

# A SOFIA/GREAT view on the Cepheus E outflow from an intermediate-mass protostar

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Rolf Güsten, Silvia Leurini

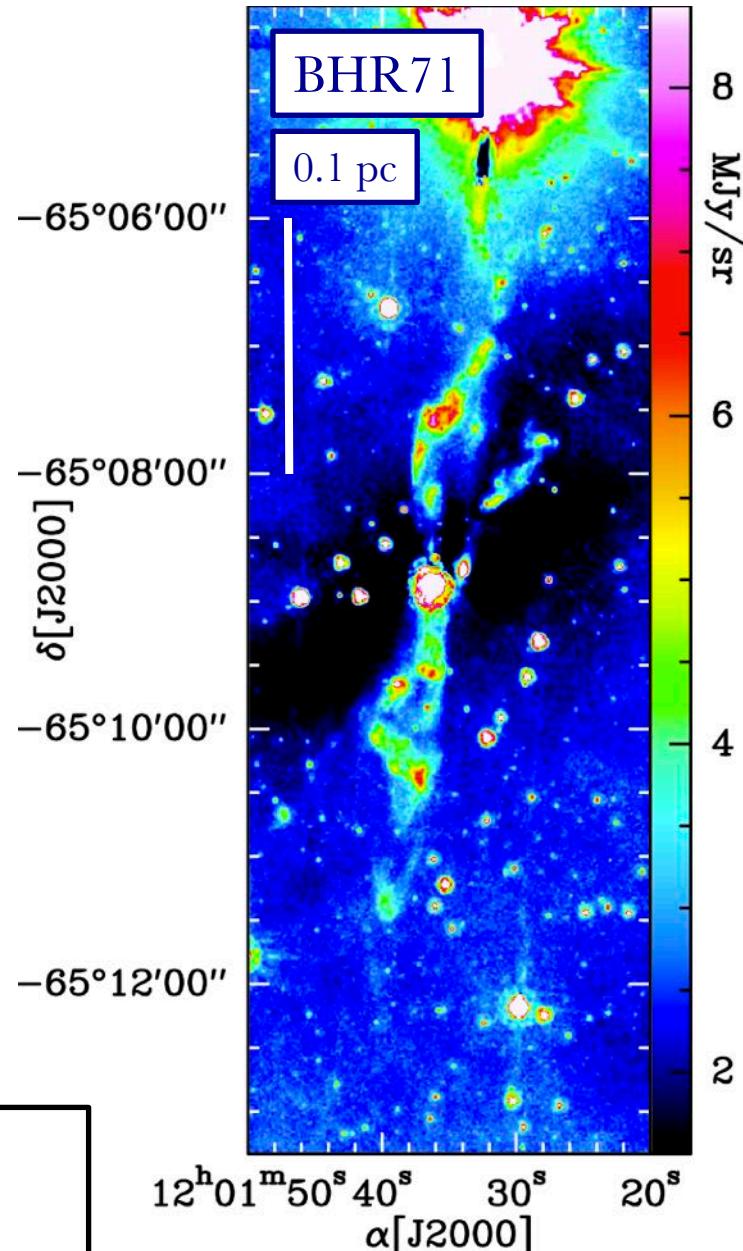
LERMA, Paris, France  
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INAF Firenze, Italy  
MPIfR, Bonn, Germany

# Outline

- Motivations
- Previous studies of Cepheus E
- The SOFIA/GREAT observations: CO
- The SOFIA/GREAT observations: [OI]
- Shock modelling in Cepheus E
- Perspectives

# Motivations

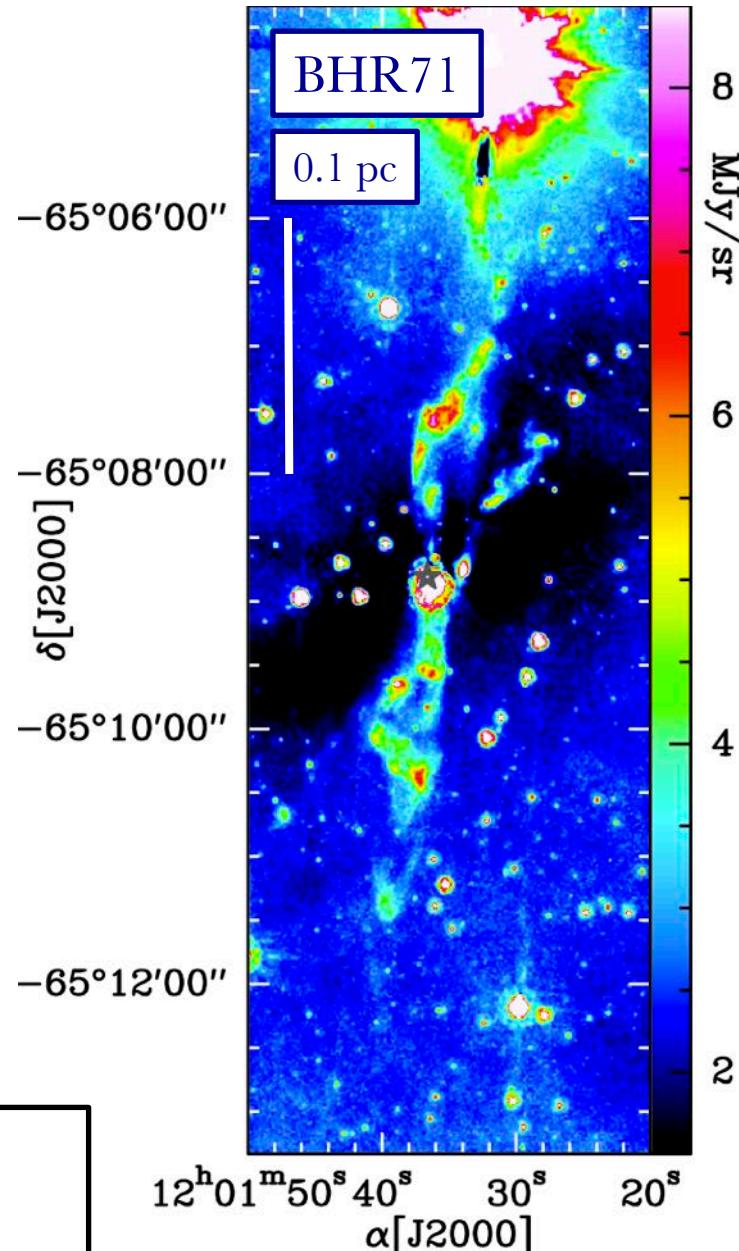
# Motivations



archive image  
*Spitzer*/IRAC 8  $\mu\text{m}$

adapted from  
Gusdorf et al. 2015

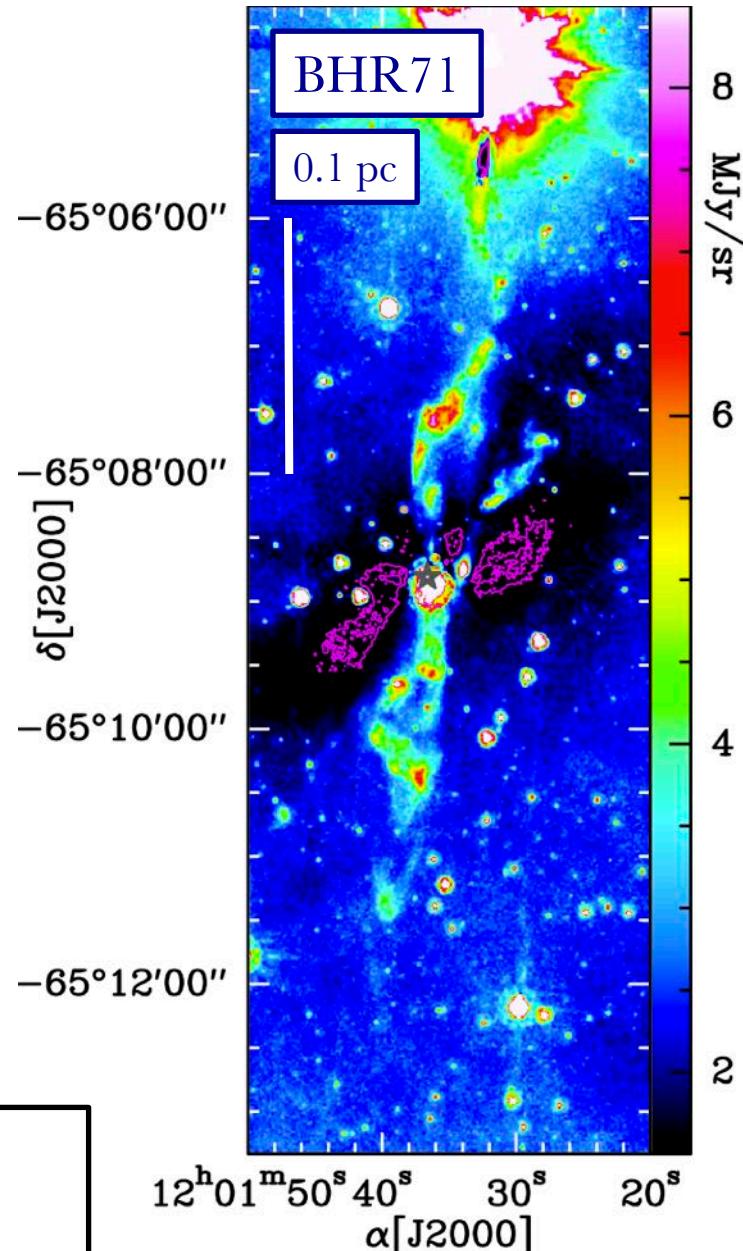
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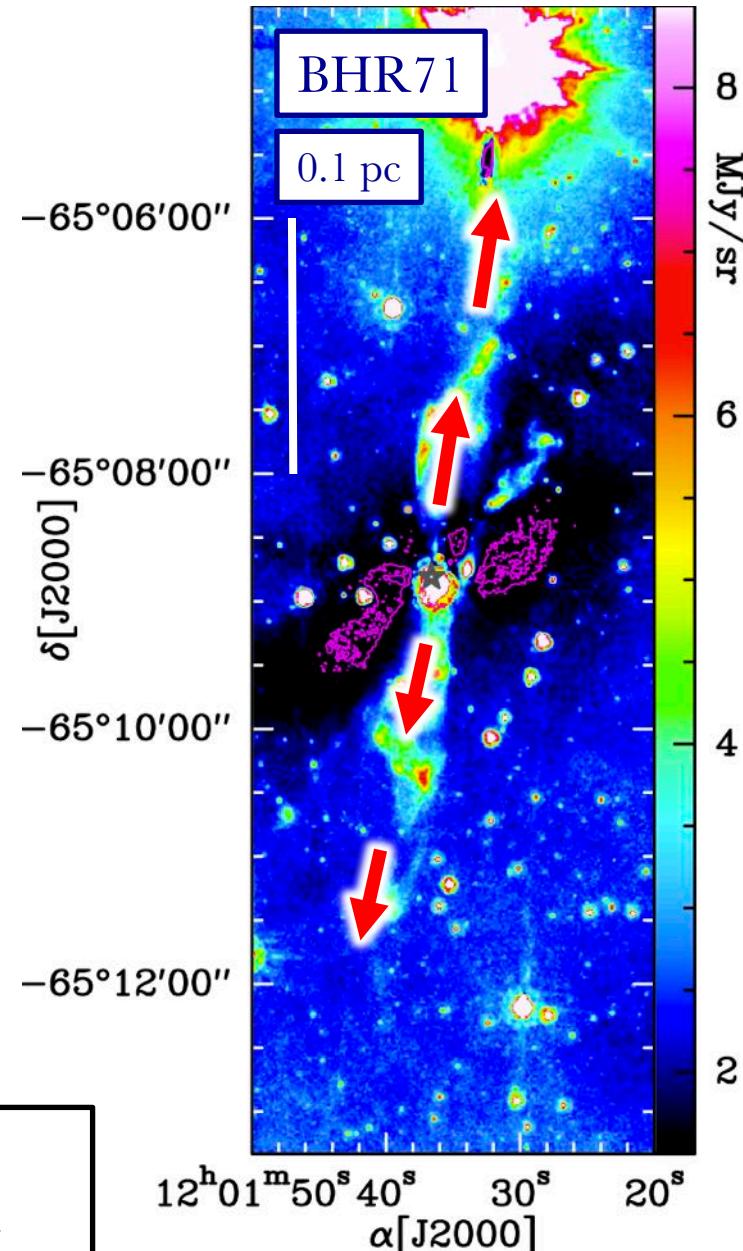
# Motivations



archive image  
*Spitzer*/IRAC 8  $\mu$ m

adapted from  
Gusdorf et al. 2015

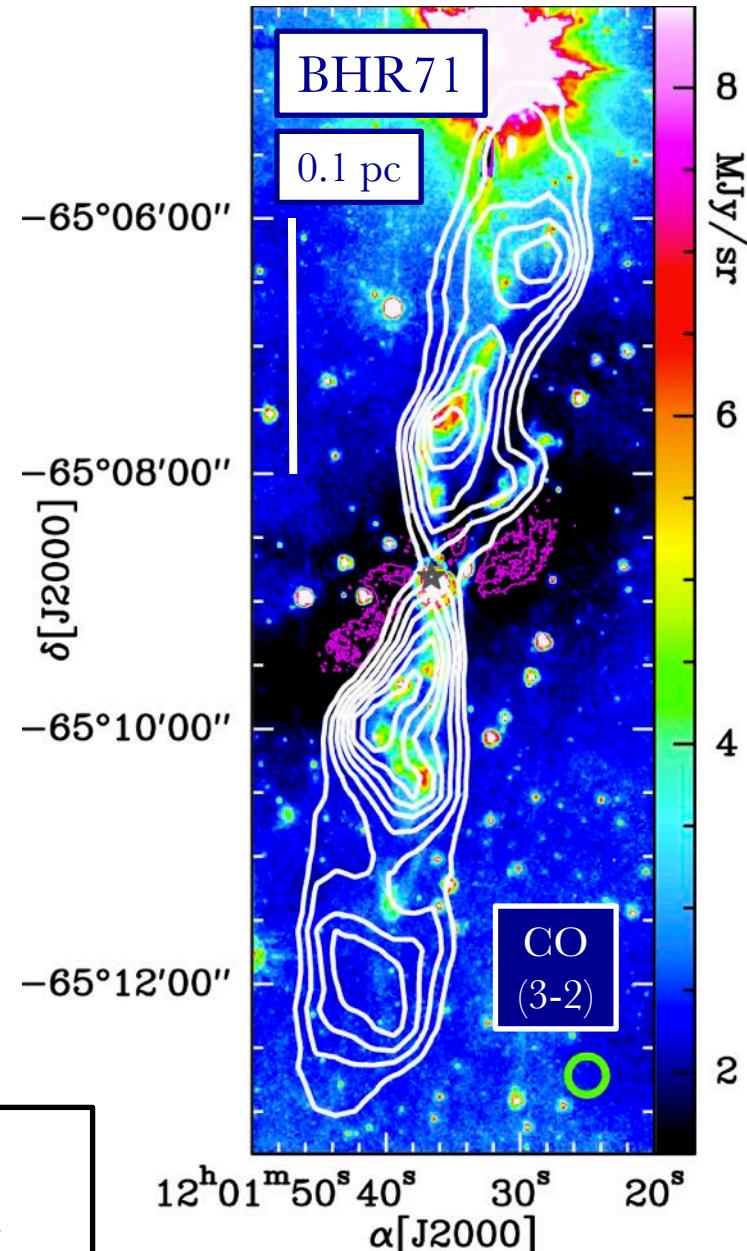
# Motivations



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*Spitzer*/IRAC 8  $\mu\text{m}$

adapted from  
Gusdorf et al. 2015

# Motivations



archive image  
*Spitzer*/IRAC 8 μm

adapted from  
Gusdorf et al. 2015

# Motivations

Physical mechanisms,  
nature of shocks:

- parametric molecular

Jimenez-Serra et al. 2008

- molecular

Neufeld & Kaufman 1996

Flower & Pineau des Forêts 2015

- radiative precursor

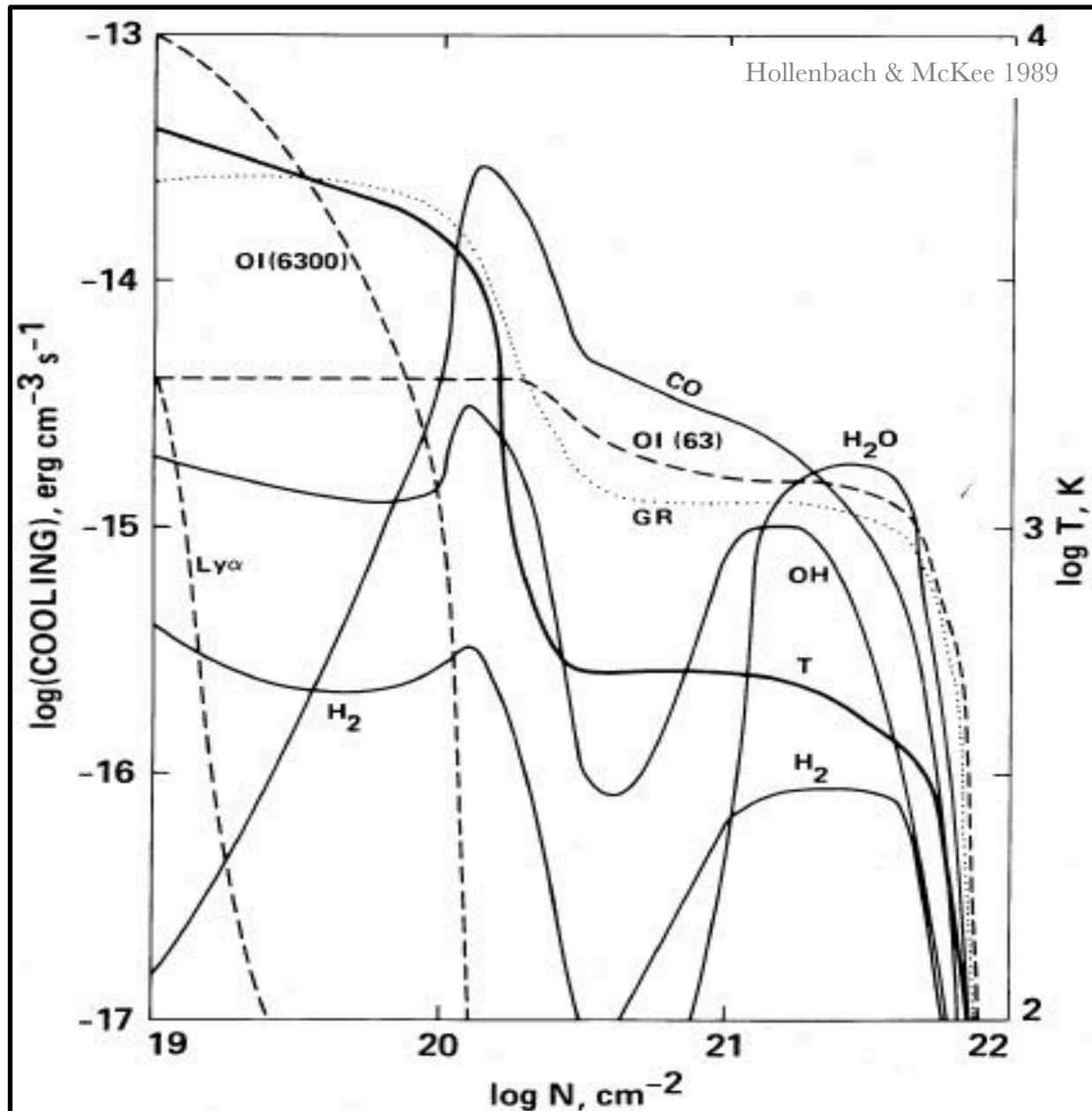
Hollenbach & McKee 1989

- irradiated

(=FUV-illuminated)

Lesaffre et al. 2013,

Melnick & Kaufman 2015



# Motivations

Physical mechanisms,  
nature of shocks

Astrochemistry:

- SiO

Caselli et al. 1997,  
Schilke et al. 1997,  
Gusdorf et al. 2008a,b, 2011,  
Guillet et al. 2009,  
Jimenez-Serra et al. 2009,  
Nguyen Luong et al. 2011,  
Anderl et al. 2013,  
Leurini et al. 2014,...

- H<sub>2</sub>O/OH/OI

WISH/WADI *Herschel* KP,  
Recent SOFIA works

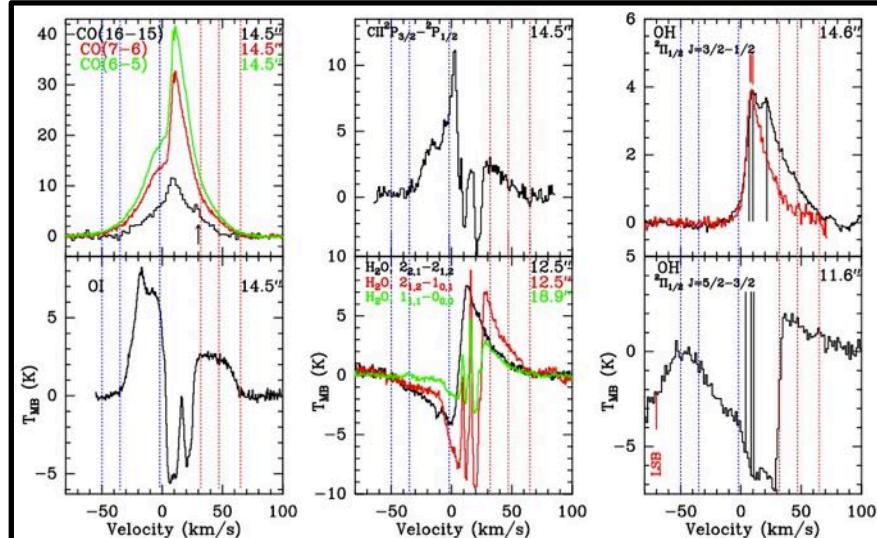
- ionized species

Podio et al. 2014 & ref. therein

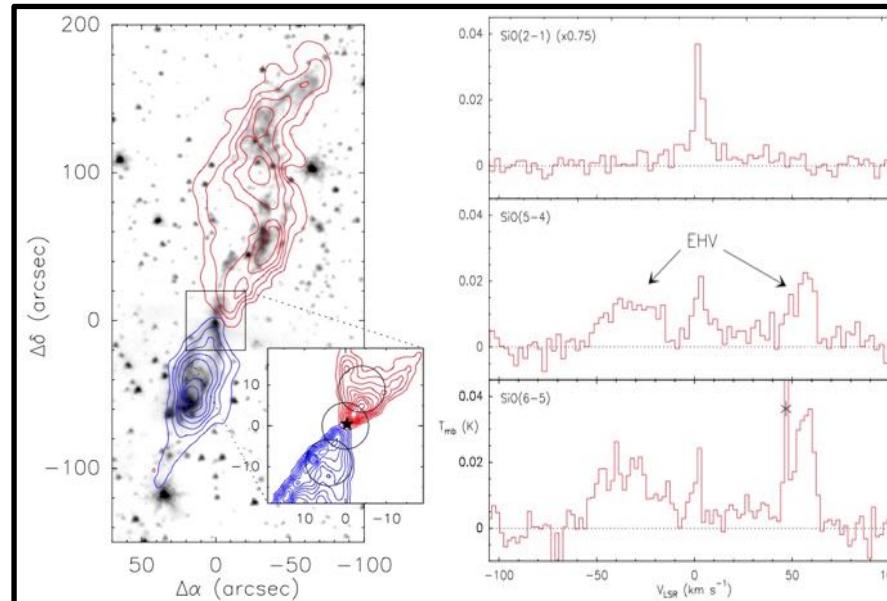
- COMs

Codella et al. 2015 & ref. therein

Also see L. Podio's talk



Leurini et al. 2015  
G5.89-0.39



Tafalla et al. 2015  
L1157

# Motivations

Physical mechanisms,  
nature of shocks

Astrochemistry

Energetic impacts and  
contribution to energetic  
balance of galaxies:  
- Galactic studies

Cabrit & Berthout 1990,

Bontemps et al. 1996,

Beuther et al. 2002,

Zhang et al. 2005,

Lopez-Sepulcre et al. 2009,

Visser et al. 2014 & ref. therein

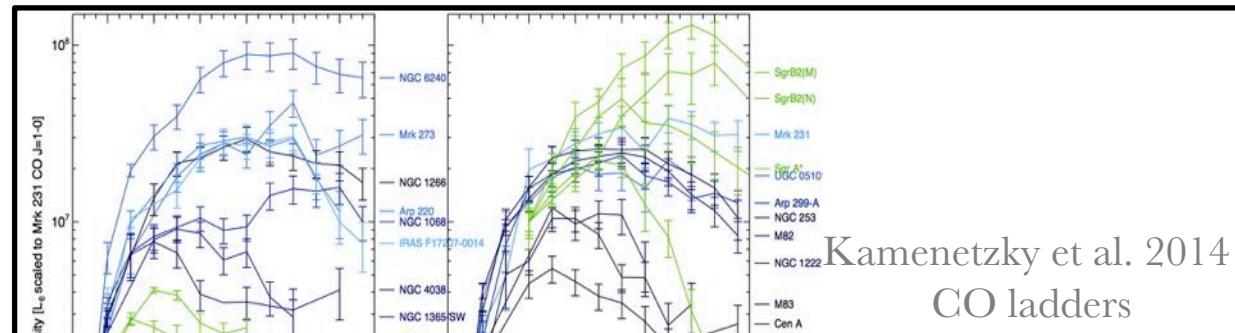
- extragalactic studies

Hailey-Dunsheath et al. 2012,

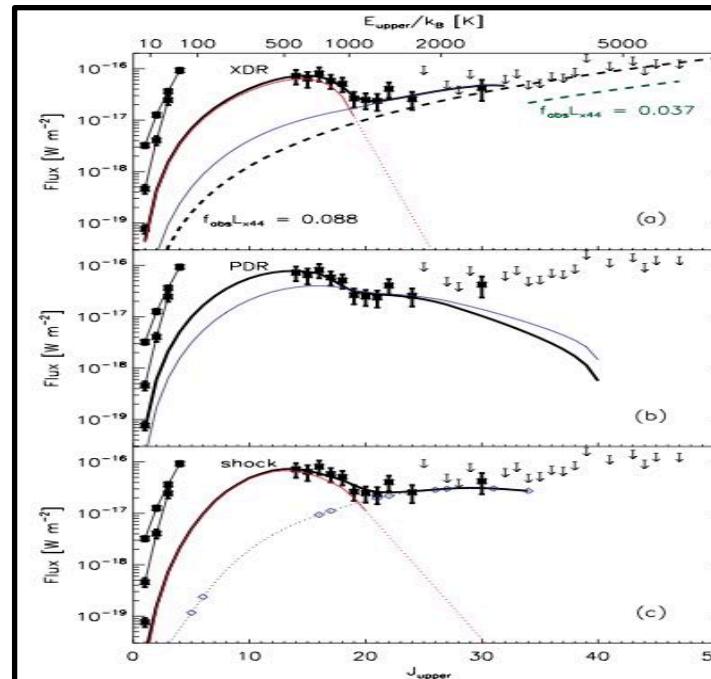
Meijerink et al. 2013,

Kamenetzky et al. 2014

& ref. therein



Kamenetzky et al. 2014  
CO ladders



Hailey-Dunsheath et al. 2012  
CO ladders in NGC1068

# Motivations

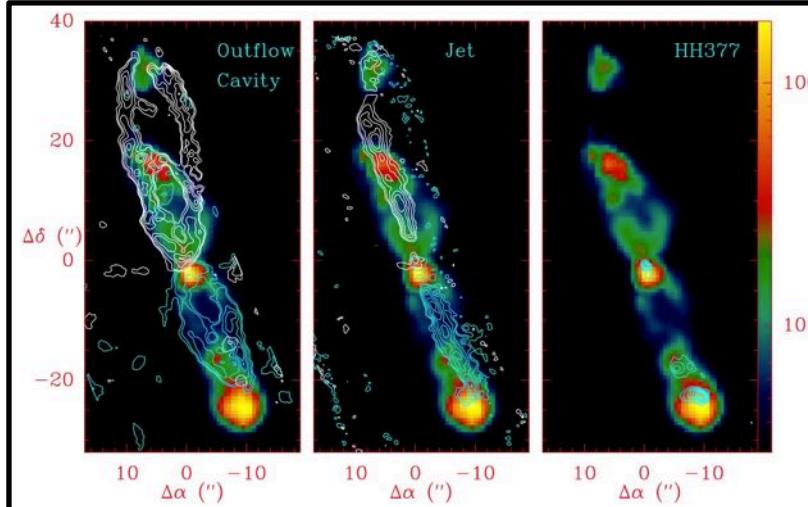
Physical mechanisms,  
nature of shocks

Astrochemistry

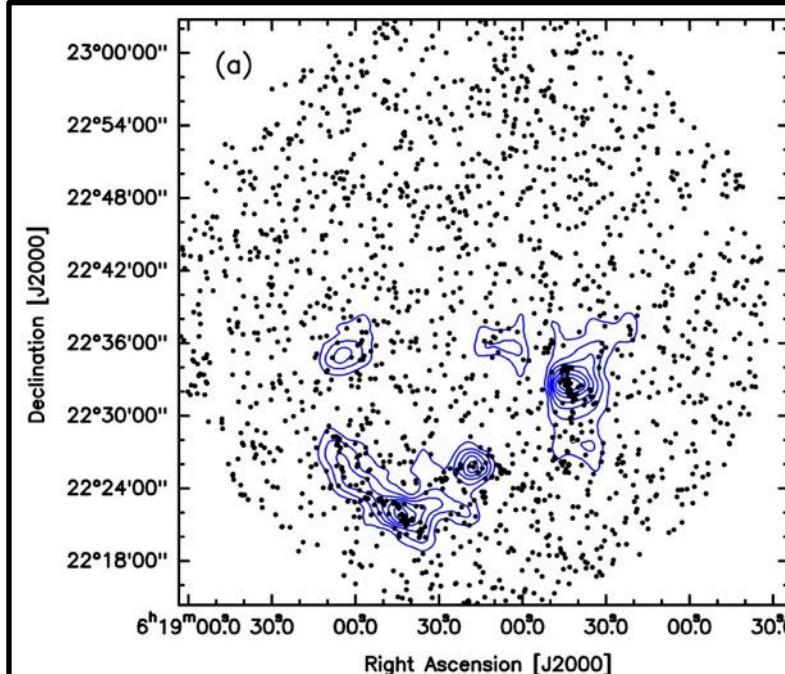
Energetic impacts and  
contribution to energetic  
balance of galaxies

Star formation scenarios  
through the mass-ladder:  
- inducing  
low-mass to high-mass

Codella et al. 2012,  
CALYPSO program,  
Lefloch et al 2015,  
Hunter et al. 2008  
- induced  
Xu et al. 2011



Lefloch et al. 2015  
CO in Cep E



Xu et al. 2011  
SF induced by SNR  
shock in IC443 ?

# Motivations

Physical mechanisms,  
nature of shocks

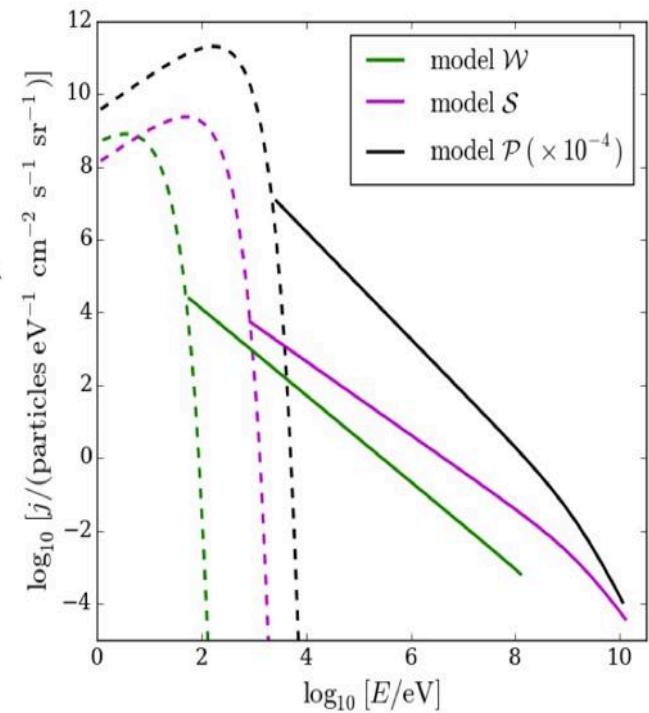
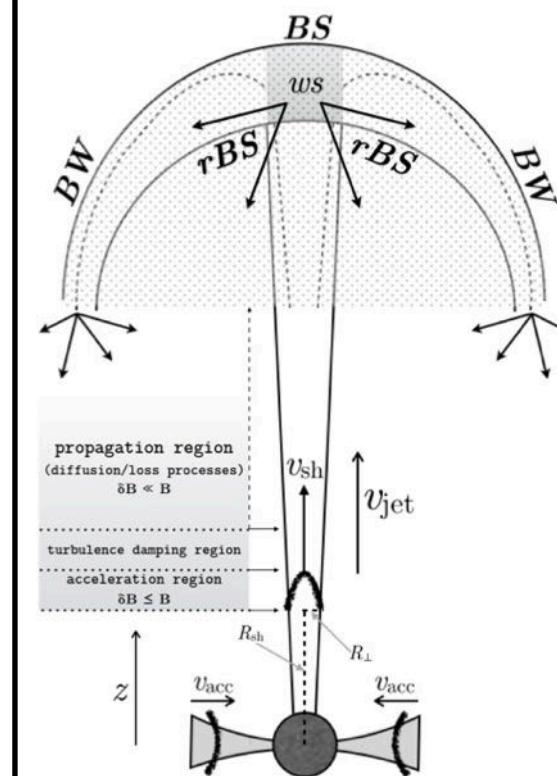
Astrochemistry

Energetic impacts and  
contribution to energetic  
balance of galaxies

Star formation scenarios  
through the mass-ladder  
(induced or inducing shocks)

CR-related questions:  
acceleration, composition,  
propagation:

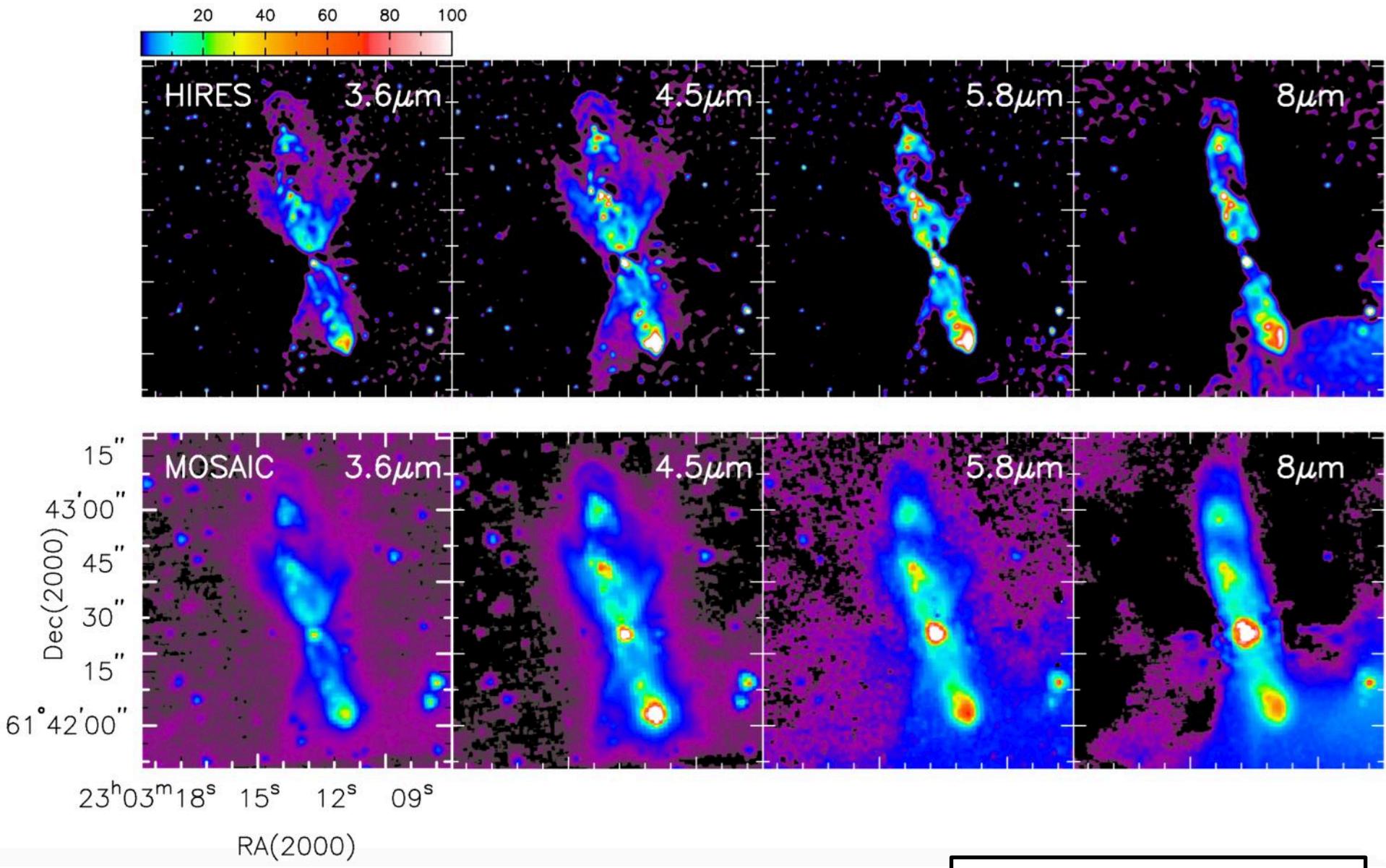
Araudo et al. 2007,  
Bosch-Ramon et al. 2010,  
Munar-Adrover et al. 2011,  
Padovani et al. 2015, 2016



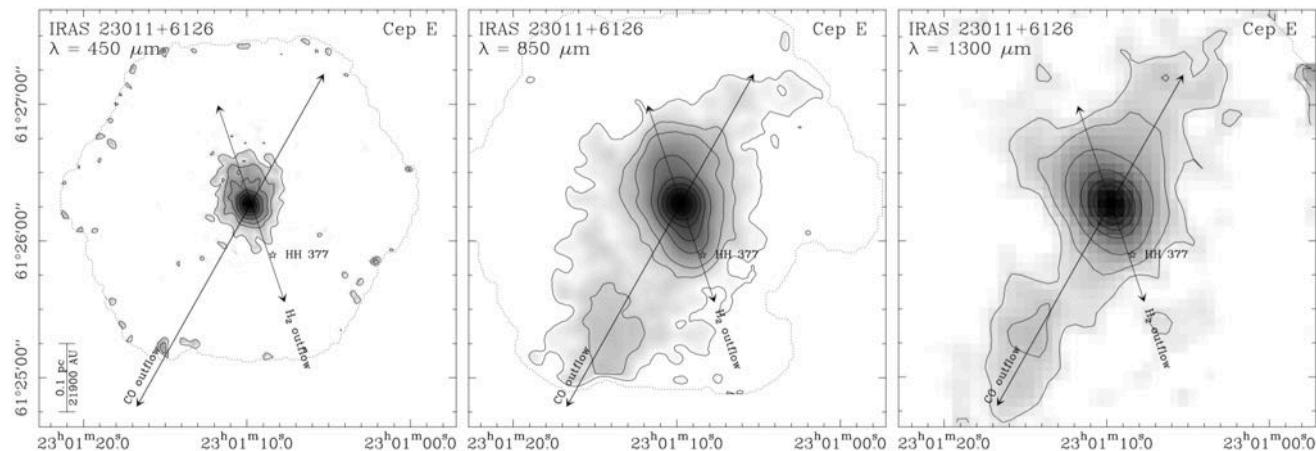
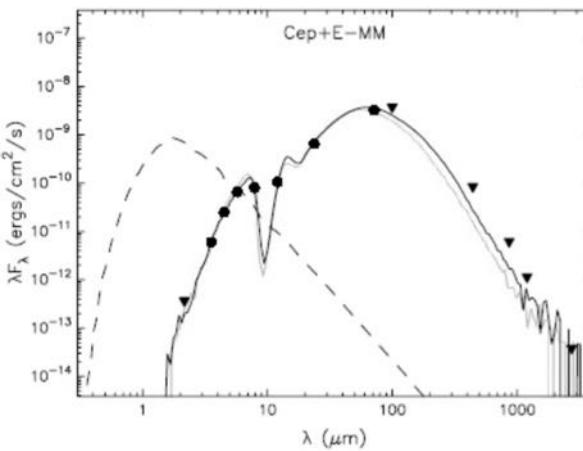
Padovani et al. 2015, 2016

# Previous studies of Cepheus E

# Previous studies of Cep E



# Previous studies of Cep E



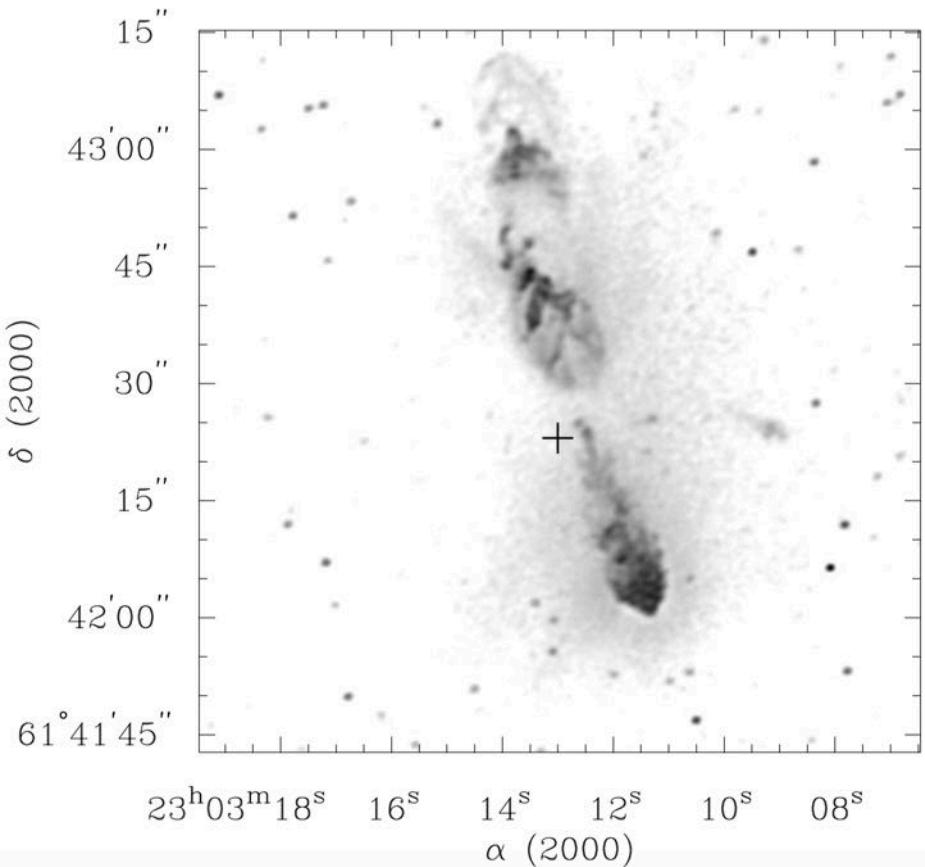
Velusamy et al. 2011

Chini et al. 2001

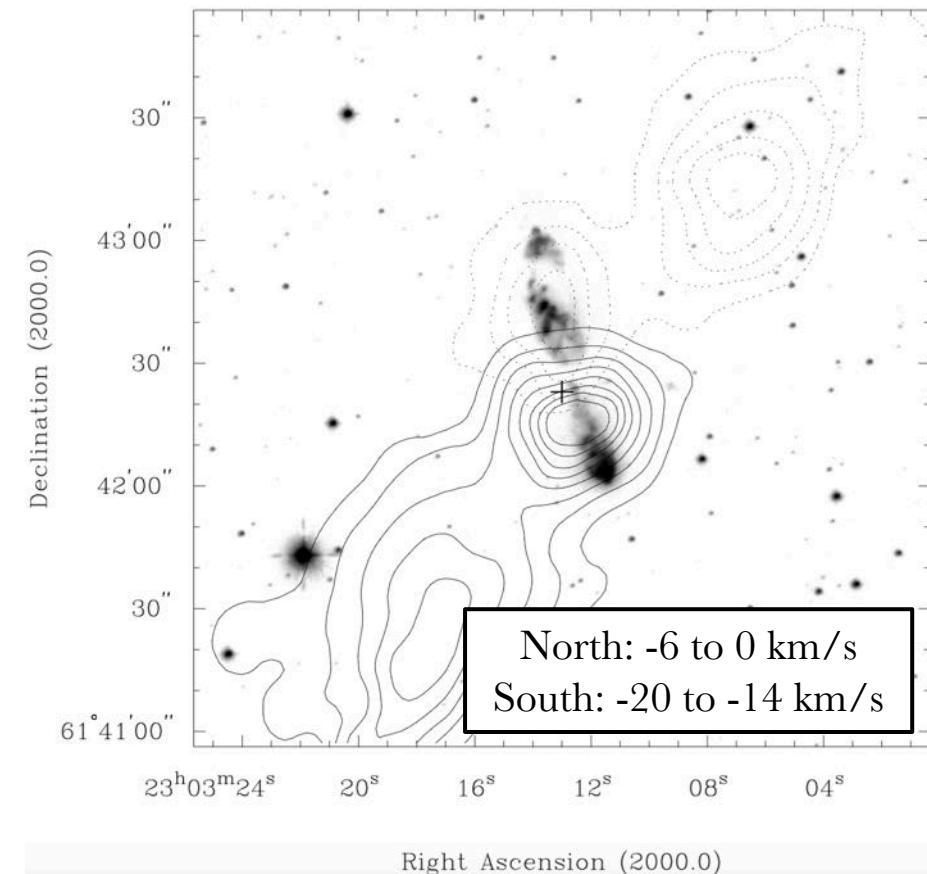
- Lefloch et al. 1996, Ladd & Hodapp 1997, Chini et al. 2001, Froebrich et al. 2003, Noriega-Crespo et al. 2004, Velusamy et al. 2011:

- Class 0, few 10 000 yrs
- $L \sim 100 L_\odot$ ;  $M \sim 3 M_\odot$ ;  $L_{\text{smm}} \sim 0.017 L_{\text{bol}}$
- $d = 730 \text{ pc}$ ,  $v_{\text{lsr}} = -11.2 \text{ km/s}$
- Compact cloud core: 0.18 pc size,  $8 M_\odot$
- $M_{\text{disk}} \sim 7 \cdot 10^{-3} M_\odot$ ;  $M_{\text{disk}}' \sim 3 \cdot 10^{-8} M_\odot/\text{yr}$
- $M_{\text{env}} \sim 7.4 M_\odot$ ;  $M_{\text{env}}' \sim 10^{-4} M_\odot/\text{yr}$

# Previous studies of Cep E

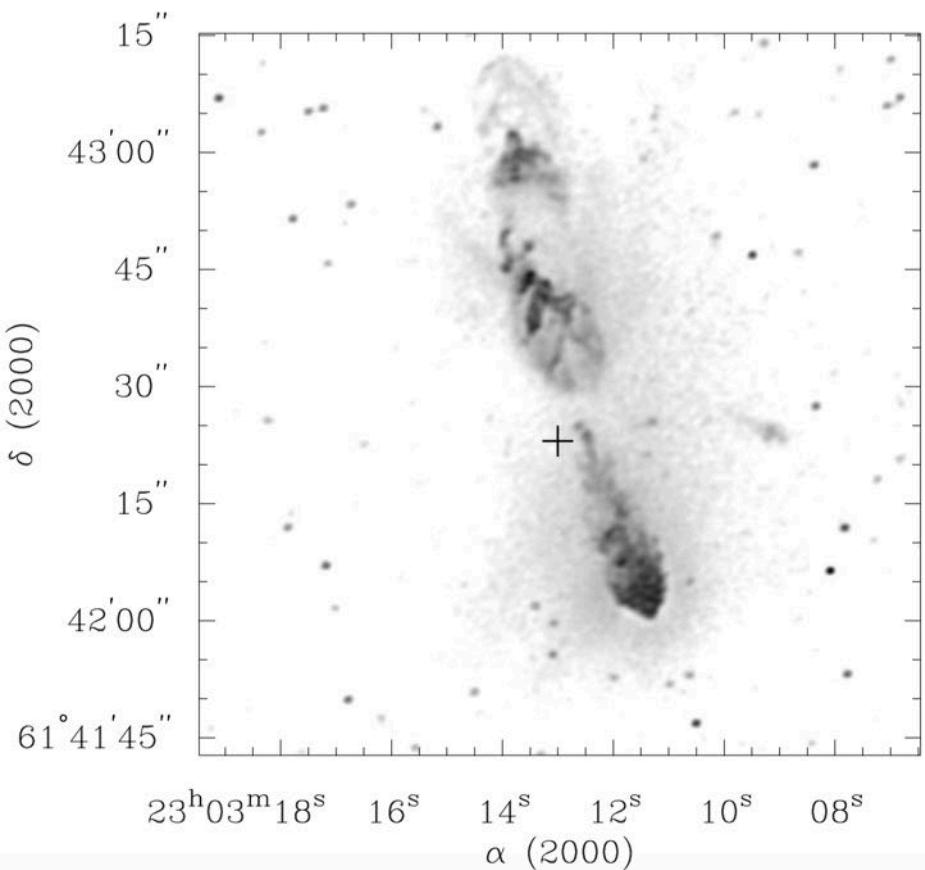


Cep E in  $\text{H}_2$  1-0 S(1) at  $2.12 \mu\text{m}$ ...

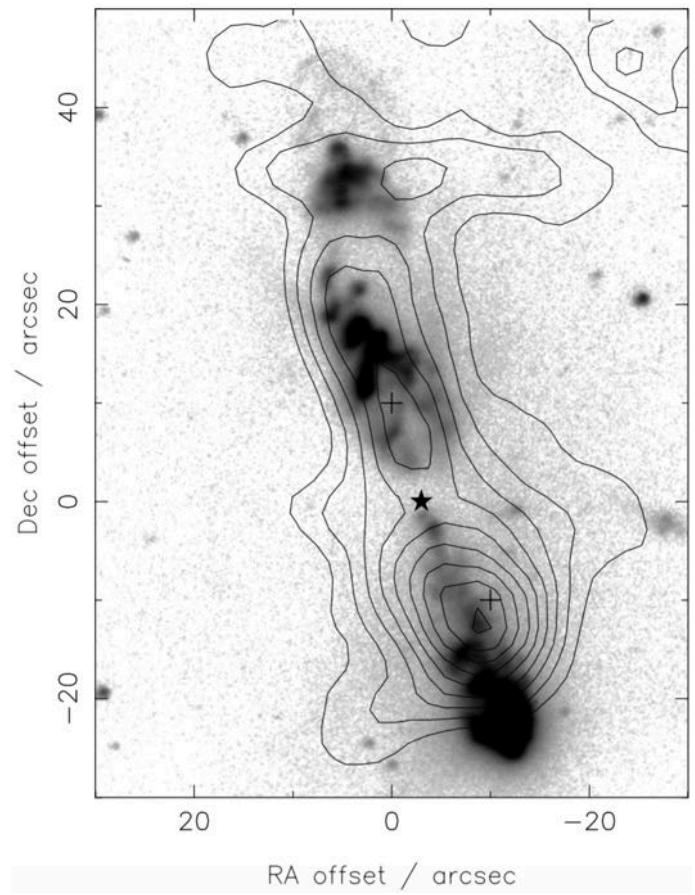


...and CO (2-1)

# Previous studies of Cep E



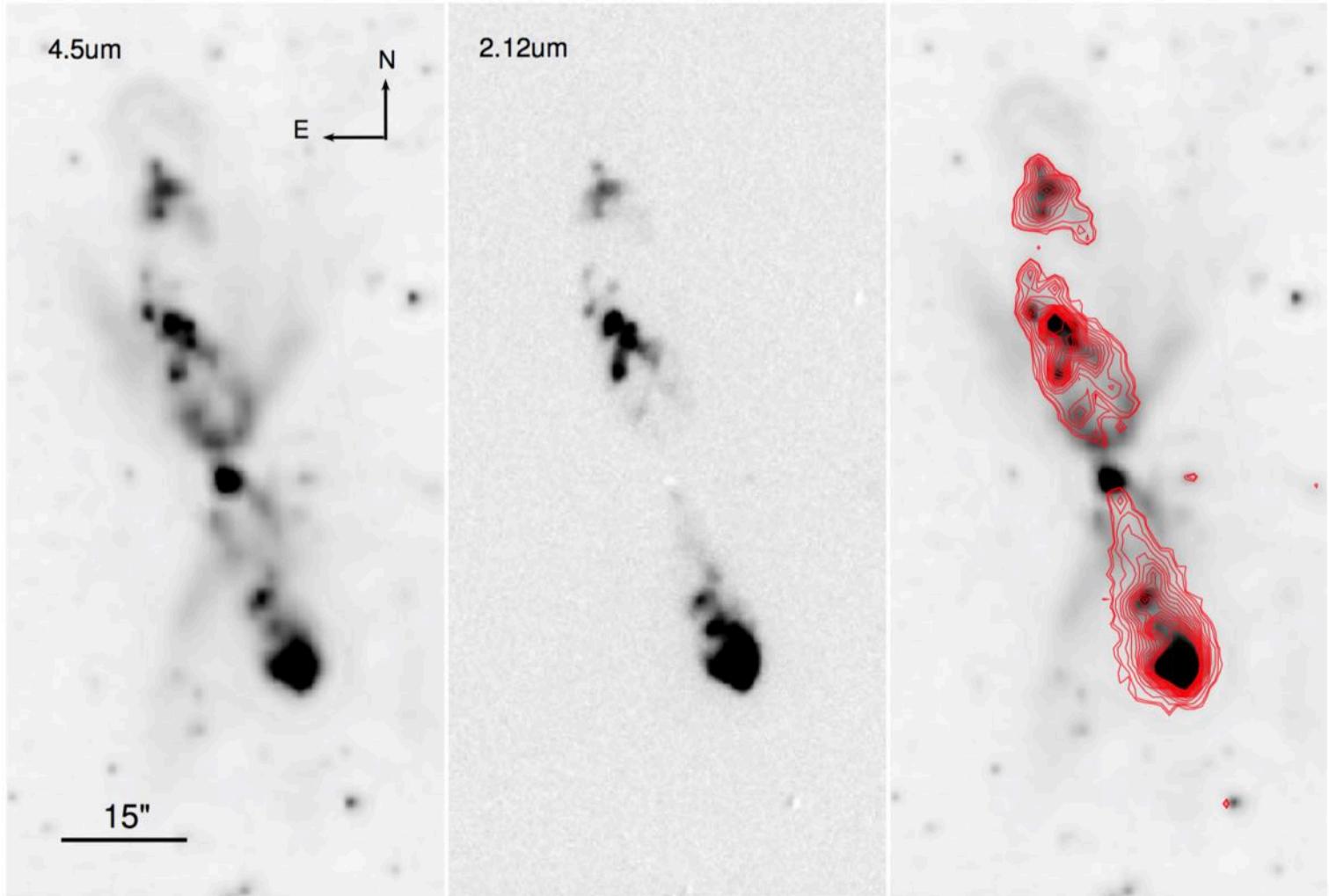
Cep E in  $\text{H}_2$  1-0 S(1) at 2.12  $\mu\text{m}$ ...



...and CO (4-3)

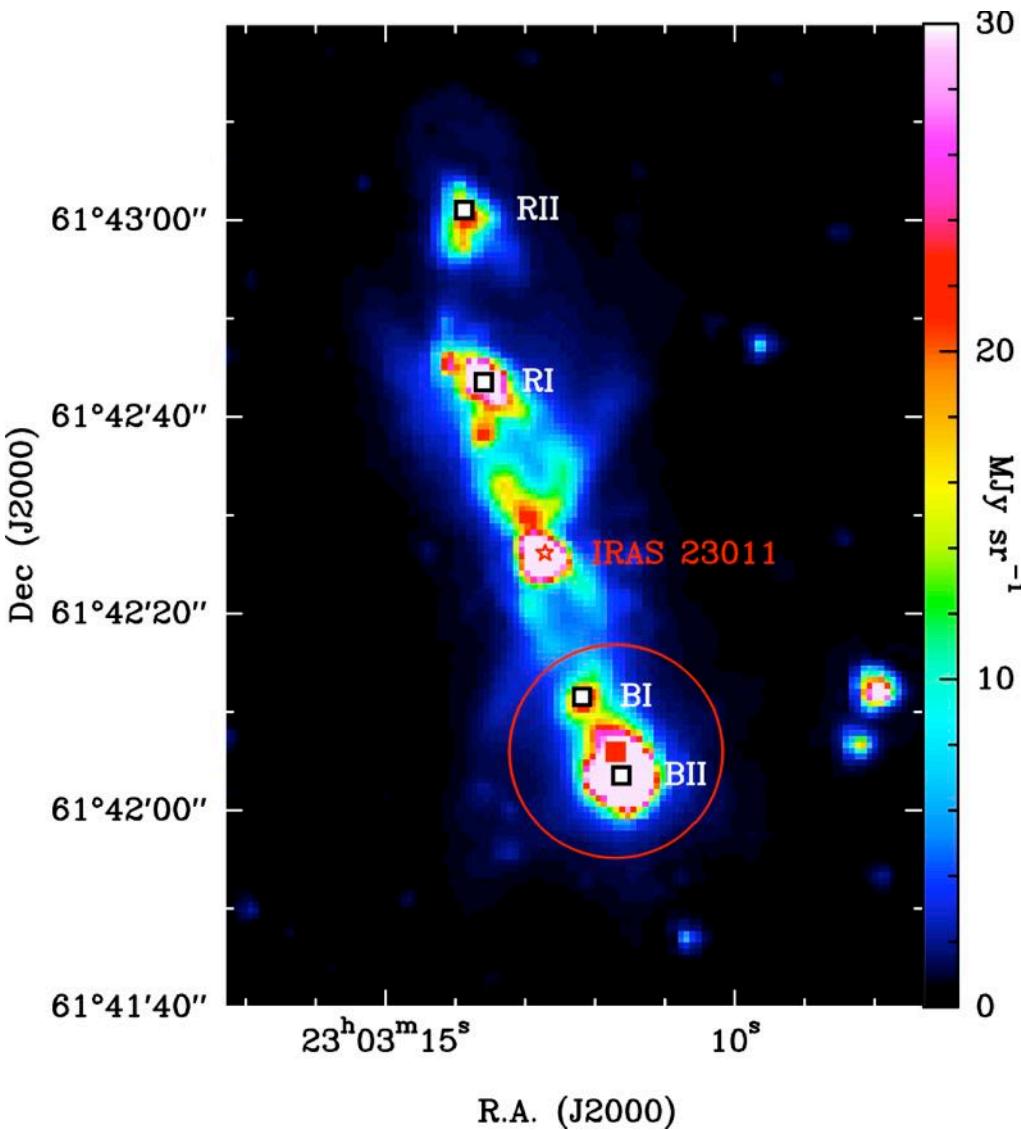
# Previous studies of Cep E

- Numerous NIR, MIR and sub-mm observations (ISO, *Spitzer*, IRAM-30m): Eisloffel et al. 1996, Noriega-Crespo et al. 1998, 2014, Moro-Martin et al. 2001, Smith et al. 2003, Lefloch et al. 2011, ...



# The SOFIA/GREAT observations: CO

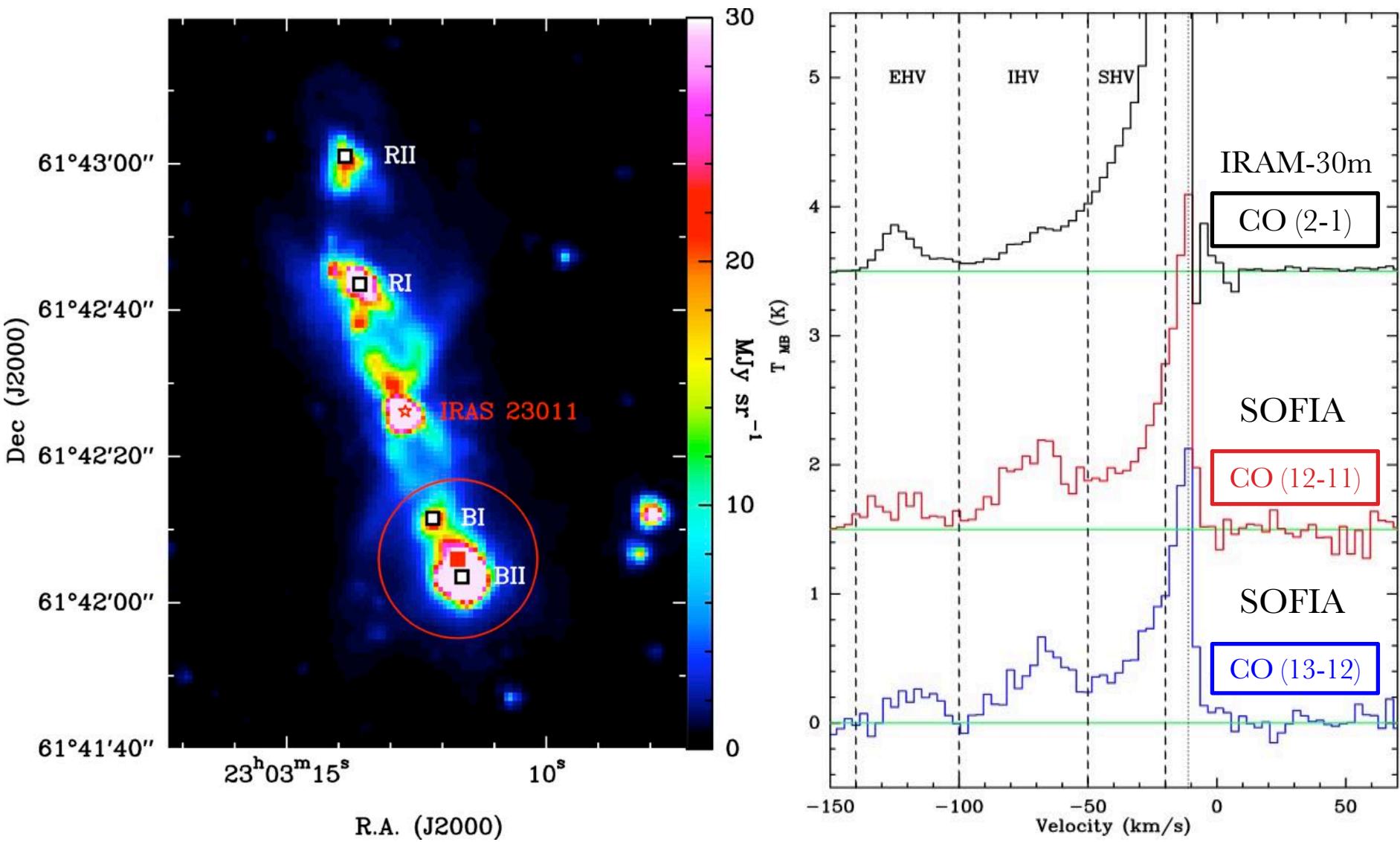
# The SOFIA/GREAT observations: CO



The Cep E outflow at 4.5  $\mu\text{m}$  by *Spitzer*/IRAC

Gomez-Ruiz et al. 2012

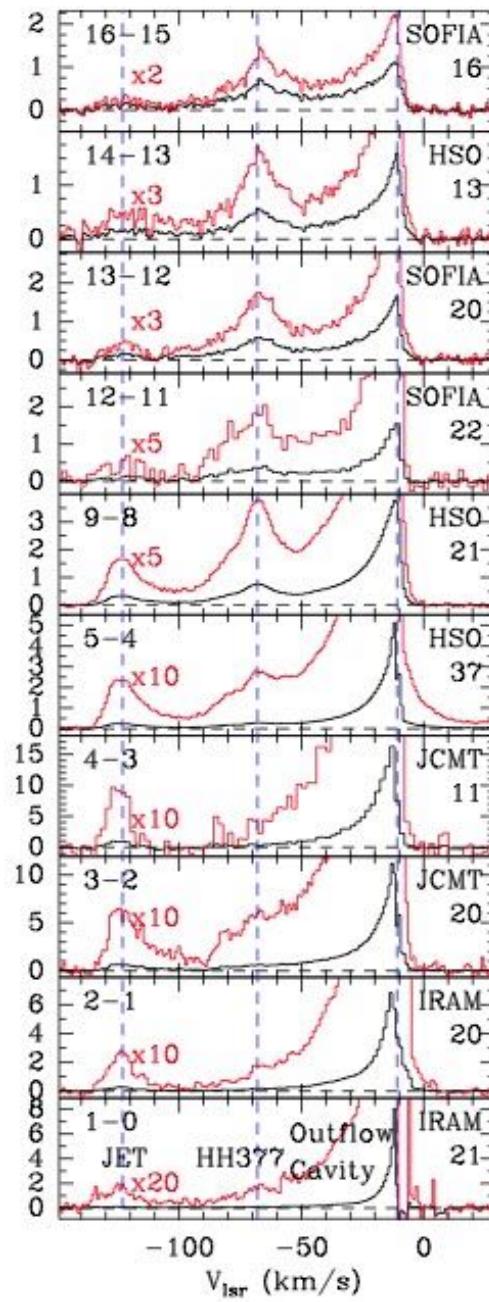
# The SOFIA/GREAT observations: CO



The Cep E outflow at 4.5 μm by *Spitzer*/IRAC

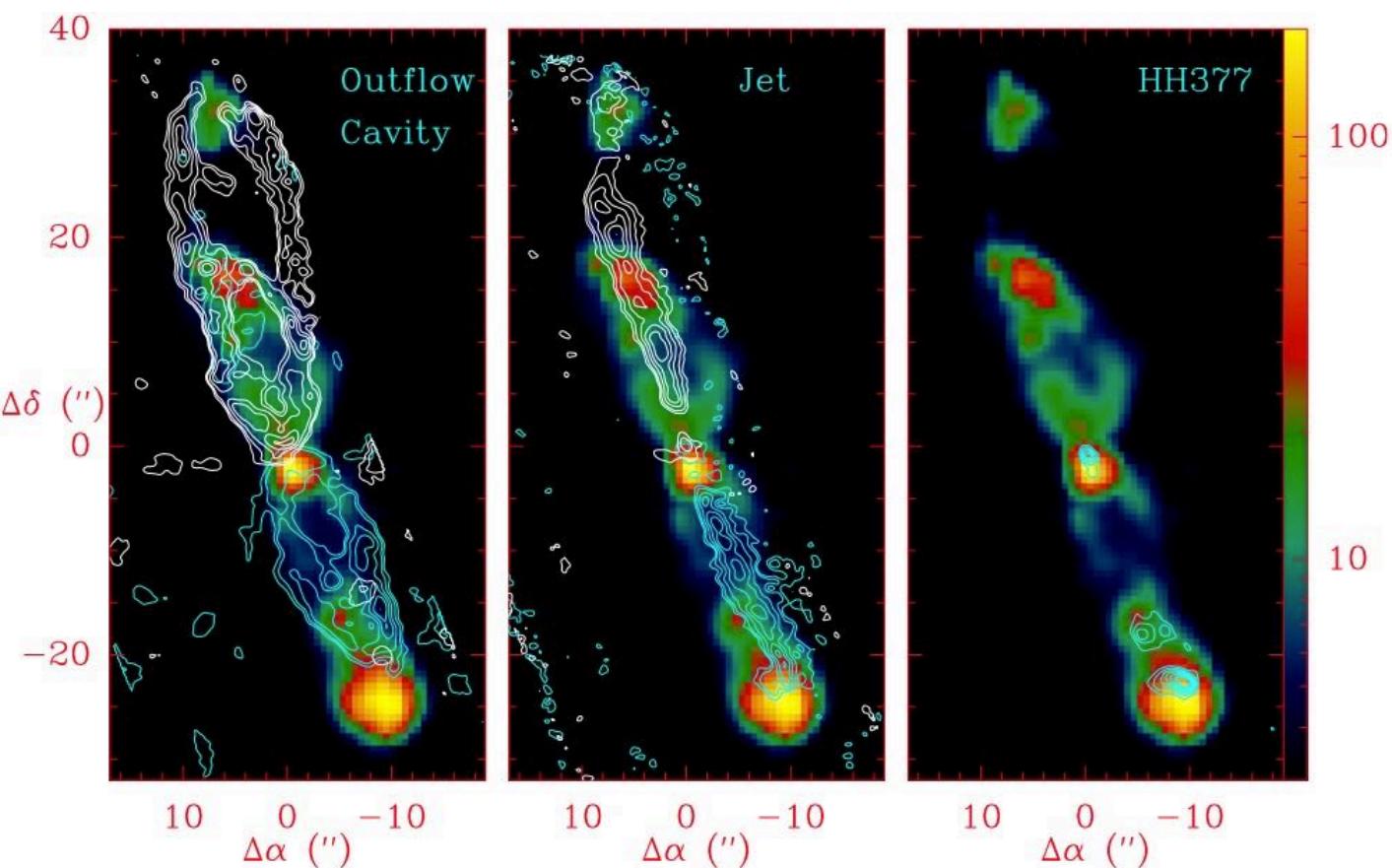
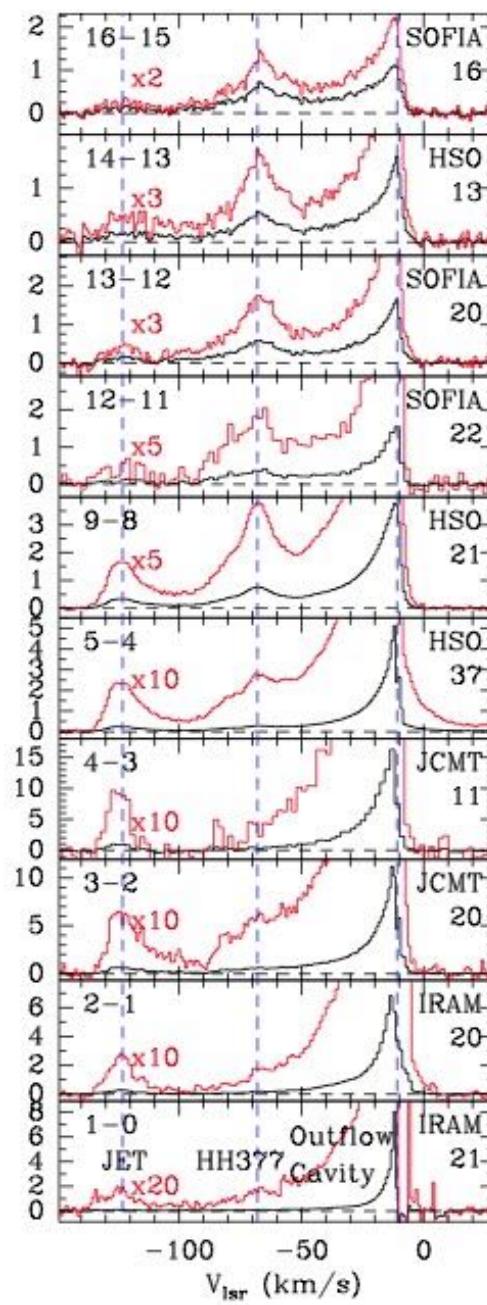
Gomez-Ruiz et al. 2012

# The SOFIA/GREAT observations: CO



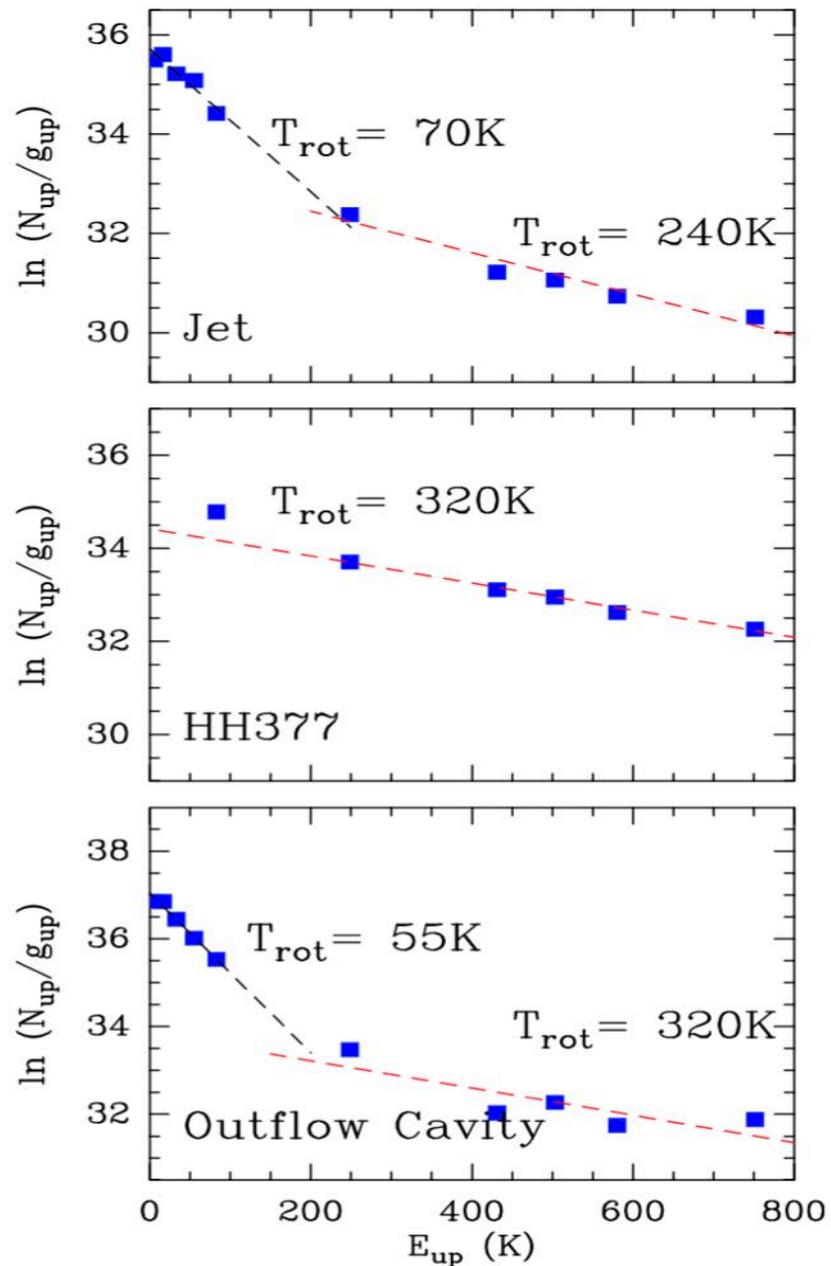
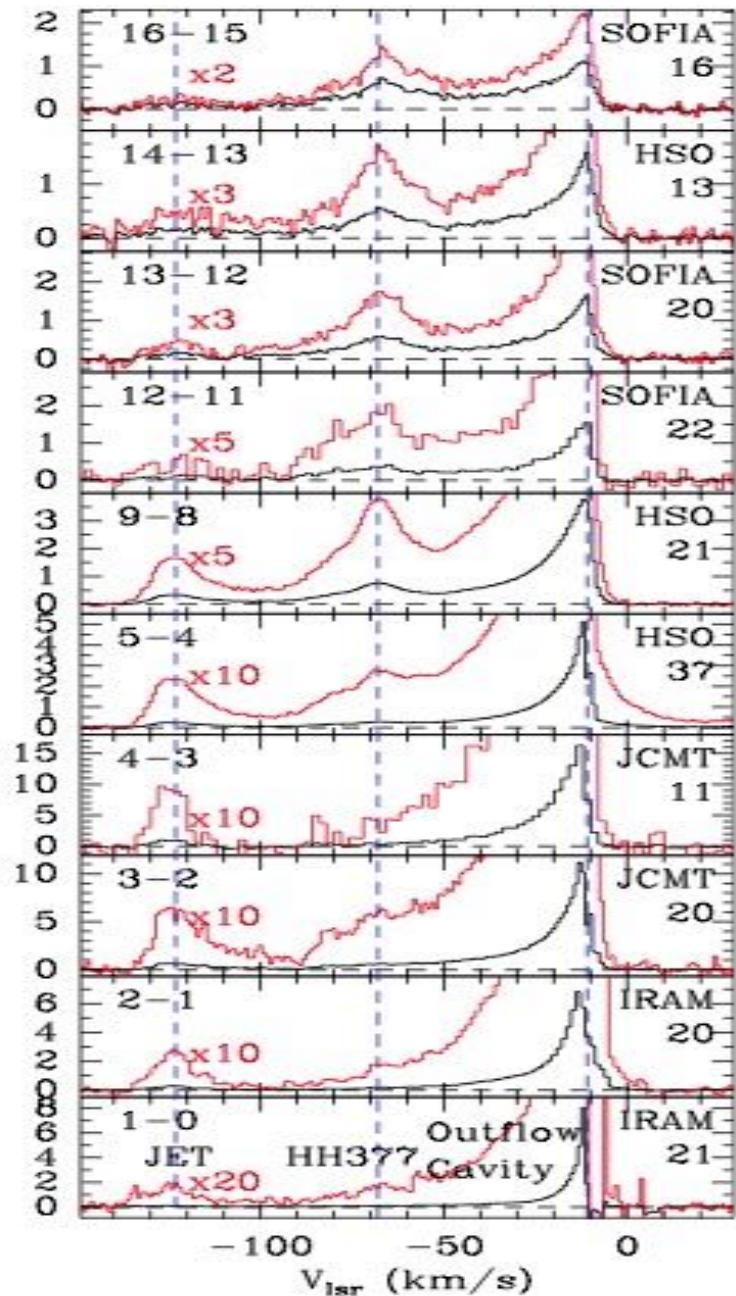
CO spectra: IRAM-30m, JCMT, *Herschel*, SOFIA  
Also:  $^{13}\text{CO}$  (2-1) by IRAM-30m and  $^{13}\text{CO}$  (5-4) by *Herschel*

# The SOFIA/GREAT observations: CO

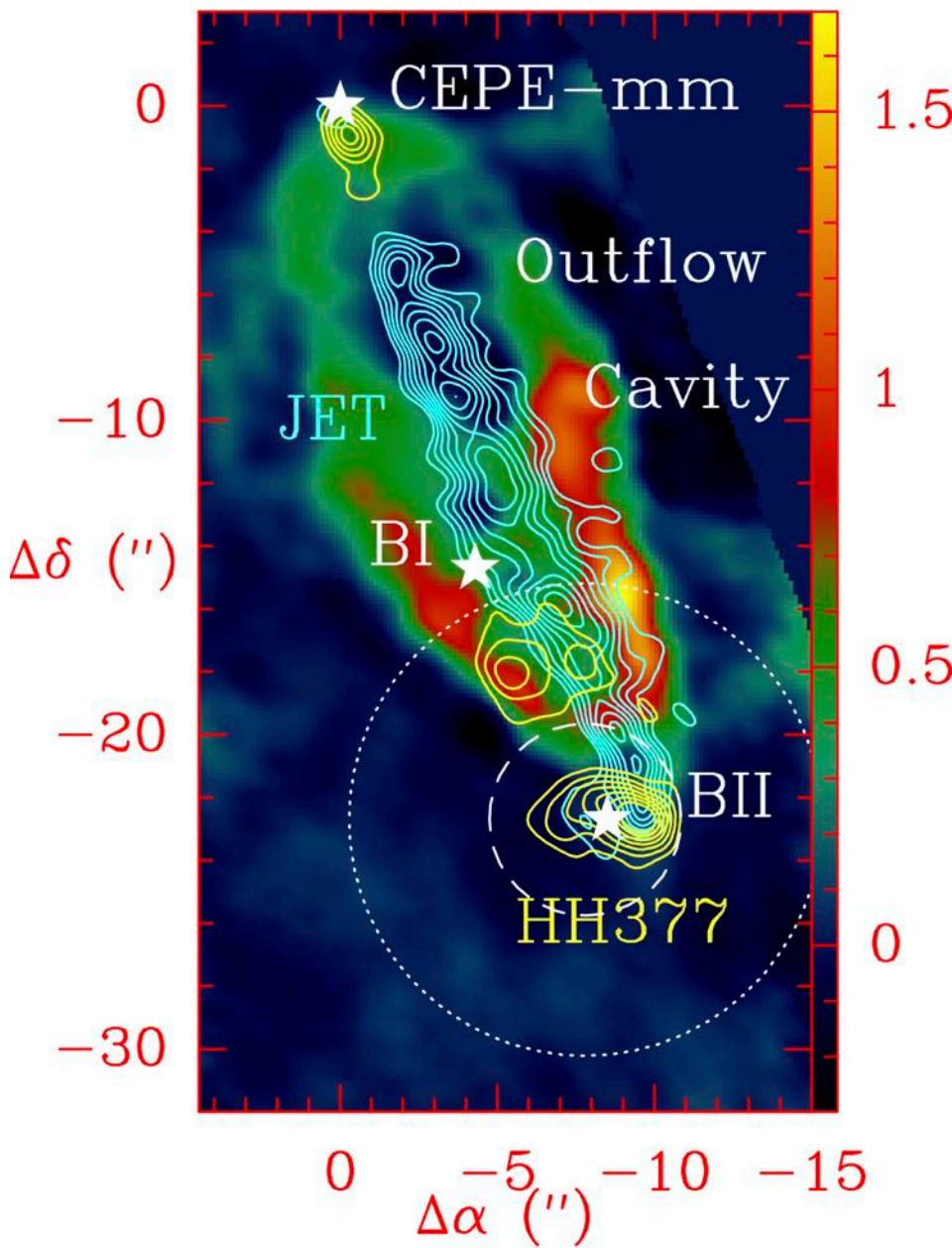


The Cep E outflow in CO (2-1) by the PdBI, Lefloch et al. 2015

# The SOFIA/GREAT observations: CO



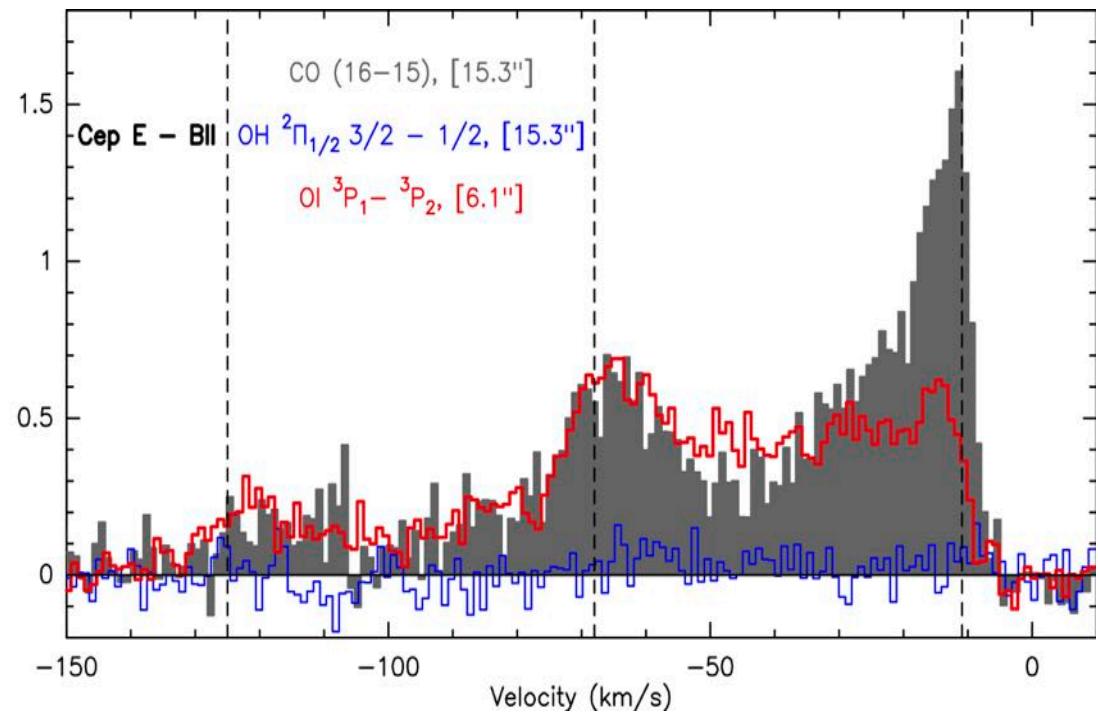
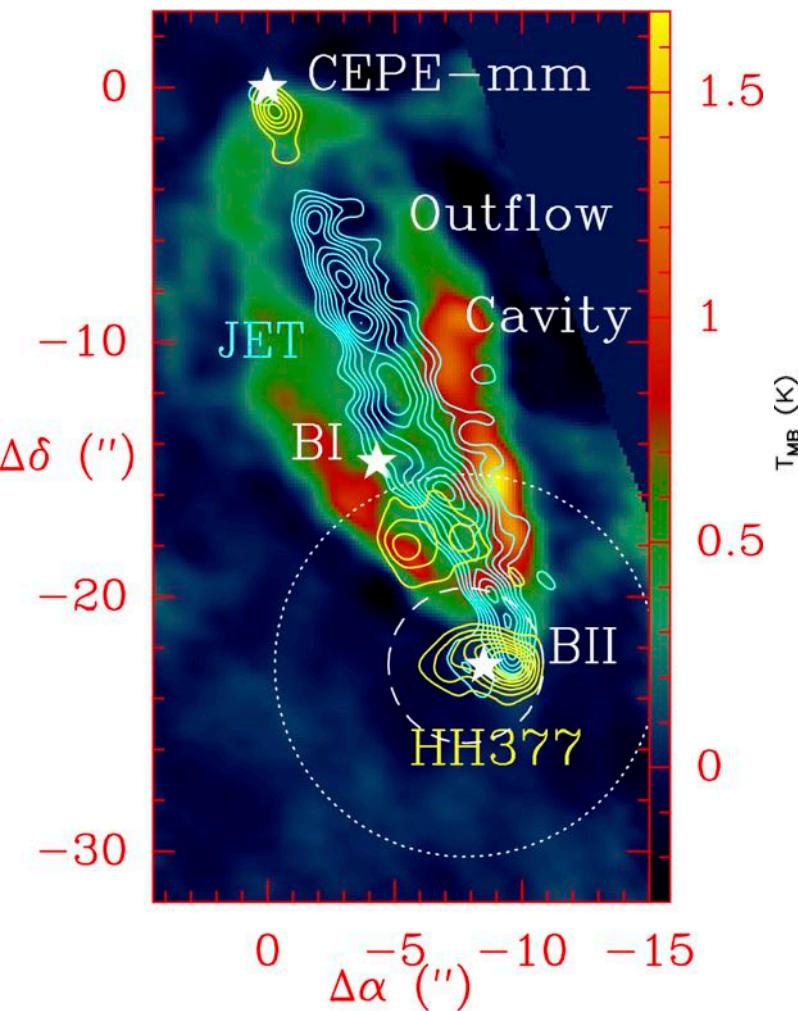
# The SOFIA/GREAT observations: CO



- Jet, cavity, HH 377: CO  
LVG analysis  $\Rightarrow$  (N,  $T_{\text{kin}}$ ) estimates
  - Southern lobe:
    - jet mass:  $0.02 M_{\odot}$ ,  $1.7 M_{\odot}\text{km/s}$
    - outflow cavity mass:  $0.32 M_{\odot}$ ,  $2.8 M_{\odot}\text{km/s}$
- $\Rightarrow$  molecular jet driven bowshock

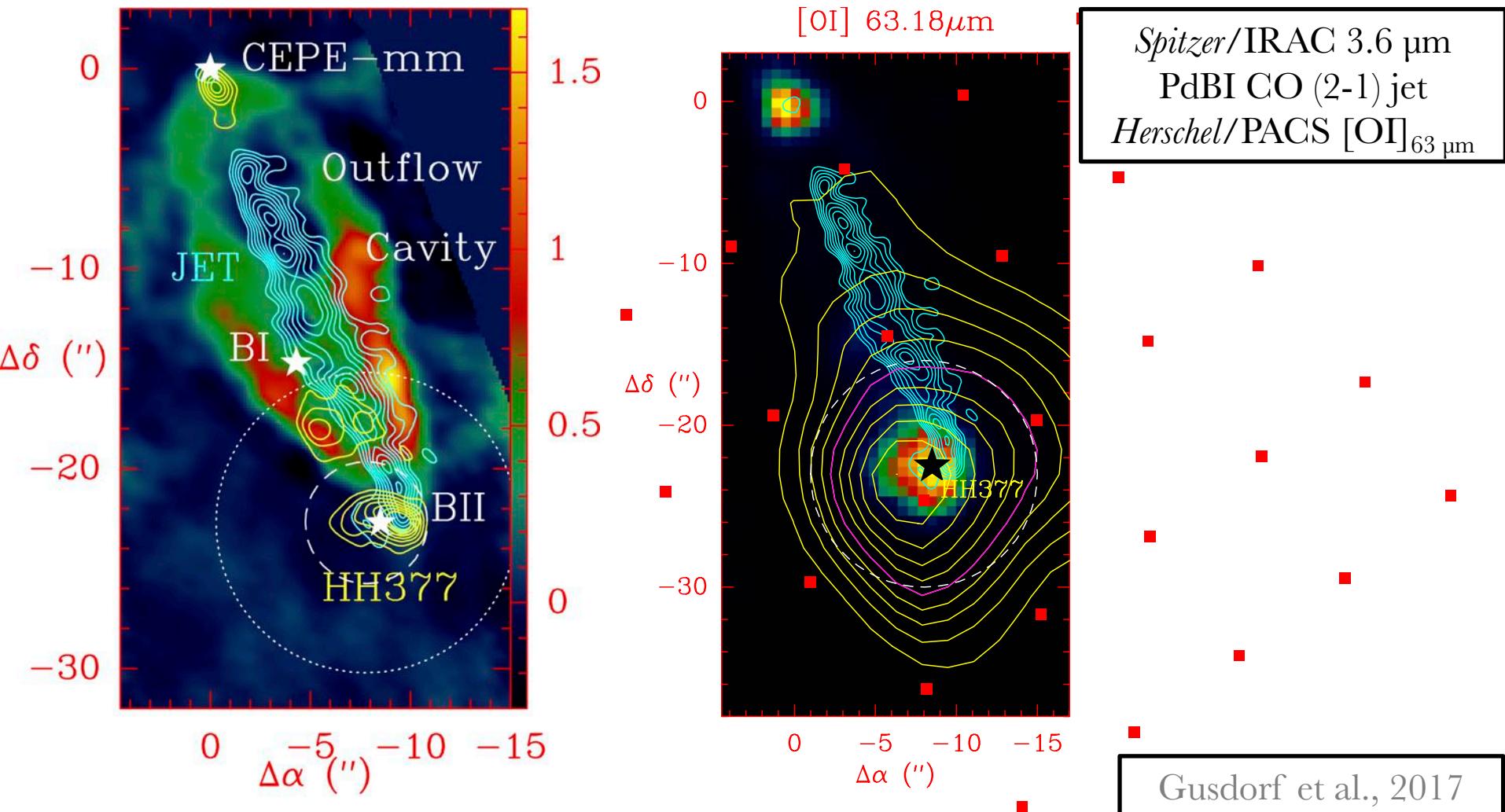
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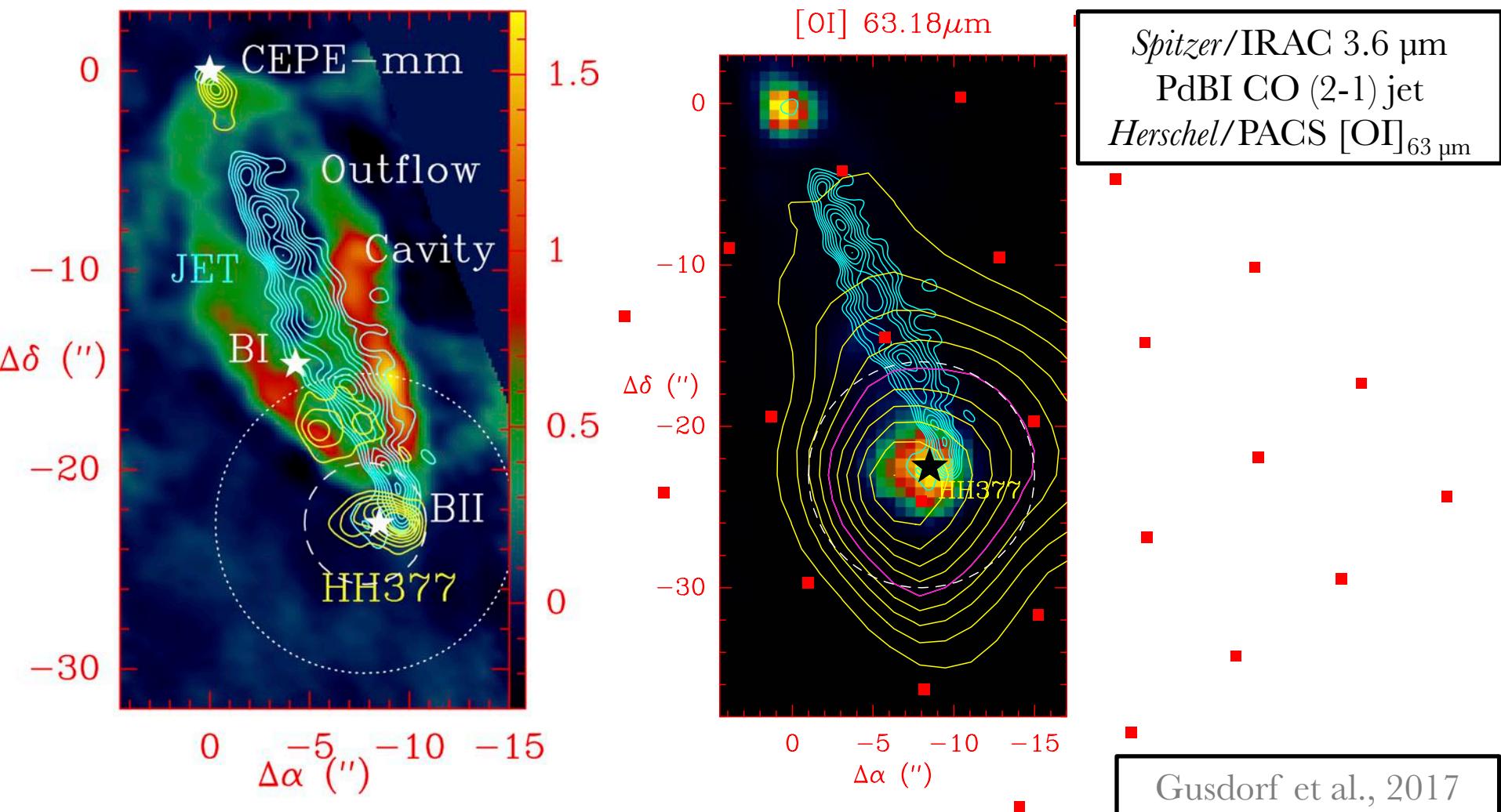


Gusdorf et al., 2017

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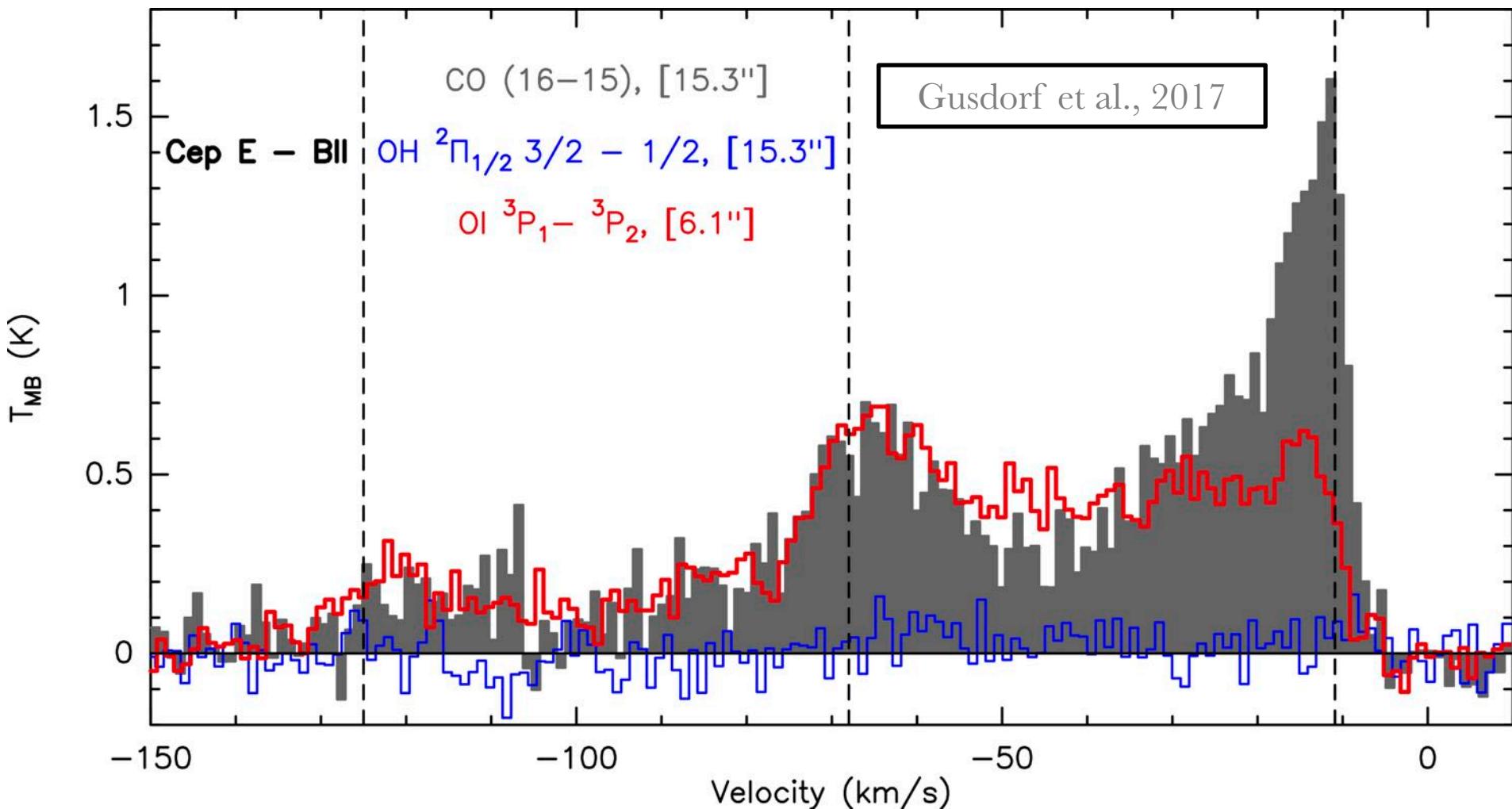
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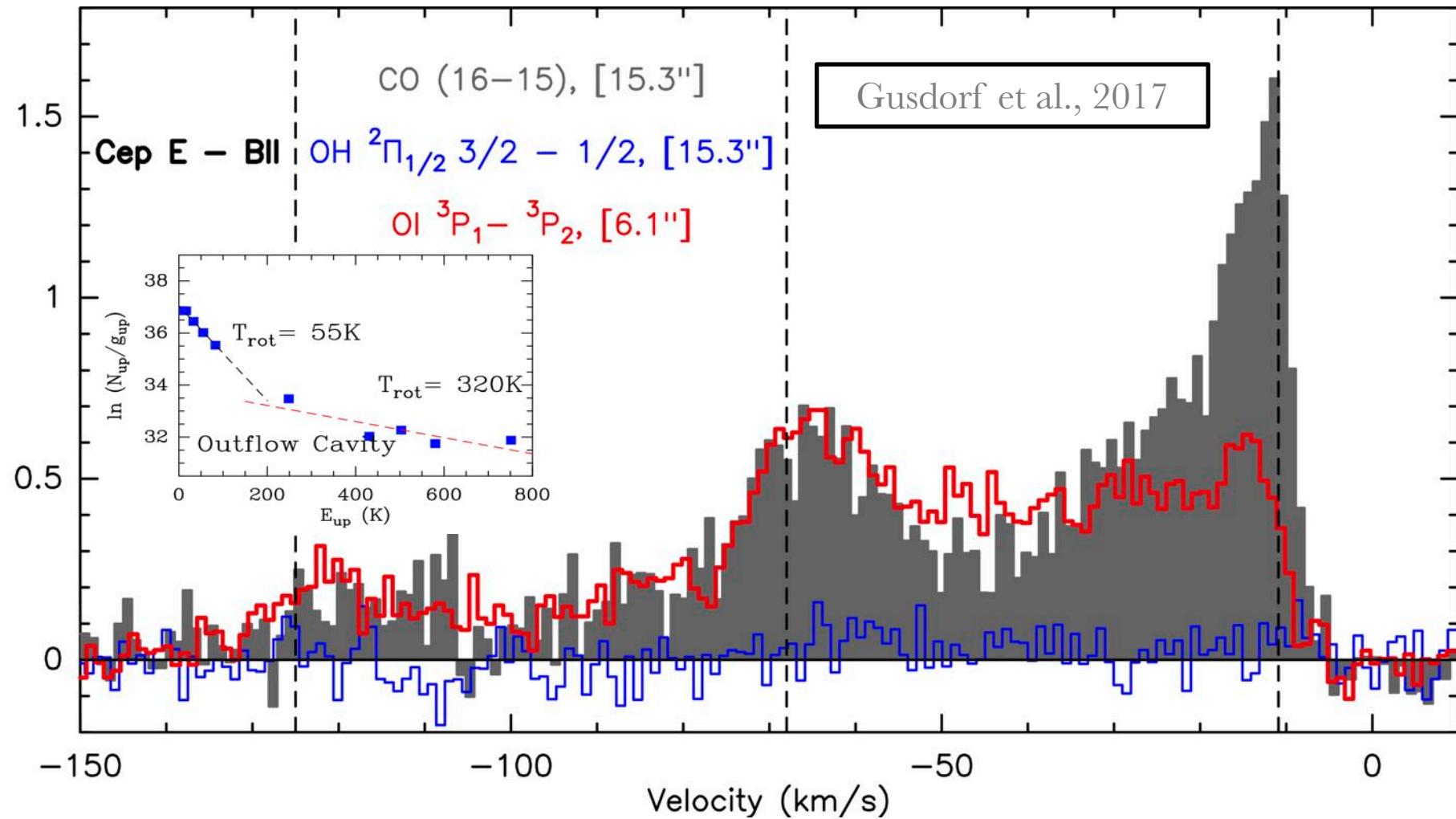
Component	Jet	Bow-shock	Outflow cavity
Filling factor OI ${}^3\text{P}_1 \rightarrow {}^3\text{P}_2$	$0.25 \pm 0.05$	$0.55 \pm 0.15$	<1
Filling factor CO (16–15)	0.09	0.06	0.15

# The SOFIA/GREAT observations: [OI]

31



# The SOFIA/GREAT observations: [OI]



# The SOFIA/GREAT observations: [OI]

	Component	Jet	Bow-shock	outflow cavity
Low-excitation assumption	$N(\text{CO}) (10^{16} \text{ cm}^{-2})$	9.0	–	70.0
	$T_{\text{kin}} (\text{K})$	80–100	–	55–85
	$n (\text{cm}^{-3})$	$(0.5\text{--}1) \times 10^5$	–	$(1\text{--}8) \times 10^5$
	$N(\text{OI}) (10^{16} \text{ cm}^{-2})$	$24.6 \pm 8.5$	–	$>24.8$
	$N(\text{OI})/N(\text{CO})$	$2.7 \pm 0.9$	–	$>0.4$
High-excitation assumption	$N(\text{CO}) (10^{16} \text{ cm}^{-2})$	1.5	10.0	6.0
	$T_{\text{kin}} (\text{K})$	400–750	400–500	500–1500
	$n (\text{cm}^{-3})$	$(0.5\text{--}1) \times 10^6$	$(1.0\text{--}2.0) \times 10^6$	$(1\text{--}5) \times 10^6$
	$N(\text{OI}) (10^{16} \text{ cm}^{-2})$	$4.0 \pm 1.0$	$2.1 \pm 0.6$	$>4.9$
	$N(\text{OI})/N(\text{CO})$	$2.7 \pm 0.6$	$0.2 \pm 0.1$	$>0.8$

# The SOFIA/GREAT observations: [OI]

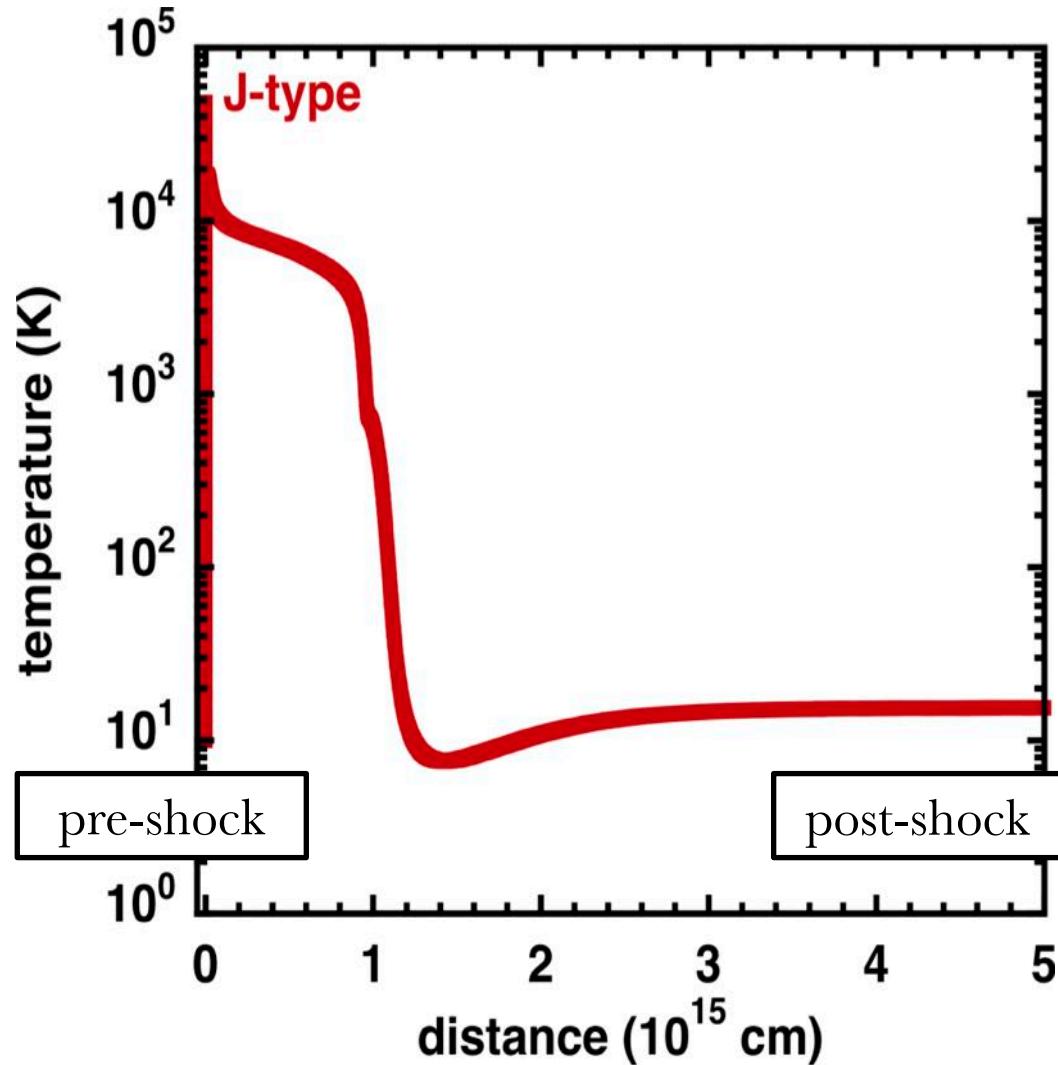
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# Shock modelling in Cepheus E

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$$n_H = 10^4 \text{ cm}^{-3}; v_s = 30 \text{ km/s}$$

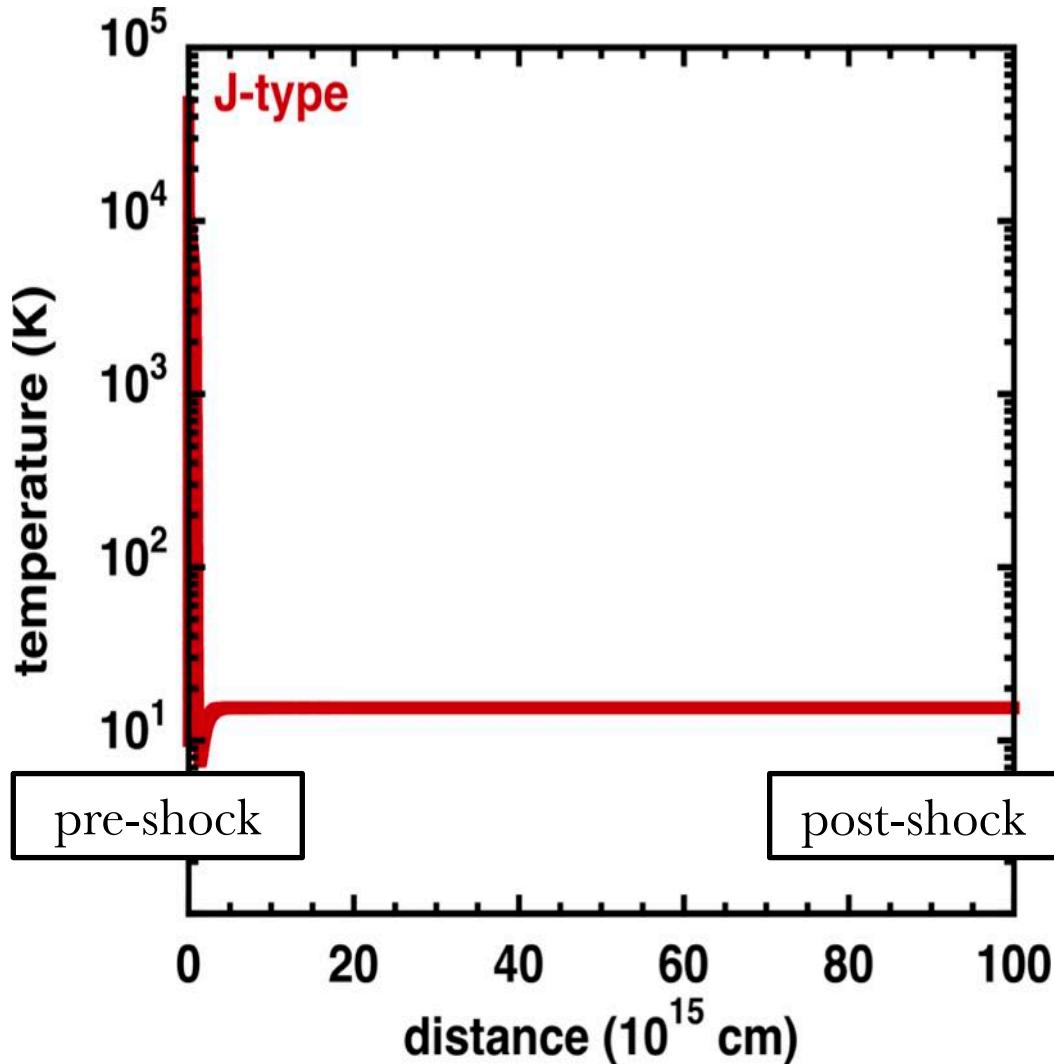
- J shock (Jump) ;
- $B = 10 \mu\text{G}$
- $v_s > v_{\text{critical}}$
- impulse heating ;
- single fluid



# Shock modelling in Cepheus E

$$n_H = 10^4 \text{ cm}^{-3}; v_s = 30 \text{ km/s}$$

- J shock (Jump) ;
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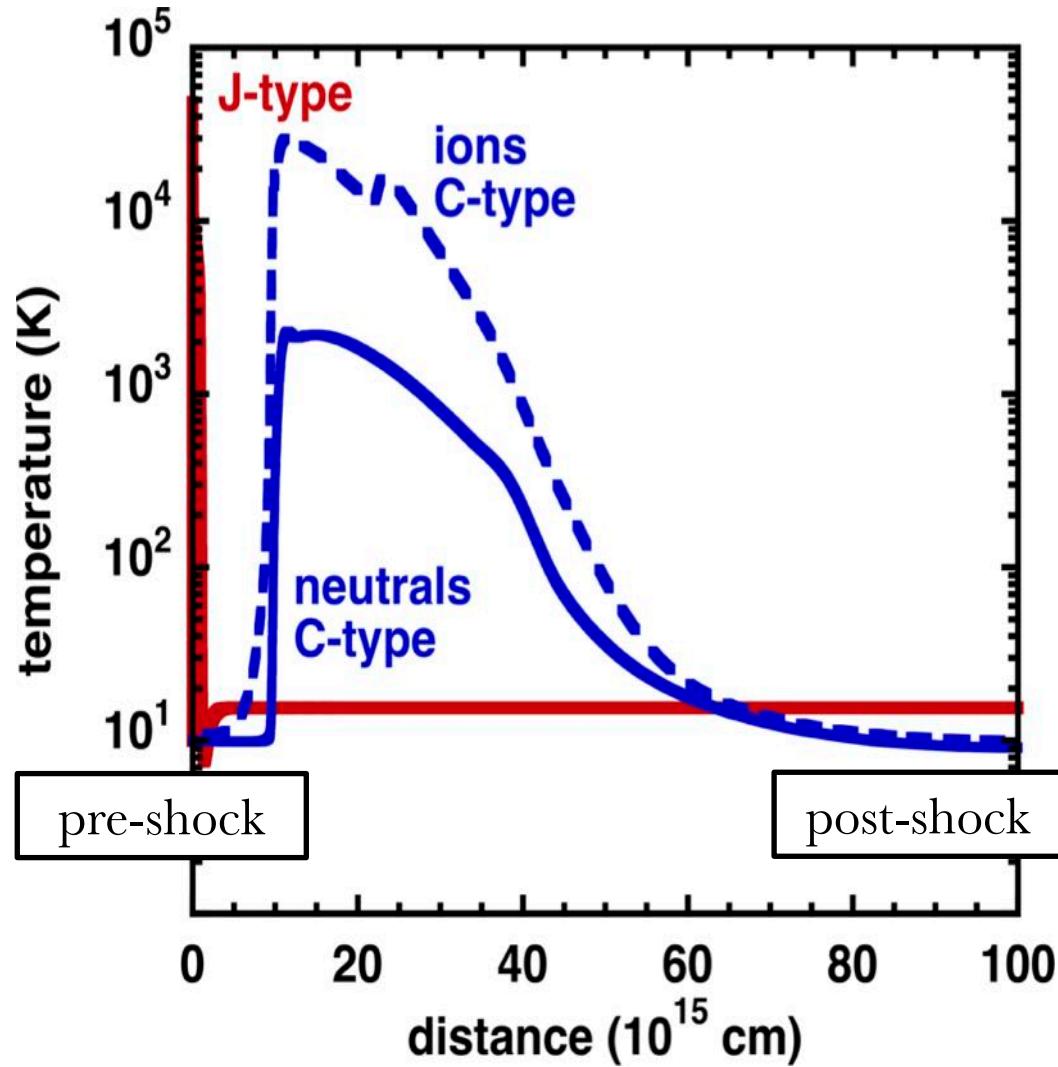


# Shock modelling in Cepheus E

$$n_H = 10^4 \text{ cm}^{-3}; v_s = 30 \text{ km/s}$$

- J shock (Jump) ;  
 $B = 10 \mu\text{G}$   
 $v_s > v_{\text{critical}}$   
 impulse heating ;  
 single fluid

- C shock (Continuous) ;  
 $B = 100 \mu\text{G}$   
 $v_s < v_{\text{critical}}$   
 ion-neutral friction ;  
 multi-fluid

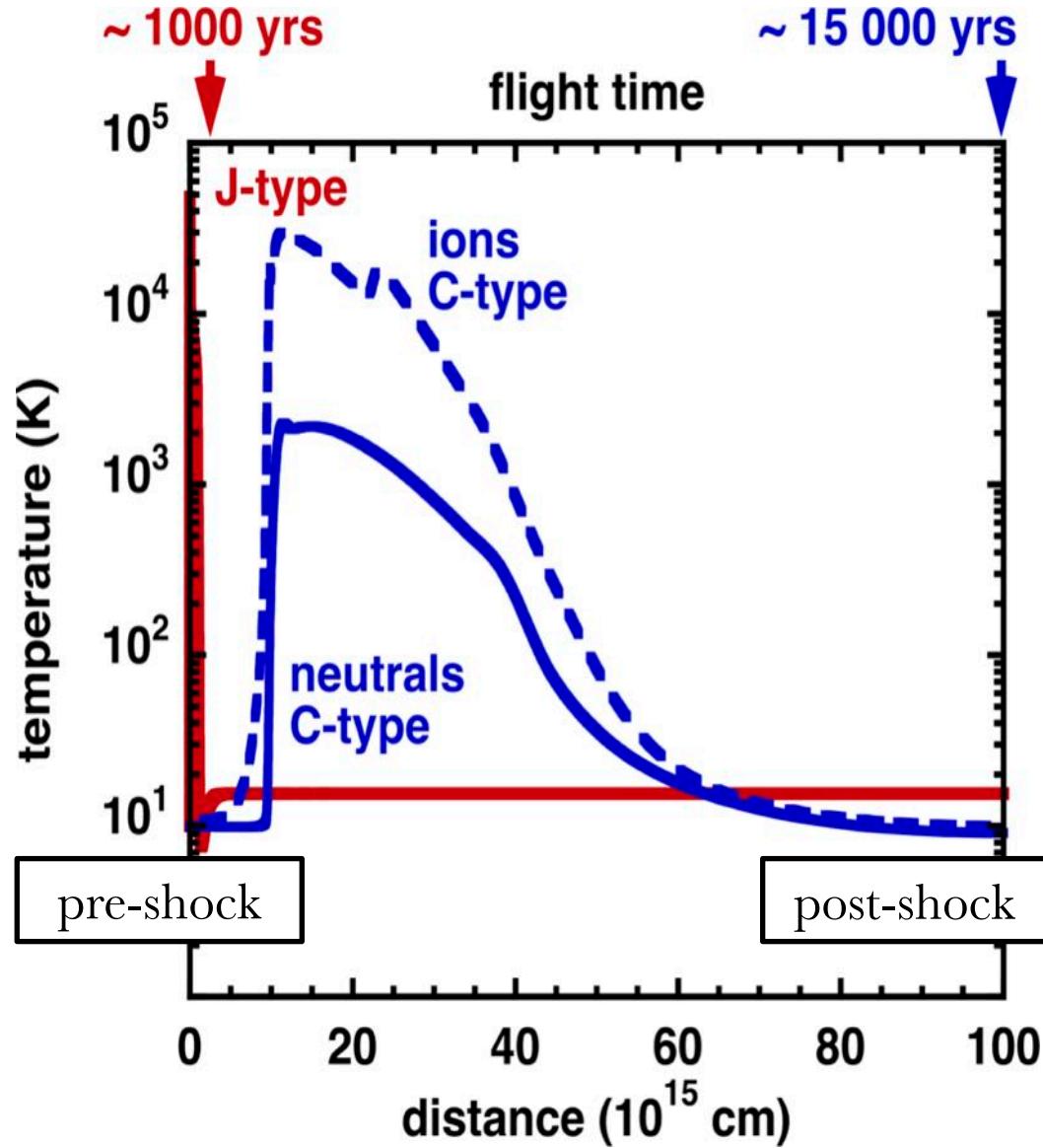


# Shock modelling in Cepheus E

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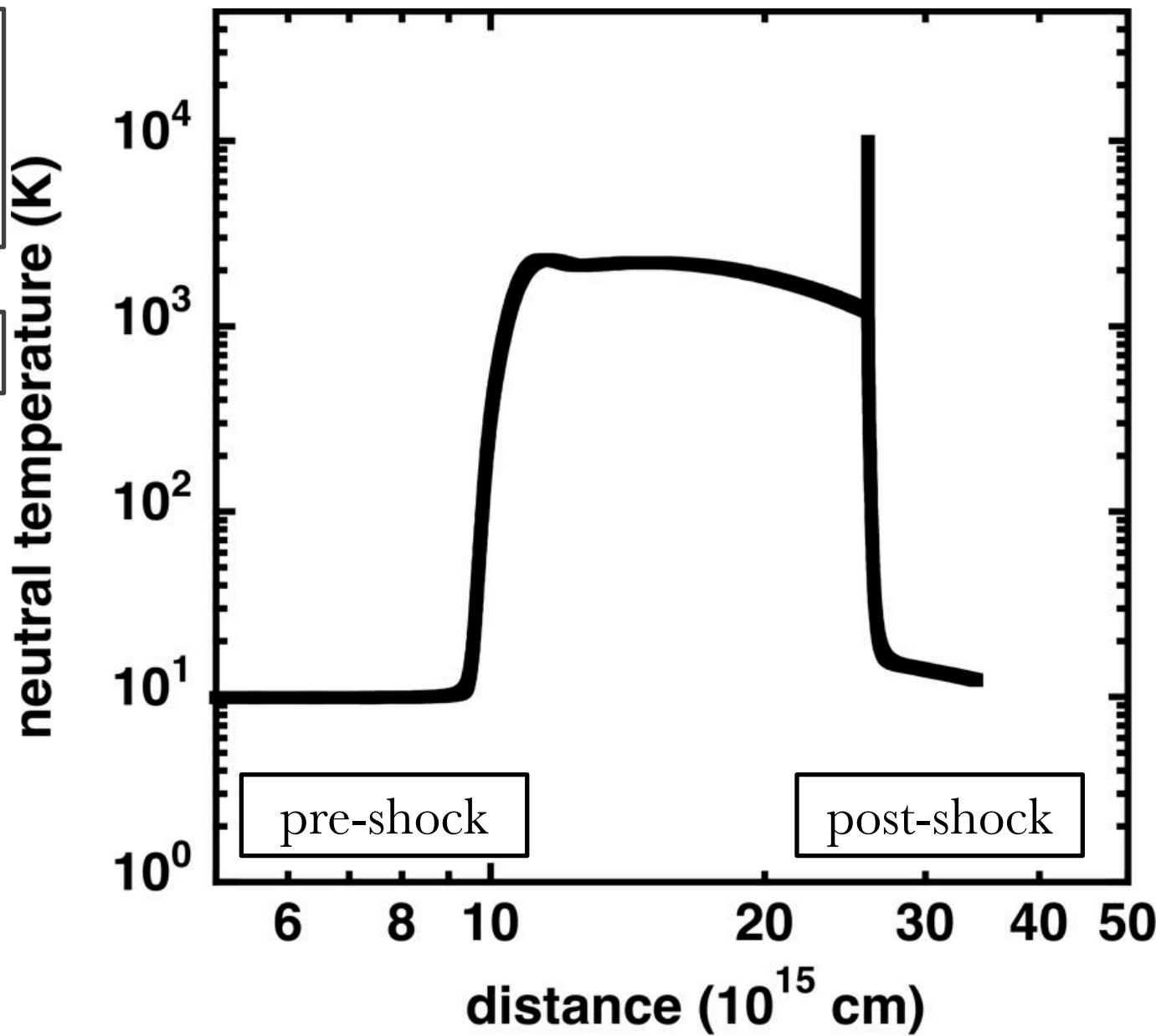
# Shock modelling in Cepheus E

$n_H = 10^4 \text{ cm}^{-3}$

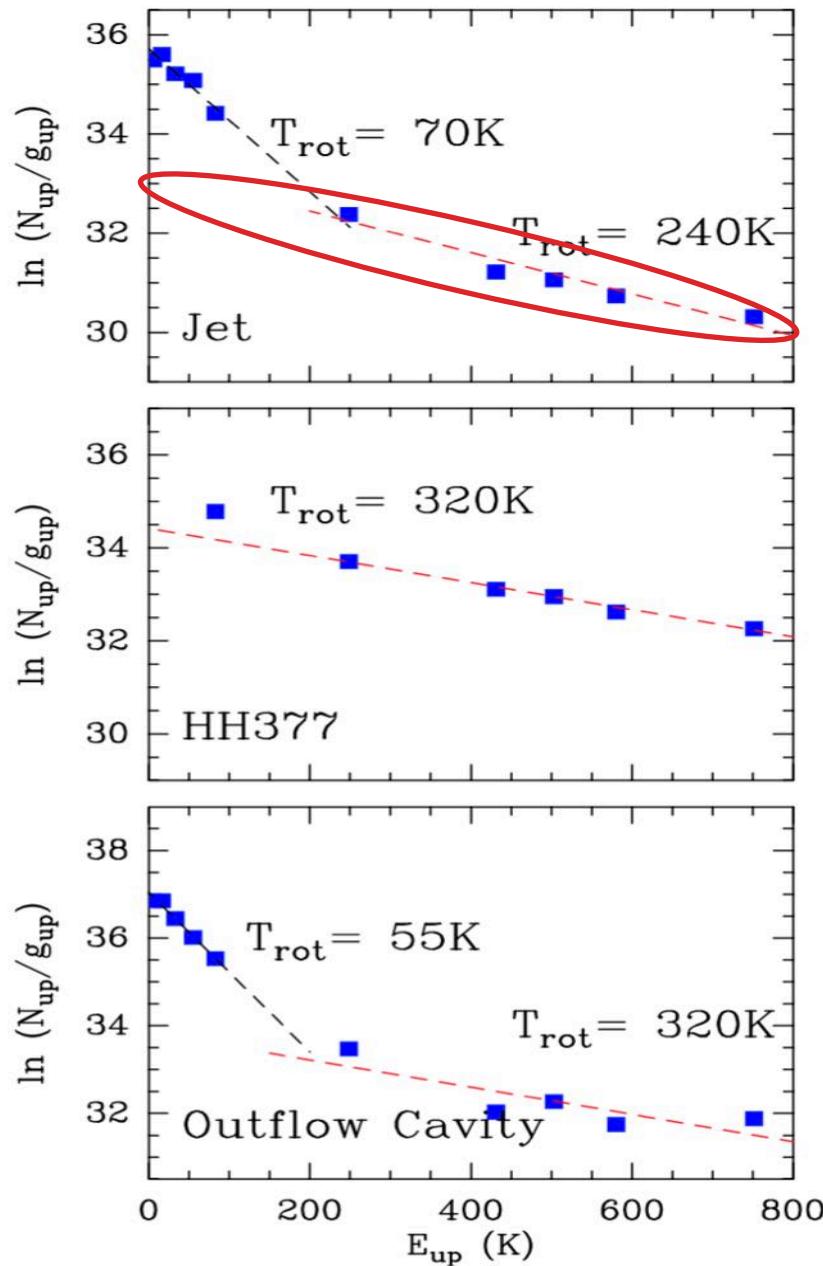
$v_s = 30 \text{ km/s}$

age 2000 yrs

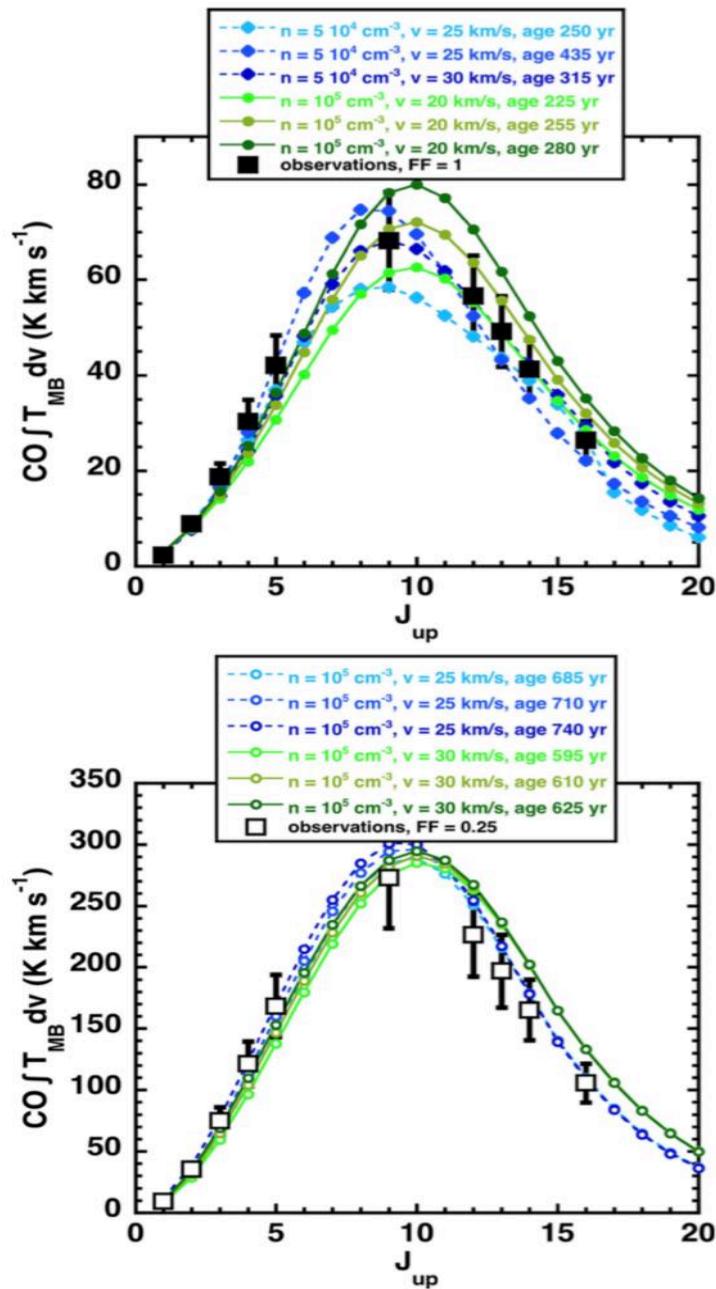
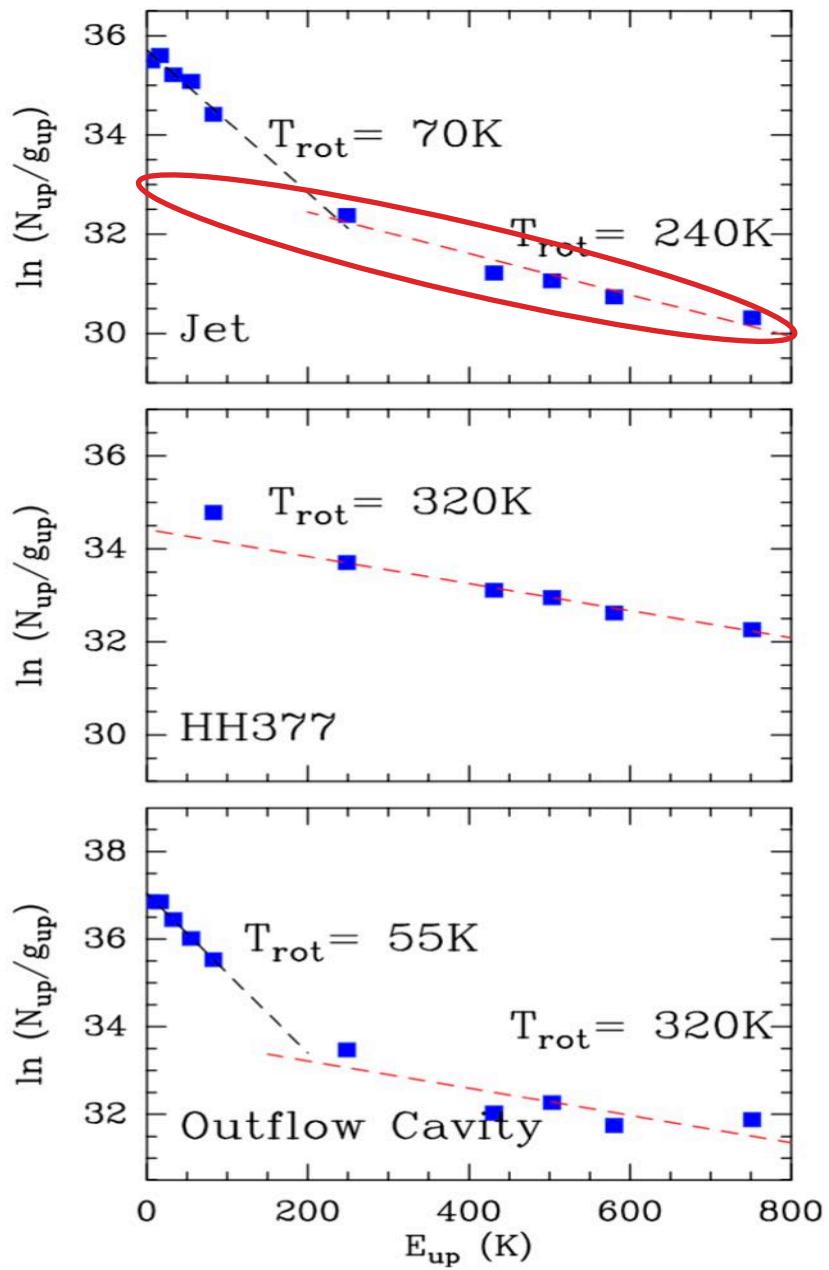
CJ shock



# Shock modelling in Cepheus E



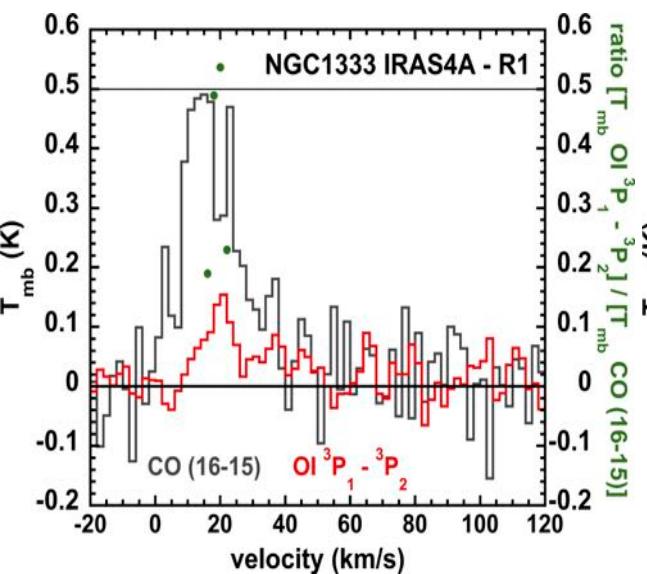
# Shock modelling in Cepheus E



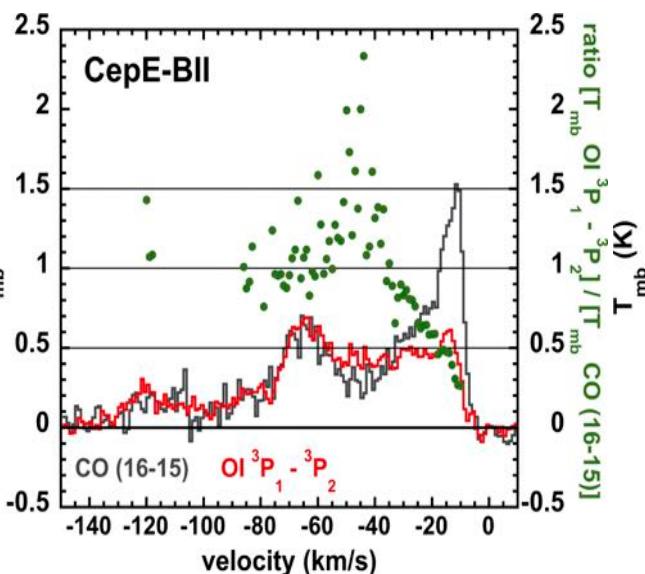
# Shock modelling in Cepheus E

From low-mass...

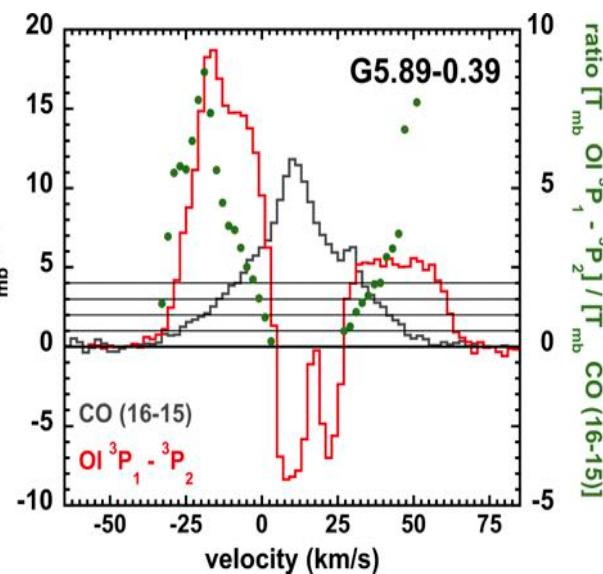
to high-mass driving proto-stars



Kristensen et al., 2017



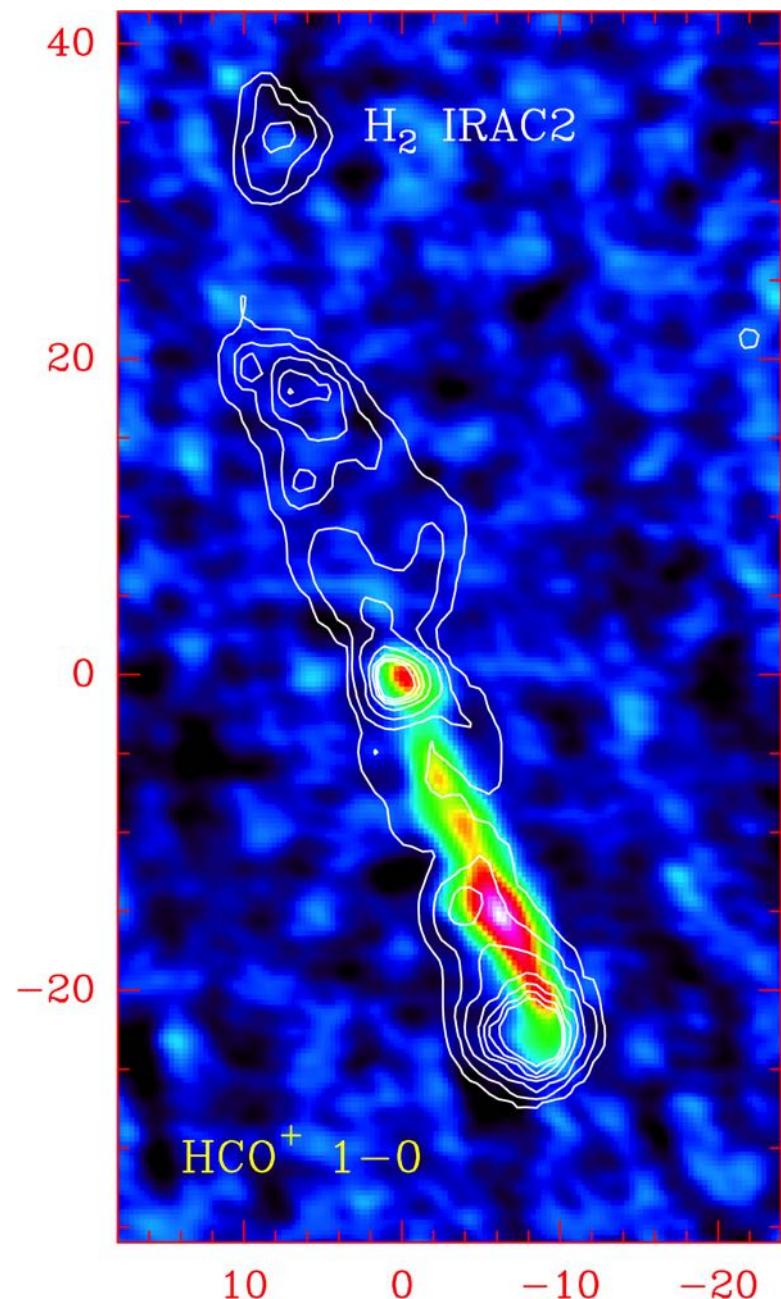
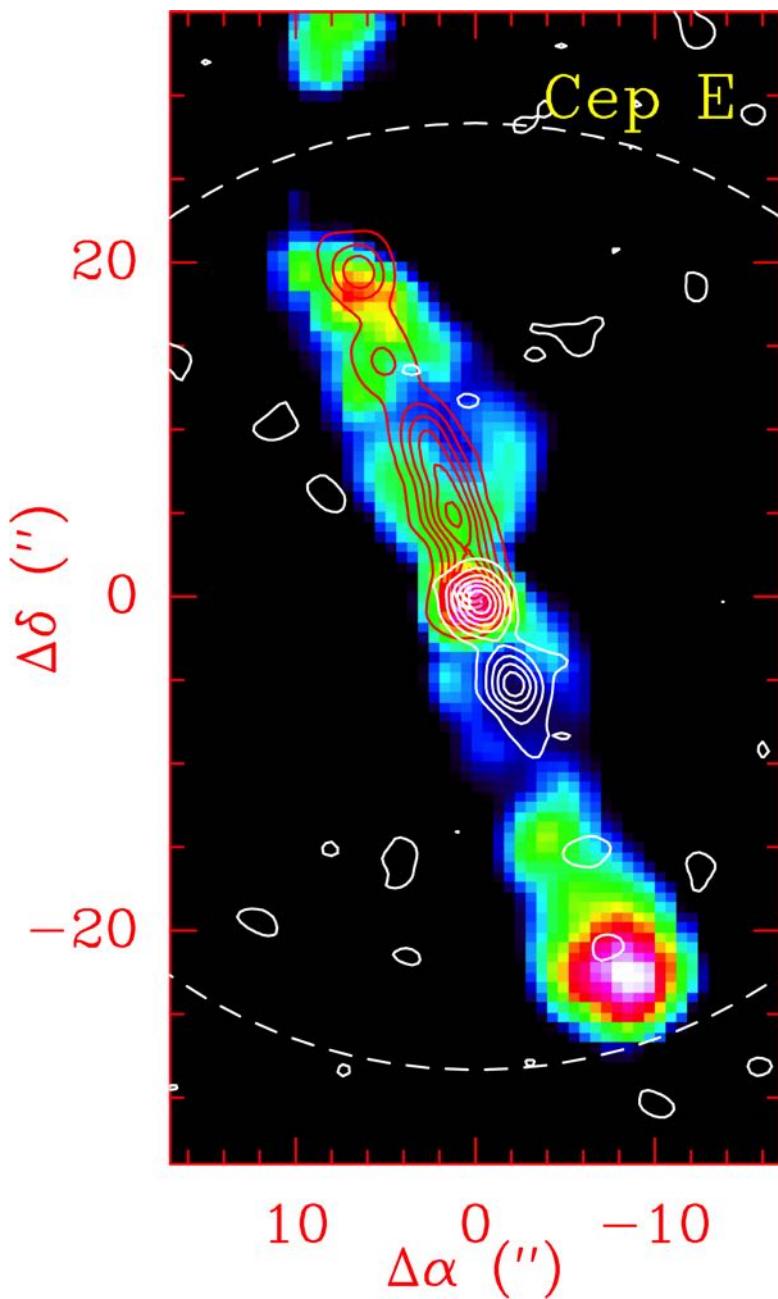
Gusdorf et al., 2017



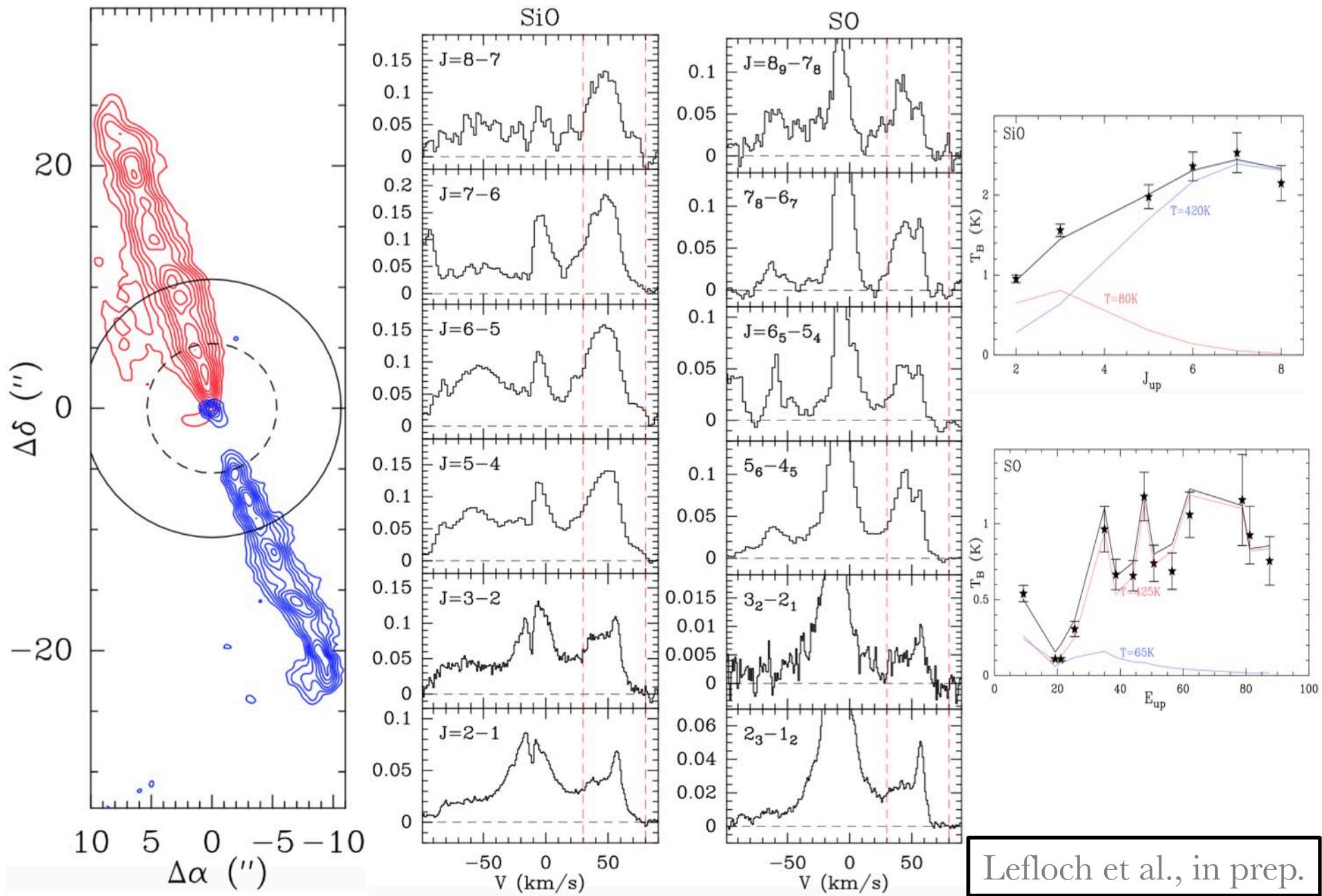
Leurini et al., 2015

# Perspectives

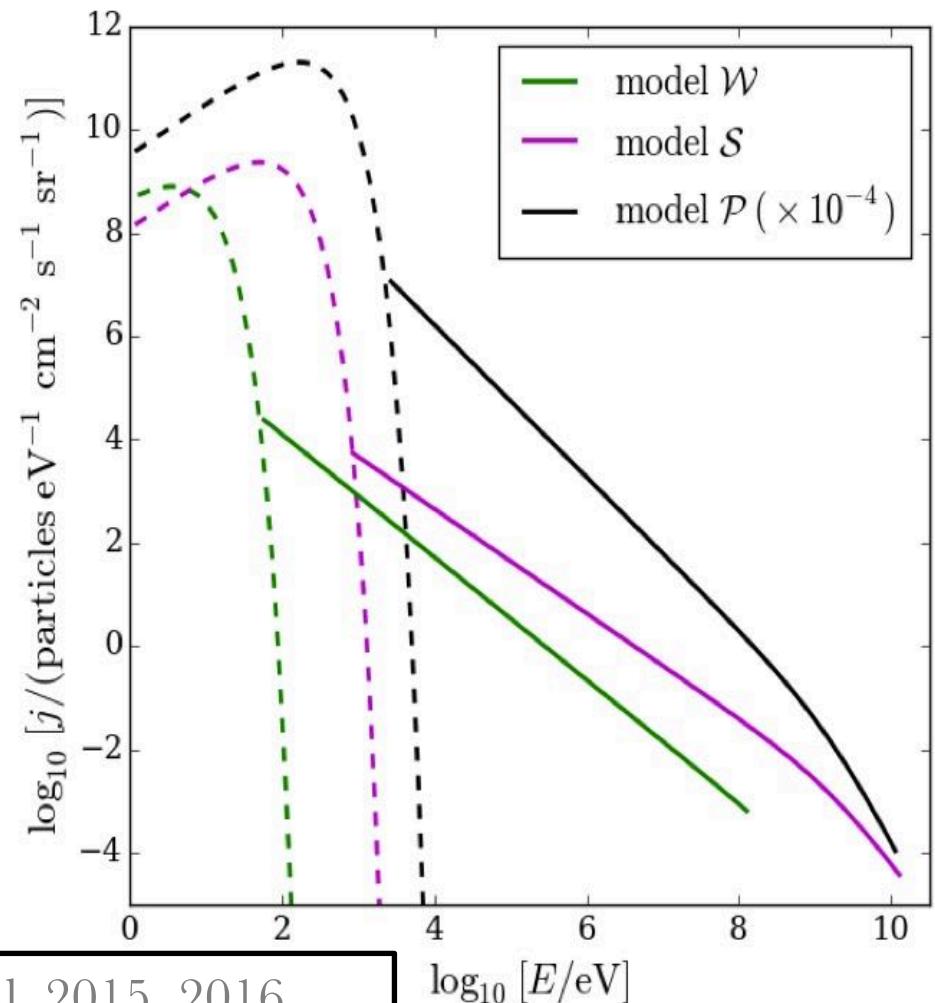
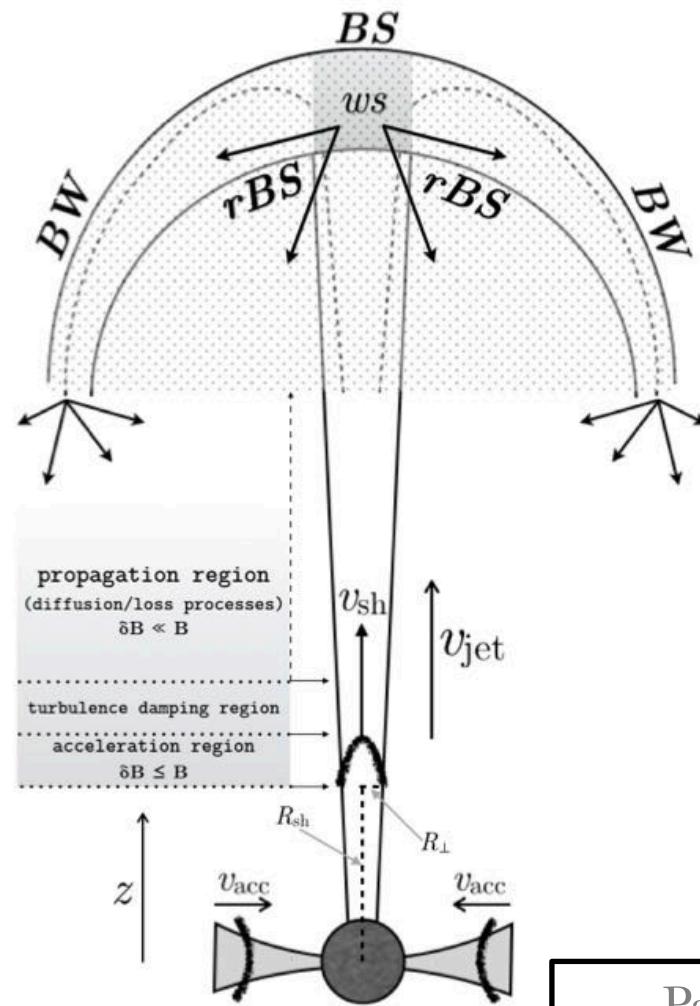
# Perspectives (1): astrochemistry (1)



# Perspectives (1): astrochemistry (2)



# Perspectives (2): cosmic rays acceleration

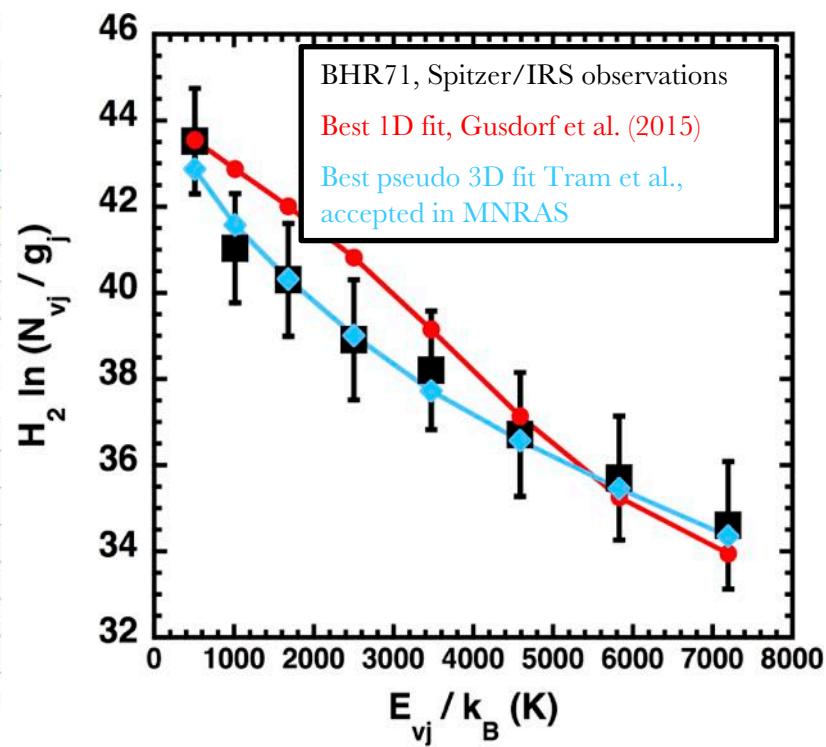
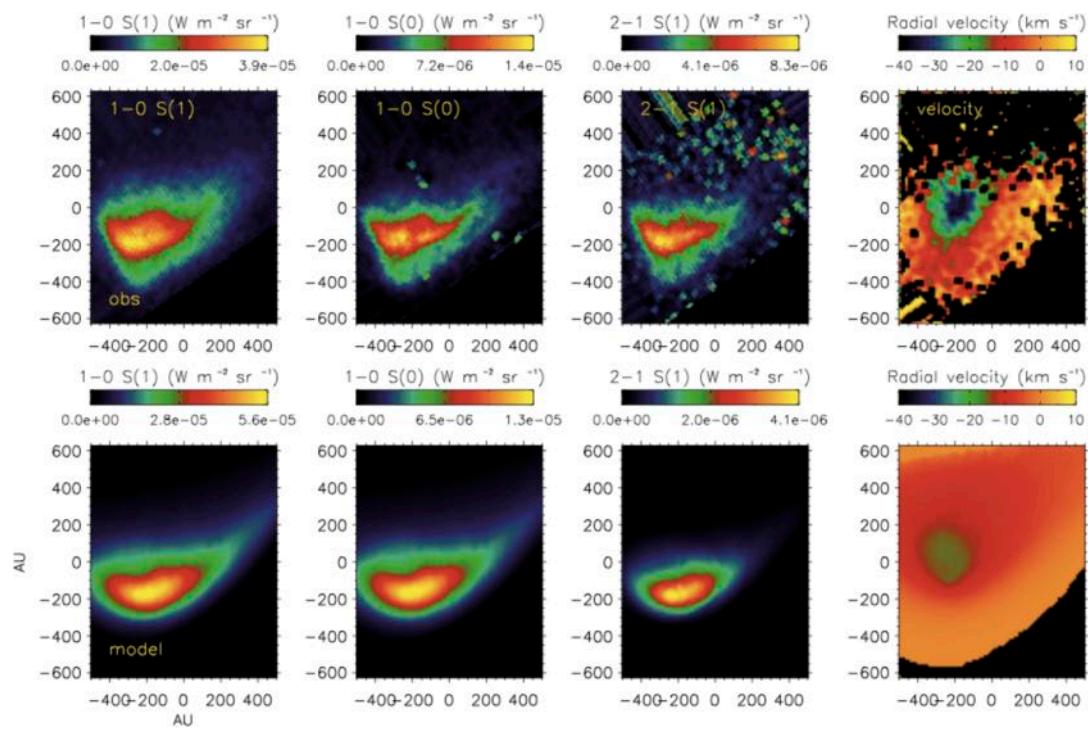
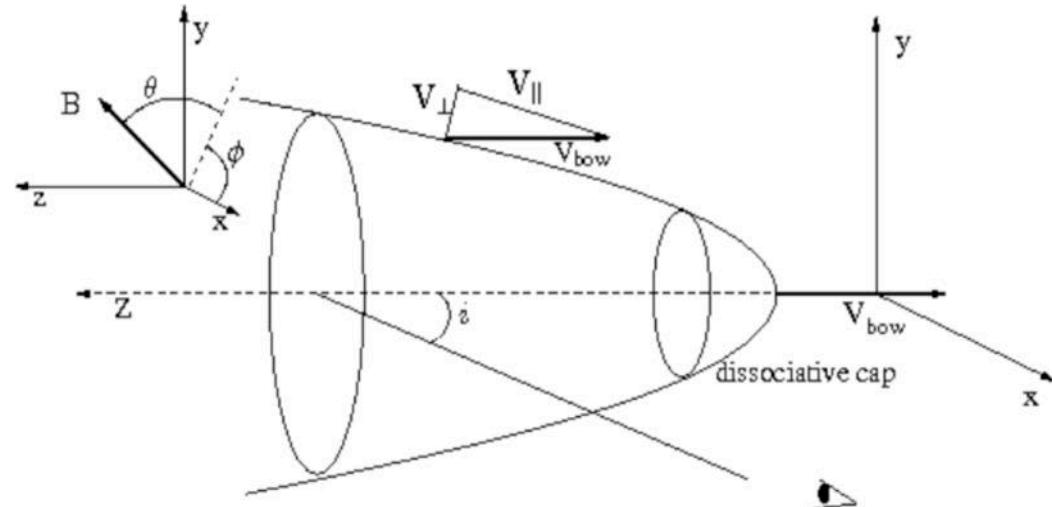


Padovani et al. 2015, 2016

model	$U$ [km s <sup>-1</sup> ]	$B$ [G]	$n_{\text{H}}$ [cm <sup>-3</sup> ]	$x$	$T$ [10 <sup>4</sup> K]	$r$	$E_{\text{max}}$ [GeV]	$\tilde{P}_{\text{CR}}$ [10 <sup>-2</sup> ]	$\lambda$	$p_{\text{inj}}$ [MeV/c]	$p_{\text{max}}$ [GeV/c]
$\mathcal{W}$	40	$5 \times 10^{-5}$	$10^5$	0.33	1	2.977	0.13	0.88	4.010	0.306	0.505
$\mathcal{S}$	160	$10^{-3}$	$6 \times 10^5$	0.60	1	3.890	12.9	4.70	4.062	1.146	13.762
$\mathcal{P}$	260	5	$1.9 \times 10^{12}$	0.30	94	2.290	11.4	0.03	3.950	2.058	12.306

# Perspectives (3): shock models

- Kristensen et al. 2008,  
Gustafsson et al. 2010  
Tram et al., accepted in MNRAS:  
 $n_H = 10^5 \text{ cm}^{-3}$ ;  
 $v_{\text{bow}} = 50 \text{ km s}^{-1}$ ;  
 $b = 3$ ;  
 $\beta = 1.7$ ;  $r_0 = 610 \text{ AU}$ ;  
 $\theta = 15^\circ$ ;  $\Phi = 180^\circ$ ;  $i = 50^\circ$



Thanks for your attention !