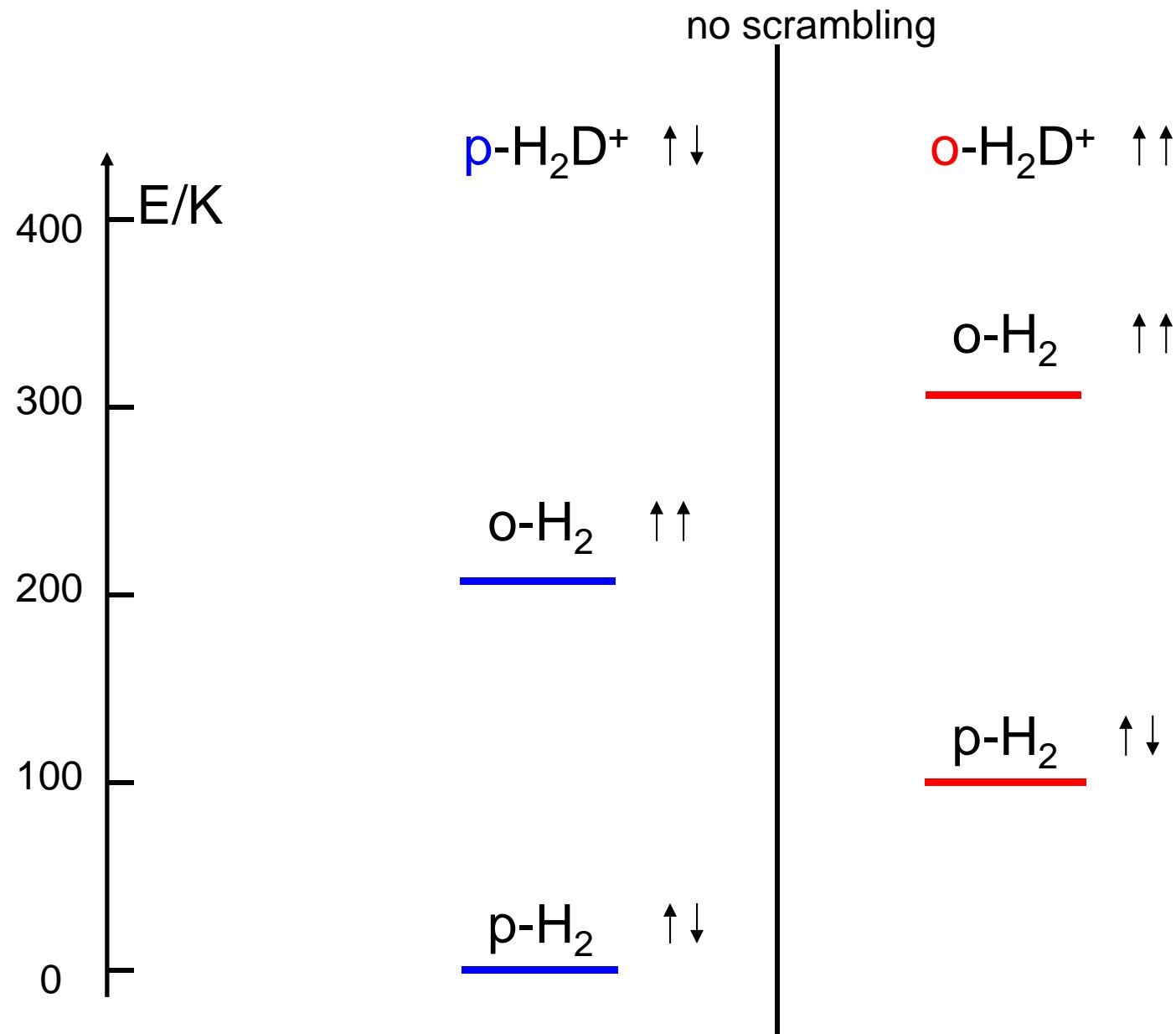




# *Role of Nuclear Spin?*

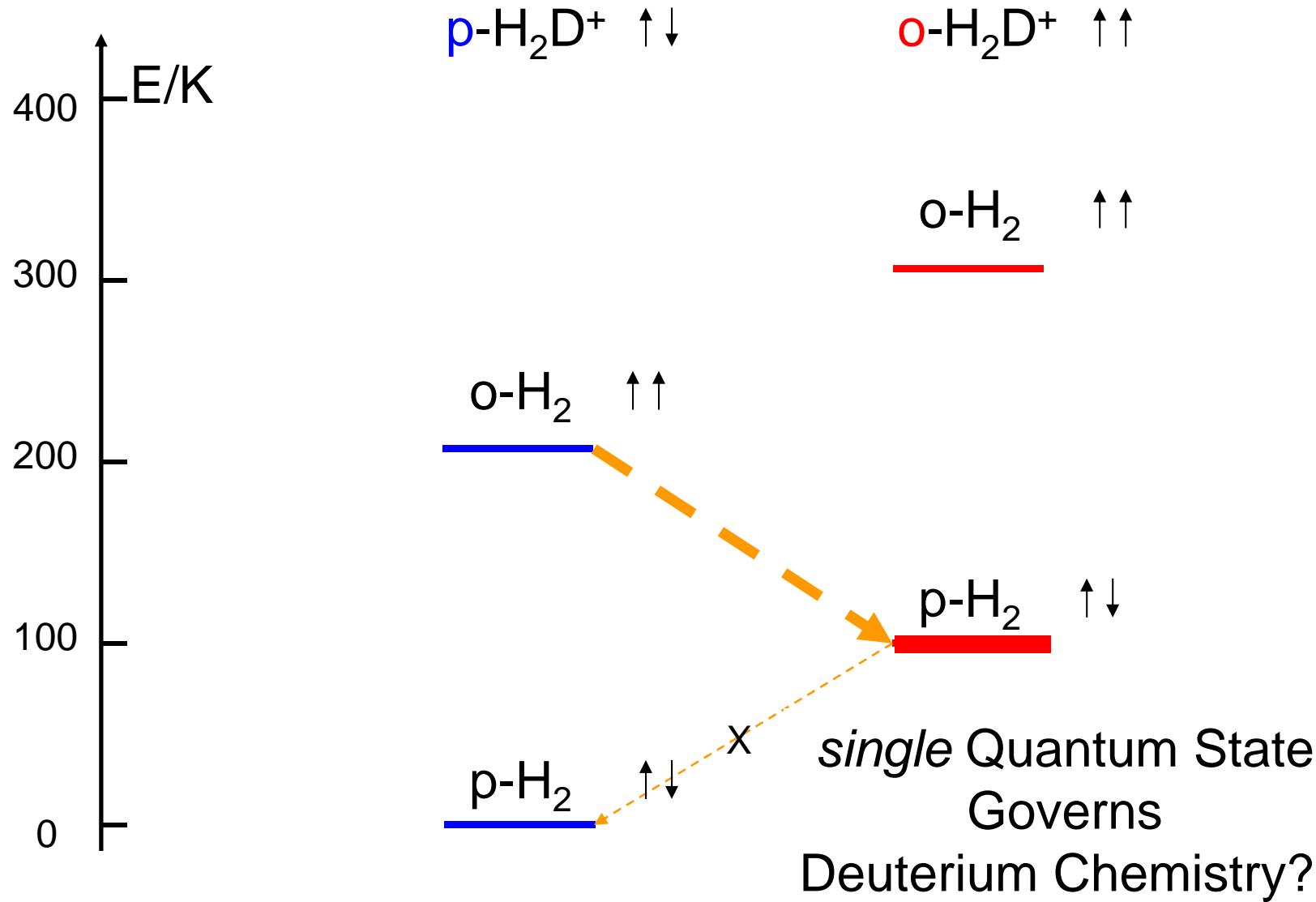
Conservation laws: E, J, P, I, ...

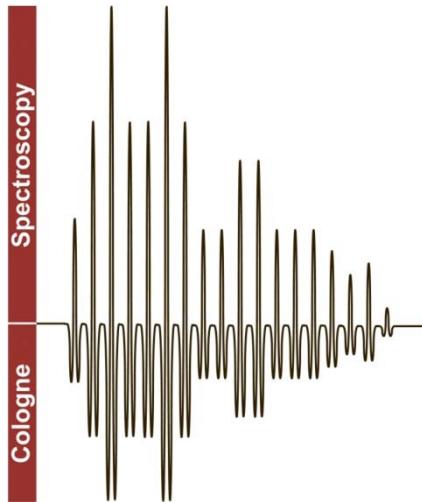
# Para – Ortho Conversion



# Para – Ortho Conversion

scrambling

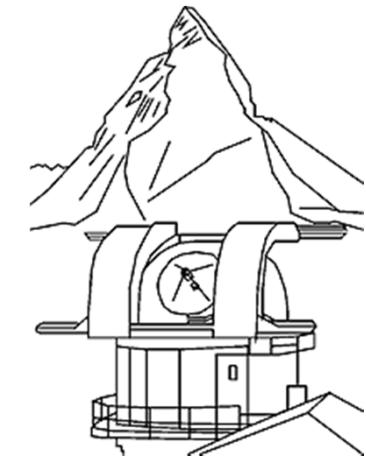




# *Cold Chemistry in Space and Laboratory*

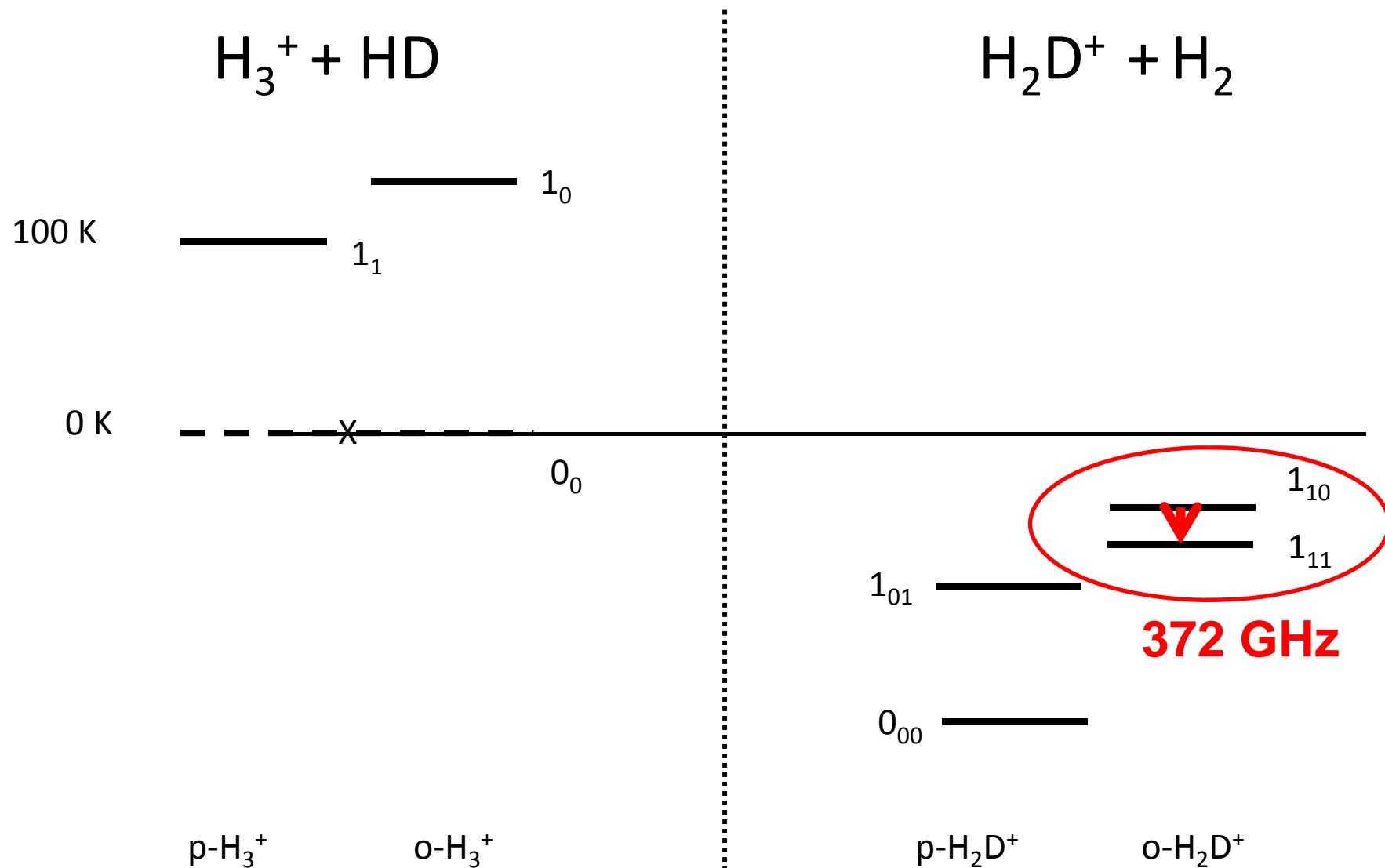
Stephan Schlemmer

Universität zu Köln

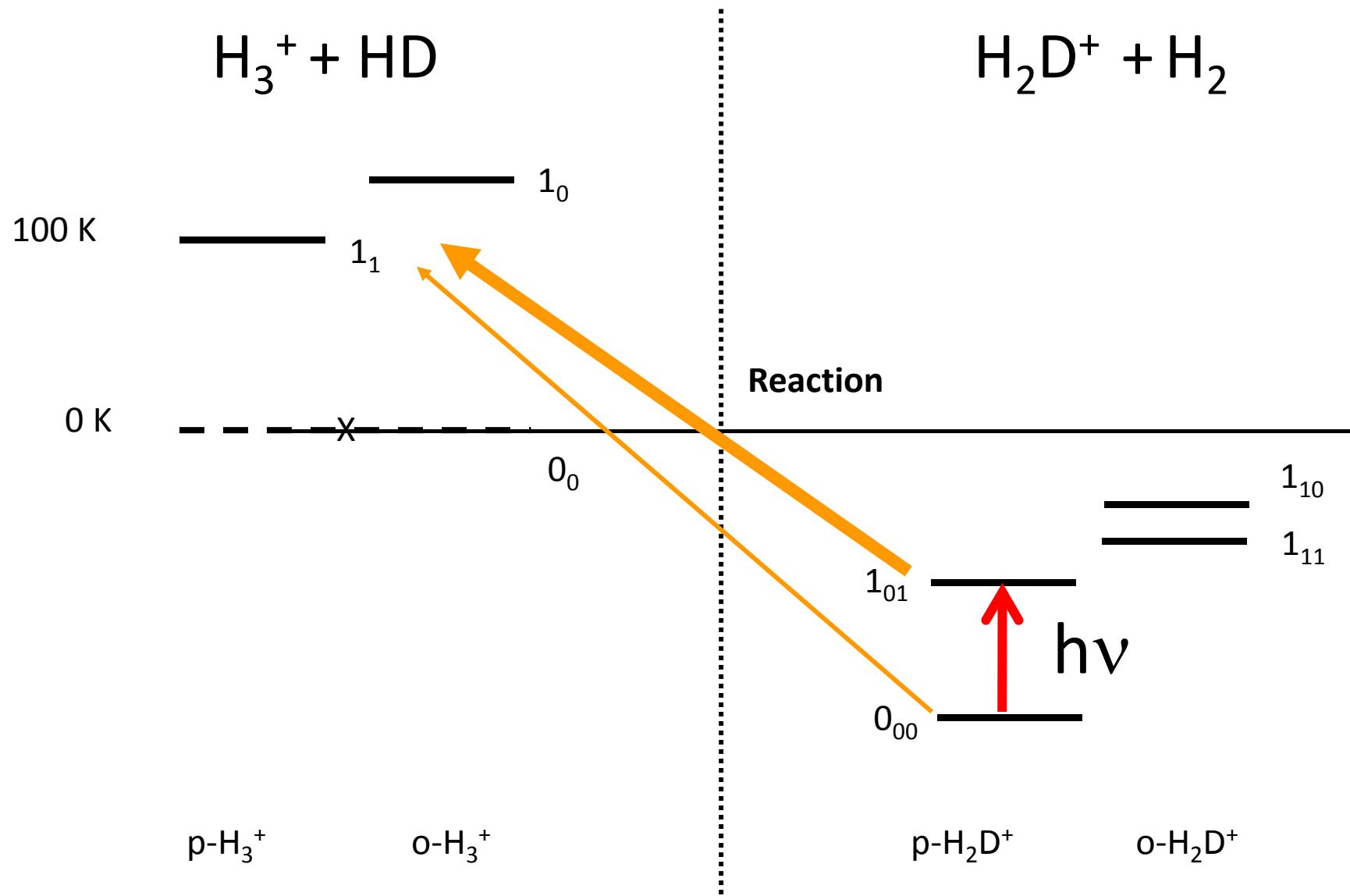


- $\text{H}_2$  Formation, OPR and Chemical Clocks
- $\text{H}_3^+ / \text{H}_2\text{D}^+$  Isotopic Fractionation,  $\text{H}_3^+/\text{H}_2\text{D}^+$ , OPR
- $\text{H}_2\text{D}^+ + \text{H}_2$  THz Spectroscopy in Lab and Space

# $\text{H}_2\text{D}^+$ Detection in Space

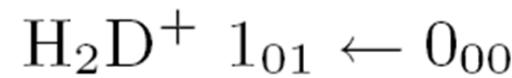
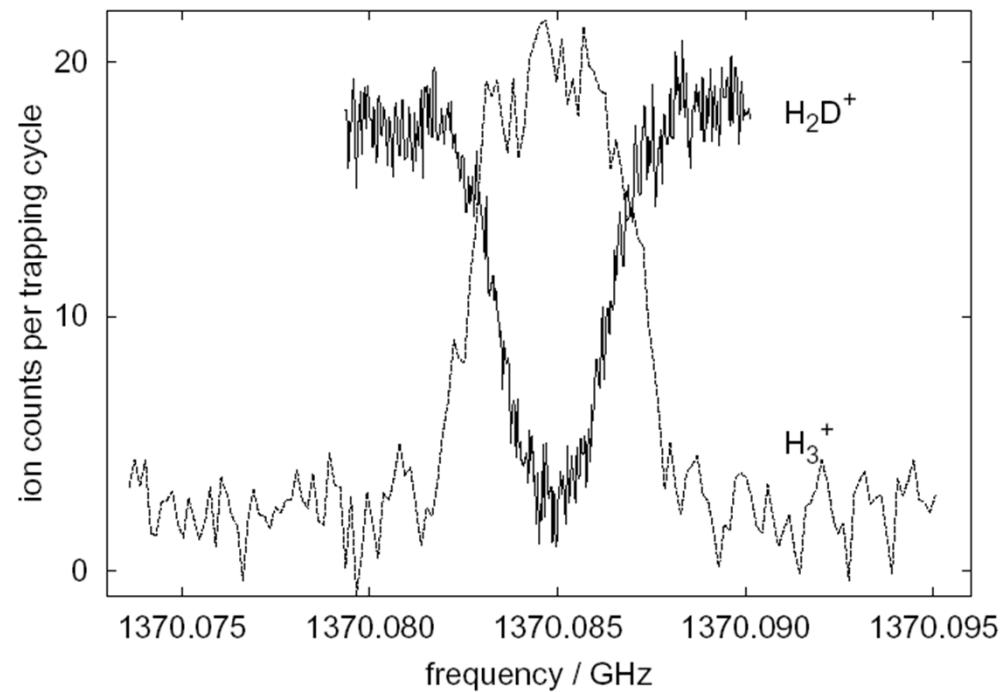


# *Light Induced Reactions probing H<sub>2</sub>D<sup>+</sup>*



## Results

$\text{H}_2\text{D}^+$



this work

1370084.880(20)

*ab initio*<sup>a</sup>

1369991.8

unpublished value<sup>b</sup>

1370146.0(3)

# SOFIA



O. Asvany et al.

$$1 \leftarrow U_{00}$$

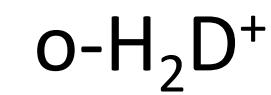
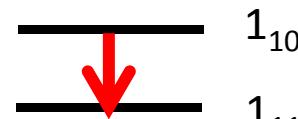
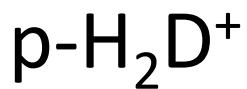
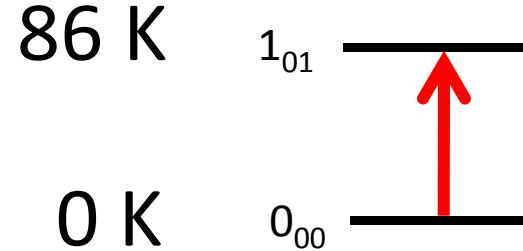
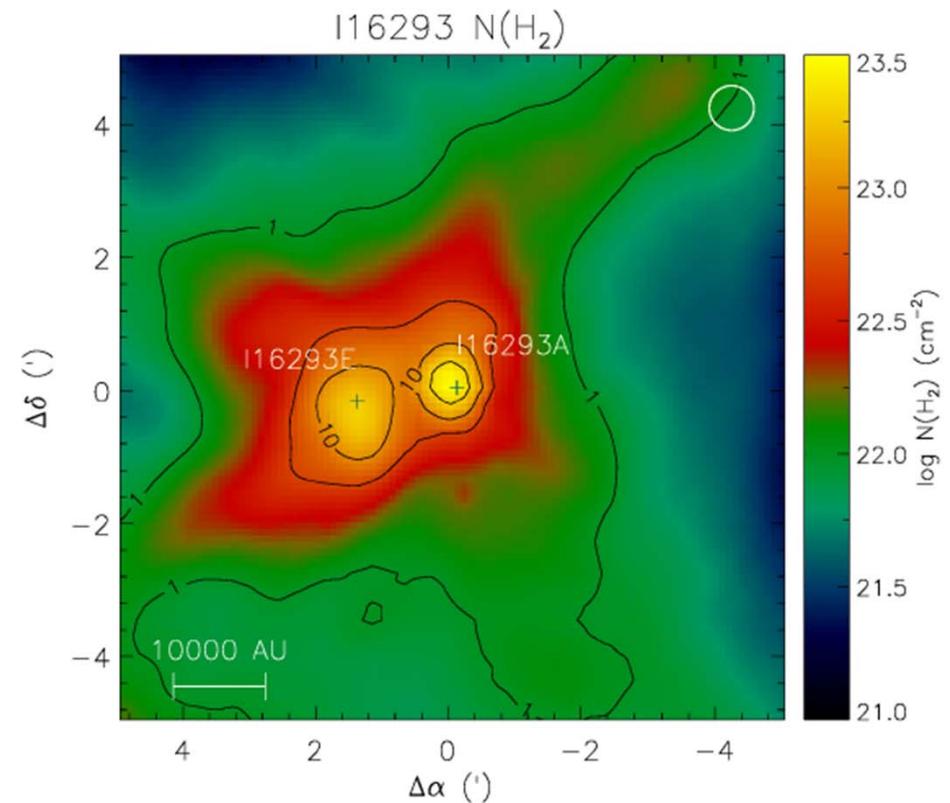
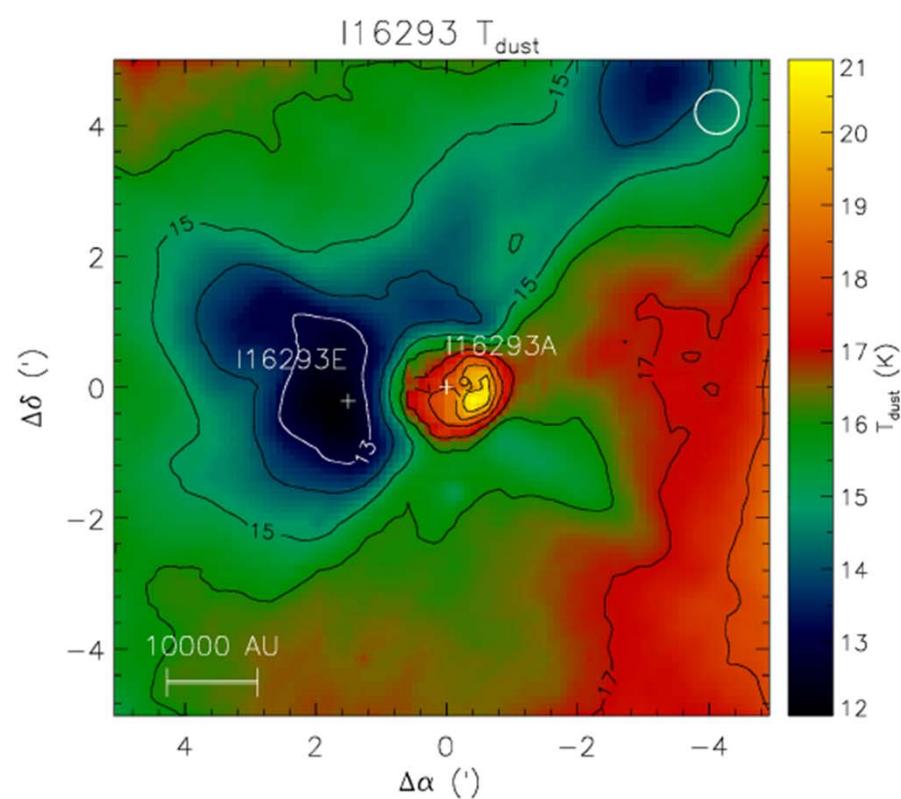
$$380(20)$$

3

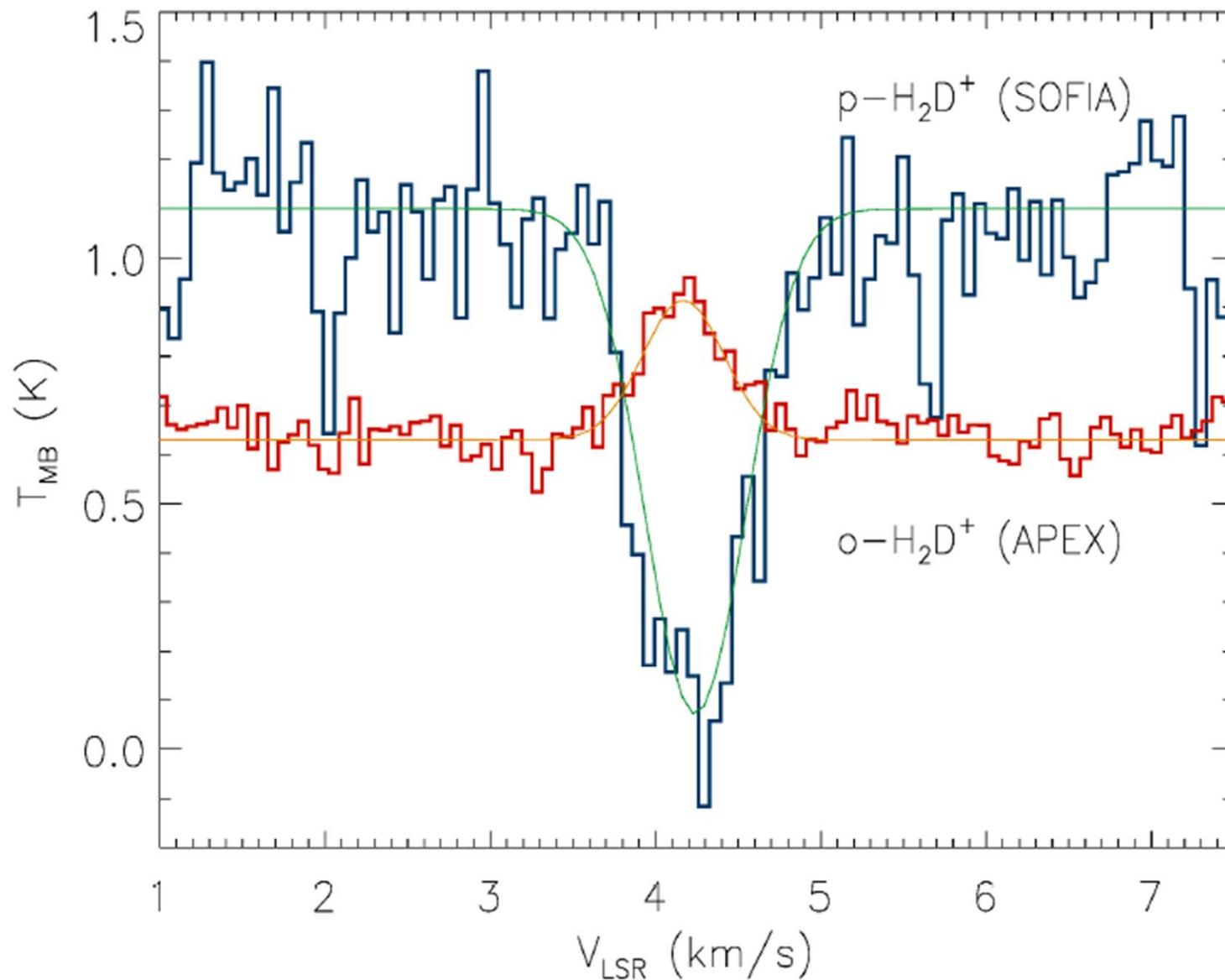
$$0(3)$$

$$\Delta = 62 \text{ MHz (!)}$$

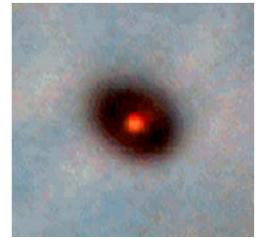
# Protostellar Cloud Core I16293A



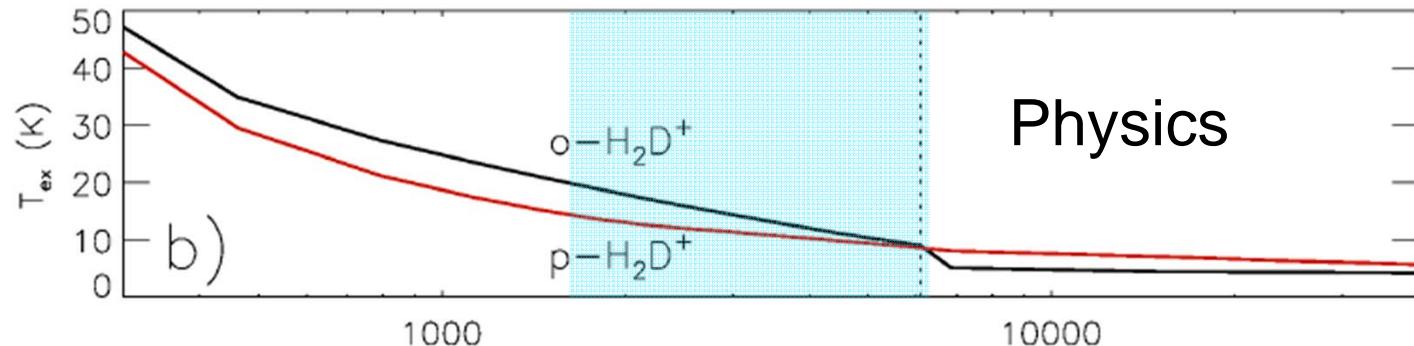
# Para- $\text{H}_2\text{D}^+$ found in Space



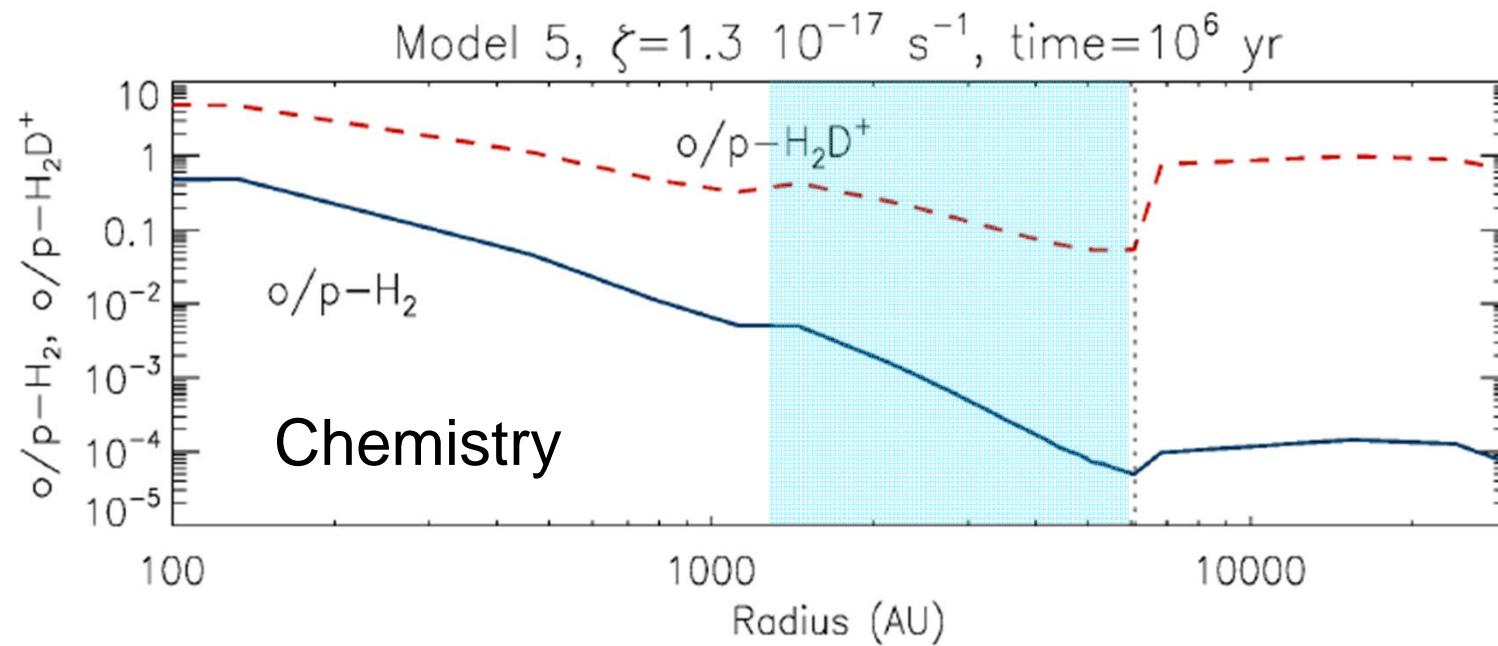
# Astrochemical Modelling



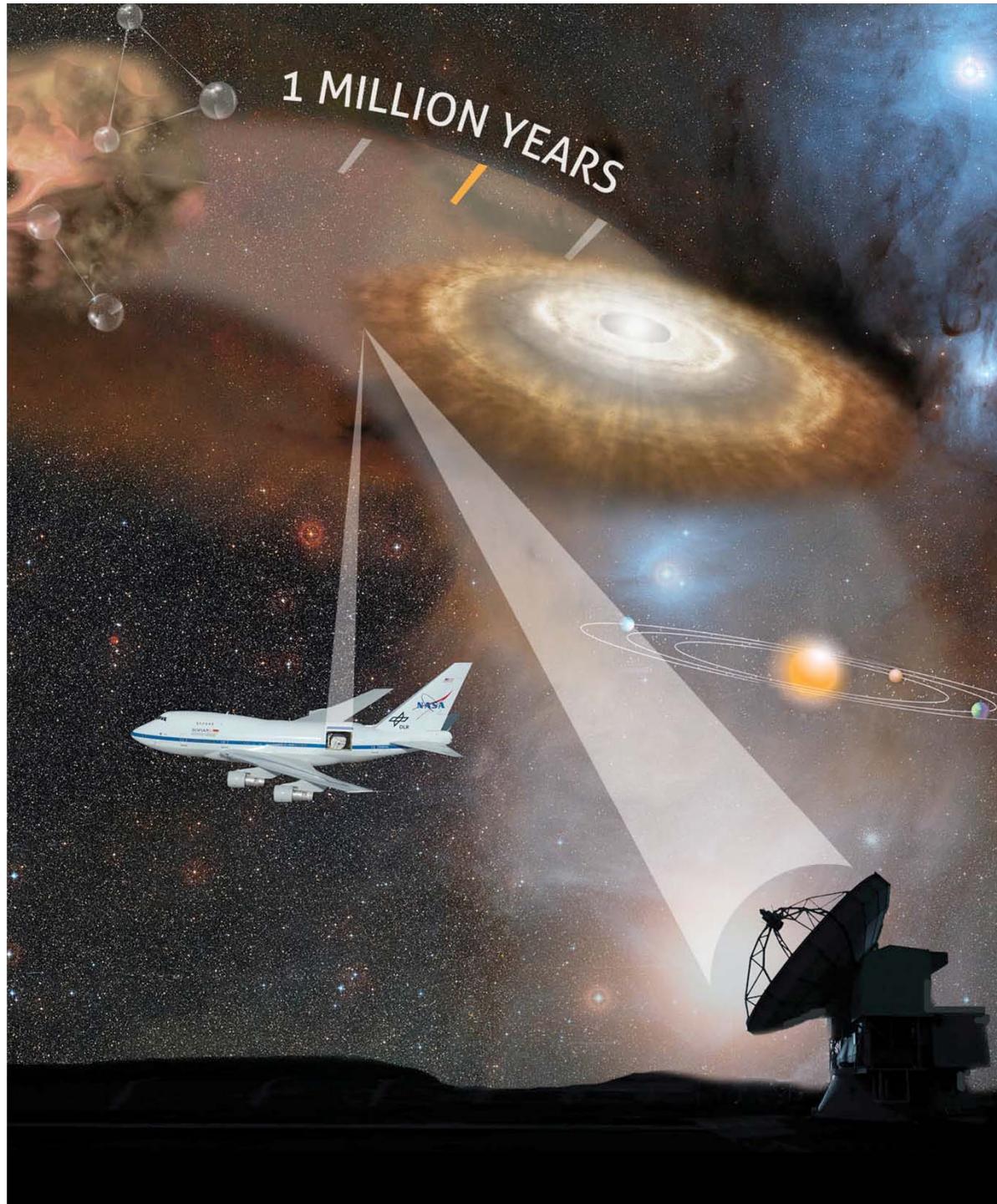
T



$\text{o/p}$

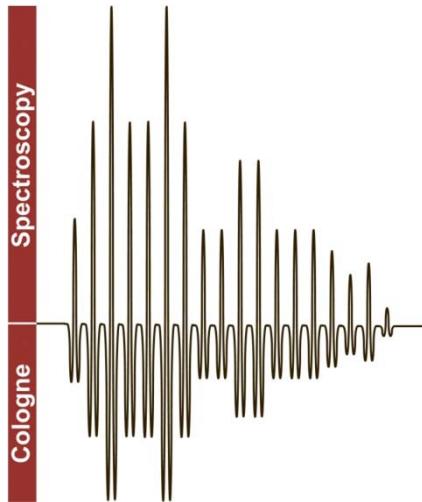


Collaboration: Jorma Harju, Olli Sipilä, Paola Caselli



$\text{H}_2\text{D}^+$  observations give an age of at least one million years for a cloud core forming Sun-like stars

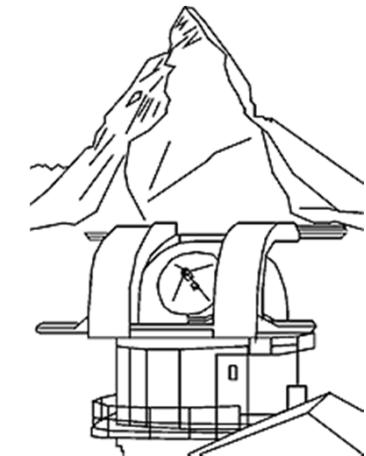
S. Brünken et al.  
Nature  
[doi:10.1038/nature13924](https://doi.org/10.1038/nature13924)



# *Cold Chemistry in Space and Laboratory*

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- H<sub>2</sub> Formation, OPR and Chemical Clocks
- H<sub>3</sub><sup>+</sup> / H<sub>2</sub>D<sup>+</sup> Isotopic Fractionation, H<sub>3</sub><sup>+</sup>/H<sub>2</sub>D<sup>+</sup>, OPR
- H<sub>2</sub>D<sup>+</sup> + H<sub>2</sub> THz Spectroscopy in Lab and Space