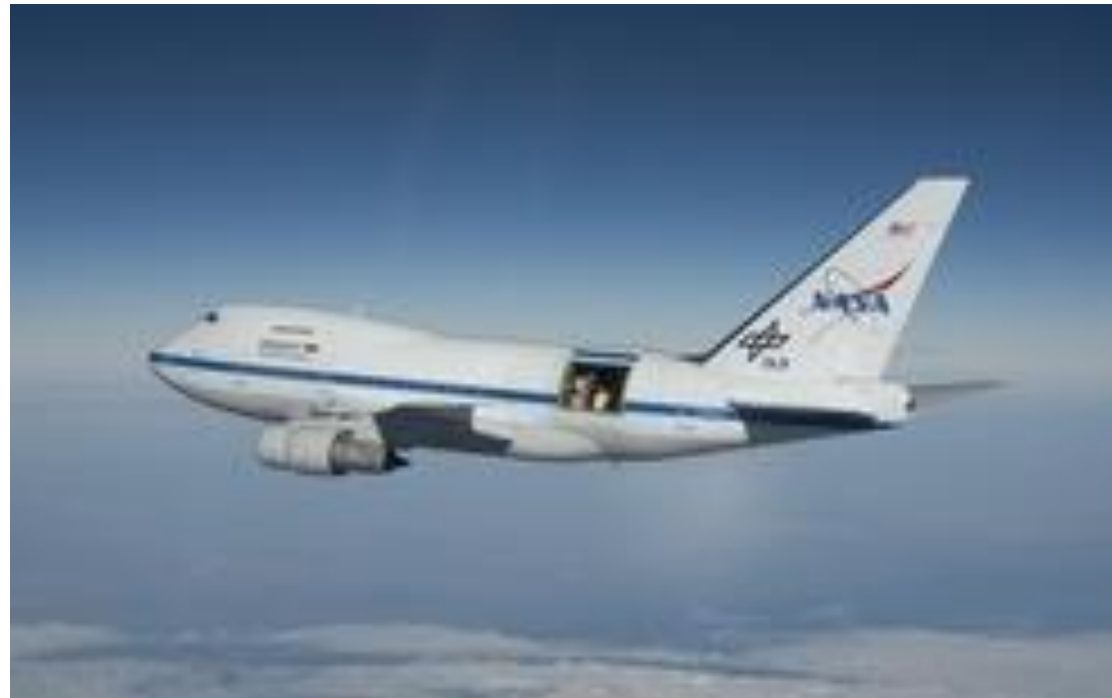


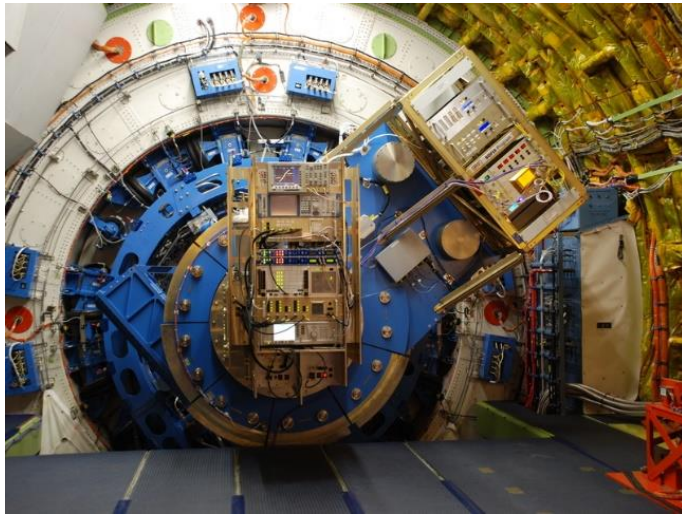
Discovery of the Linear Carbon Chain Molecules ^{13}CCC and C^{13}CC Towards SgrB2(M)

Thomas Giesen

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SOFIA GREAT Receiver



Front-End: LFAH

Two 7-pixel arrays at two polarizations (H,V). Both polarizations can be tuned to the same frequency (LFAH range), or two separate frequencies on a best effort basis.

Frequencies (GHz)

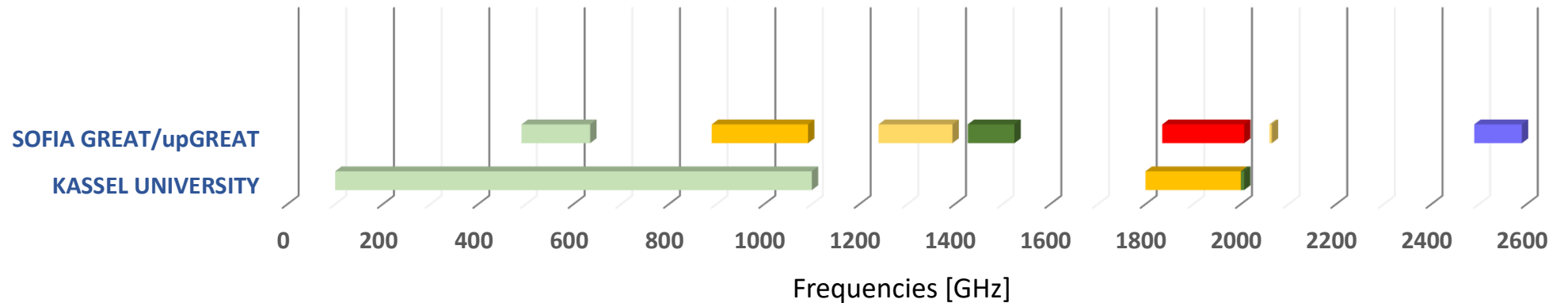
1835–2007

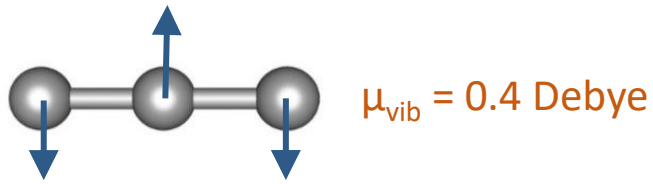
Lines of Interest

[CII] 158 μm , CO, OH, $^2\Pi_{1/2}$, ^{12}CH , ^{13}CH

NEW: C_3 ^{13}CCC C^{13}CC

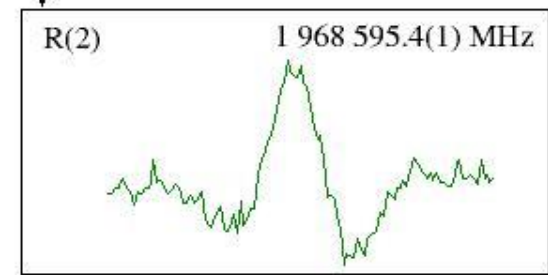
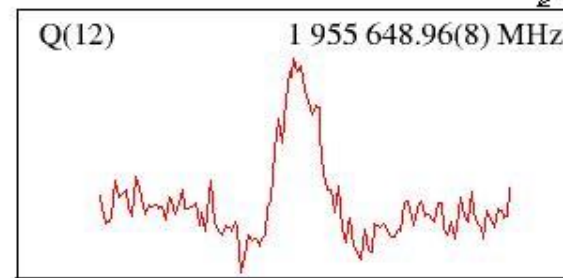
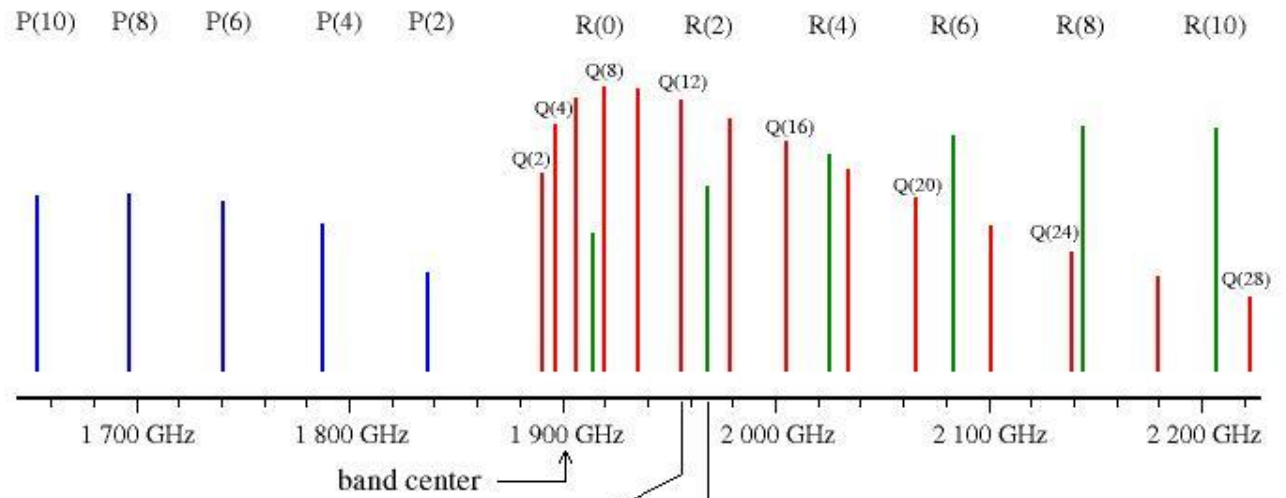
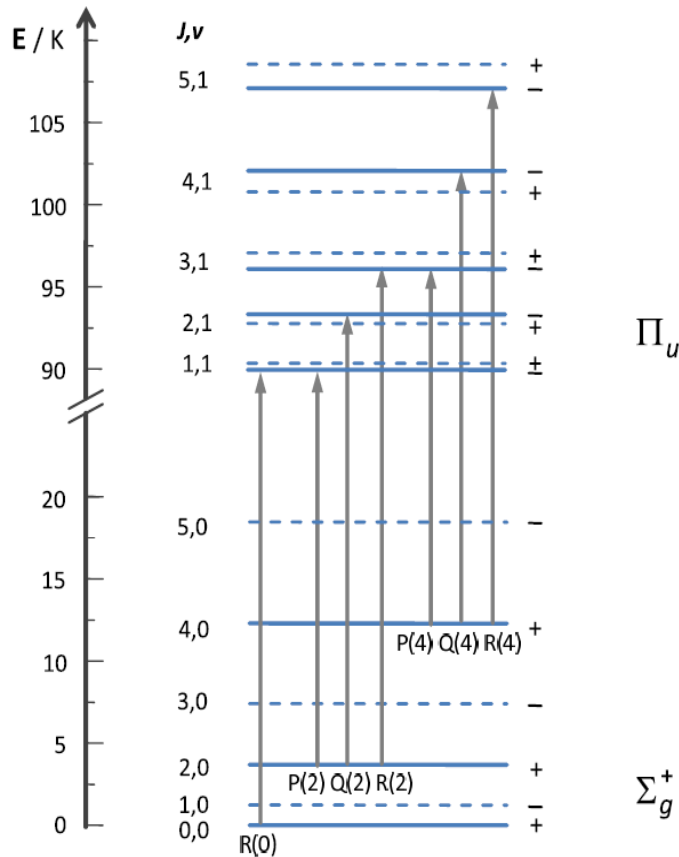
SOFIA GREAT HETERODYNE RECEIVERS





Lowest bending vibration of C₃ at 1.9 THz

COSSTA Terahertz-Sideband Spectrometer / Cologne



Selection Rules: $\Delta J = -1 \quad 0 \quad +1$
 $P(J) \quad Q(J) \quad R(J)$

Experimental setup at University of Kassel / Germany

Supersonic Jet Spectrometer for Terahertz Applications (SuJeSTA)

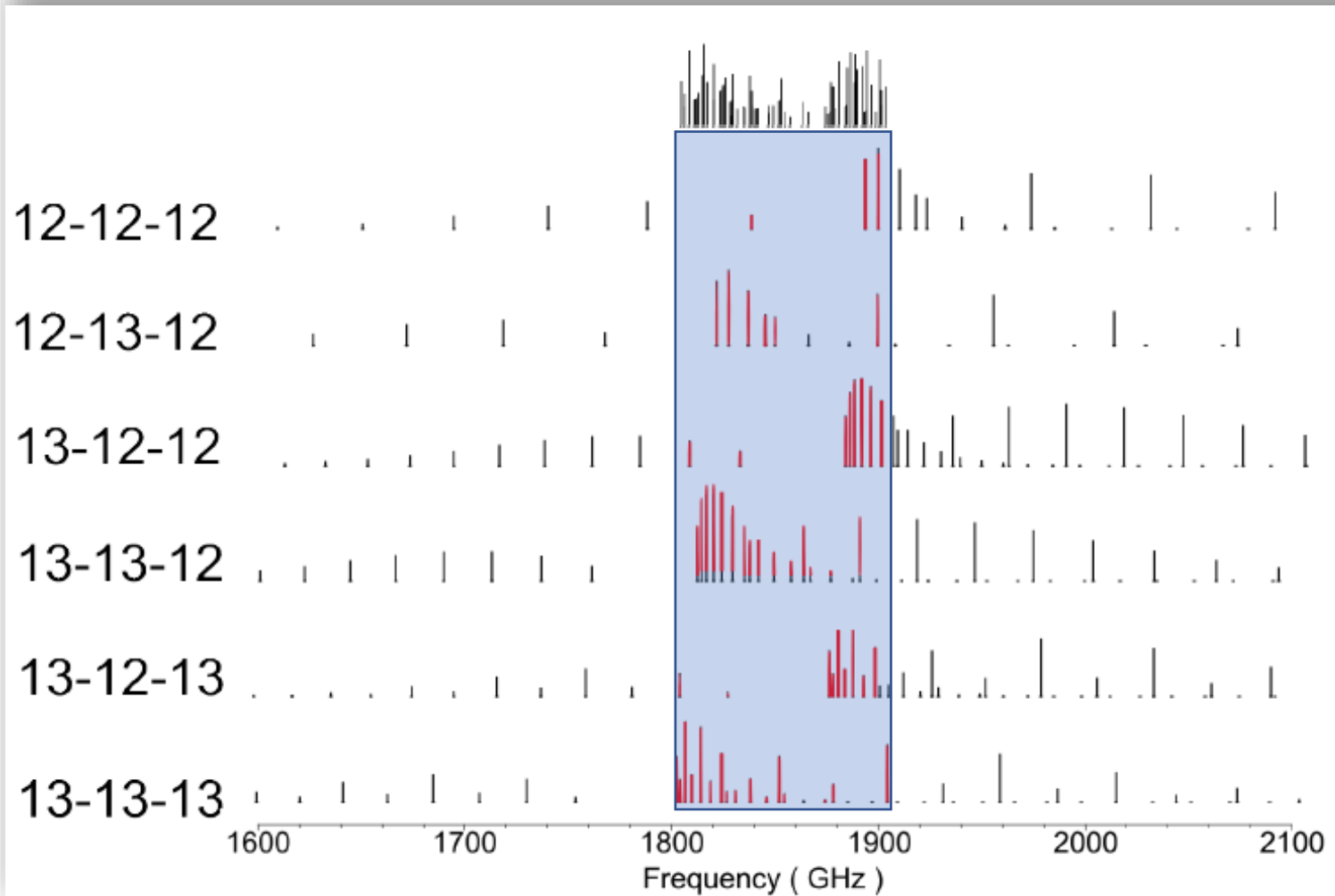
Reaction Chamber



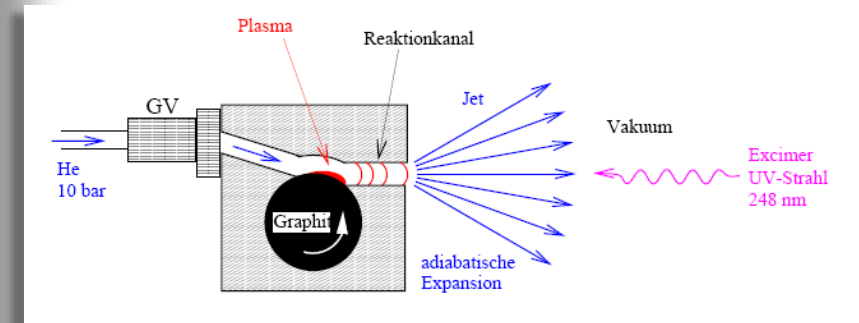
Lowest bending mode of ^{13}C -substituted C_3 and an experimentally derived structure

A.A. Breier, T. Büchling, R. Schnierer, V. Lutter, G.W. Fuchs, K.M.T. Yamada, B. Mookerjee, J. Stutzki, T.F. Giesen, J. Chem. Phys. 145, 23 (2016)

Laboratory Spectra of C_3 and its Isotopologues at 1.8 – 1.9 THz



Laser Ablation of Graphite



Supersonic Jet

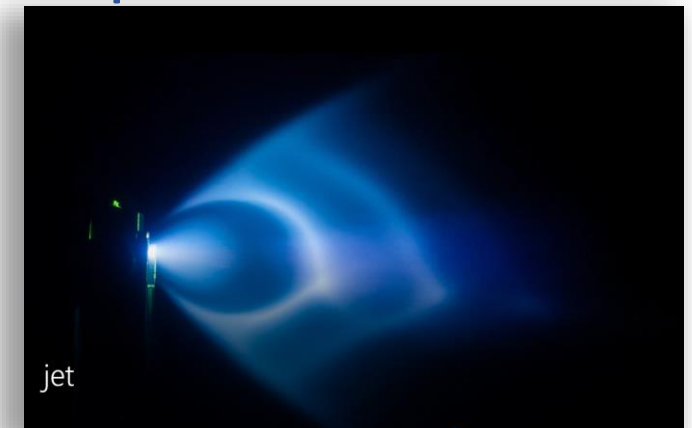
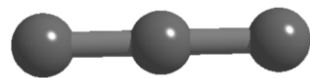
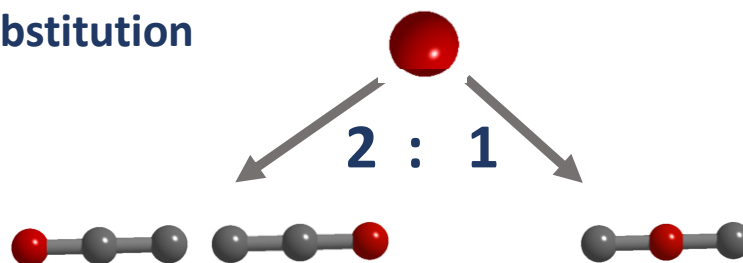


Photo by Björn Waßmuth

Abundances and Intensities of $^{12}\text{C}^{13}\text{C}^{12}\text{C}$ and $^{13}\text{C}^{12}\text{C}^{12}\text{C}$



^{13}C Substitution



Partitionfunction $Q^{13\text{C}^{12}\text{C}^{12}\text{C}} : Q^{^{12}\text{C}^{13}\text{C}^{12}\text{C}}$

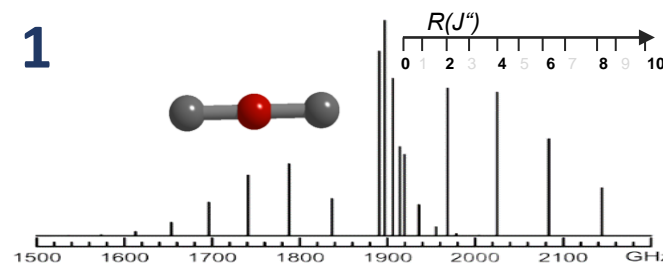
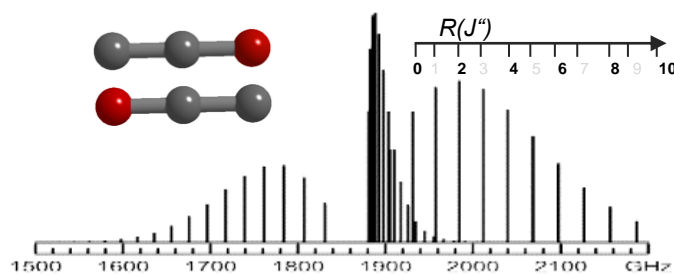
2 : 1

Intensity of individual lines

$$I \sim \frac{N}{Q}$$

Line Intensity Ratio $I^{13\text{C}^{12}\text{C}^{12}\text{C}} : I^{^{12}\text{C}^{13}\text{C}^{12}\text{C}}$

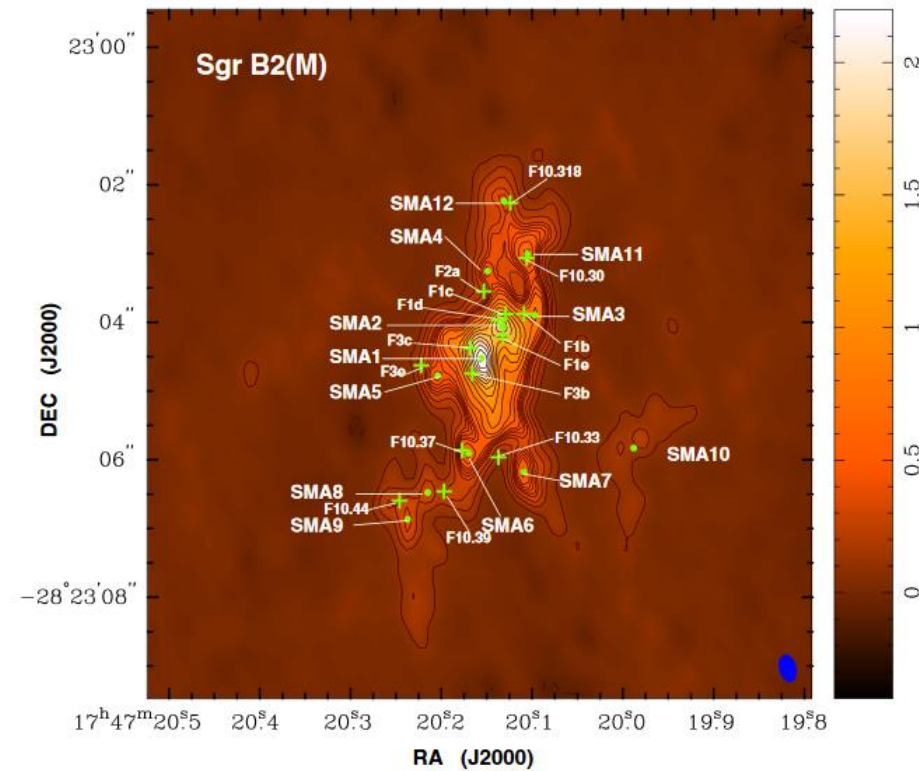
1 : 1



^{13}CCC and C^{13}CC observations with SOFIA GREAT and upGREAT

Observations started from New Zealand

Target : SgrB2(M)



Data Analysis

averaged $^{13}\text{CCC}_Q(4)$ spectrum:

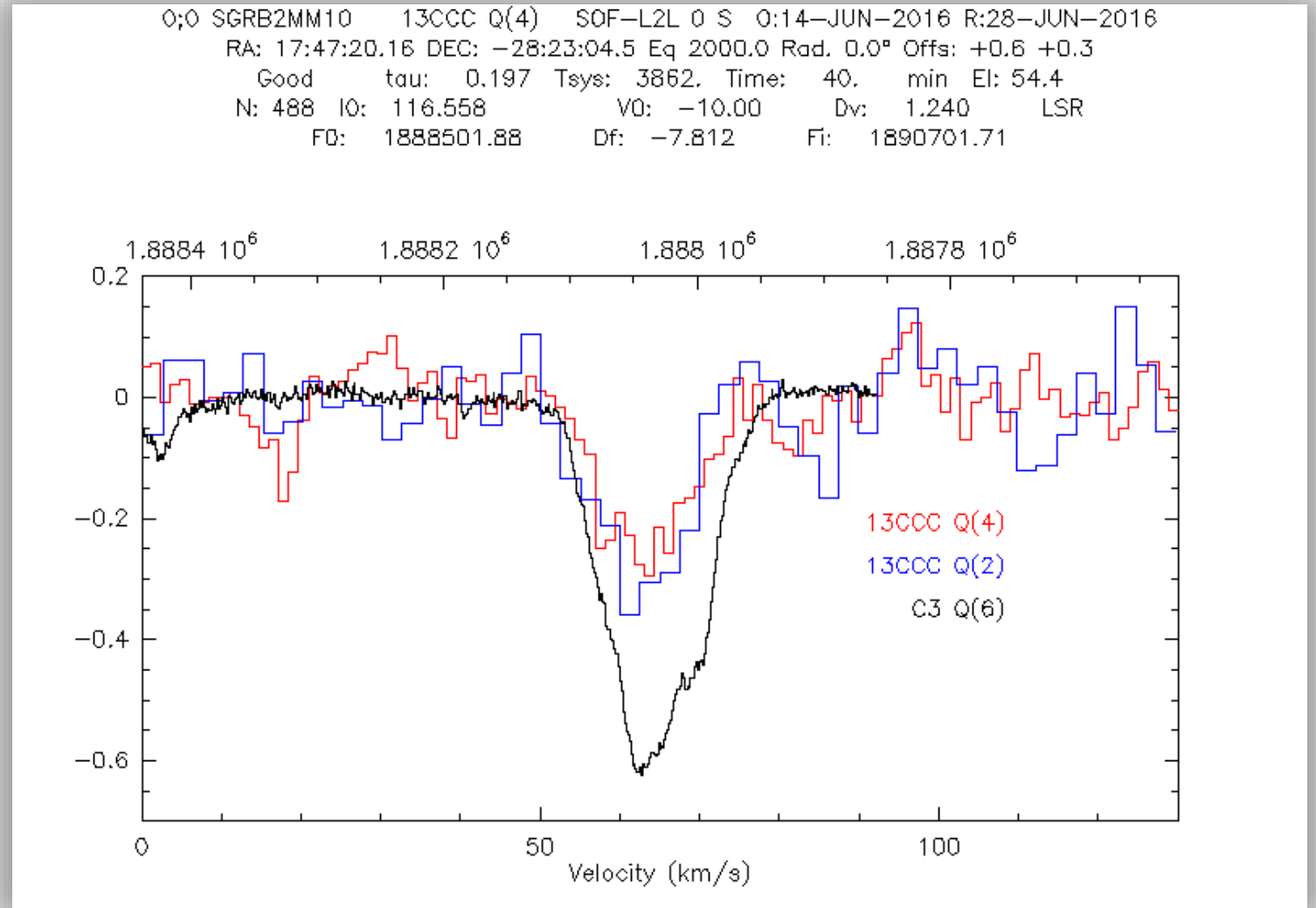
- 3 tunings from 2015
- 1 tuning from 2016,
corrected for sideband
absorption as derived
from 2015 tunings

$^{13}\text{CCC}_Q(2)$
CCC_Q(6)

- convincing detection of ^{13}CCC !
- two lines with consistent
profile
 - profile consistent with lower
velocity component of C₃
absorption

-2015: CCC_Q(6), $^{13}\text{CCC}_Q(2)$, $^{13}\text{CCC}_Q(4)$ (3 tunings) during 40 minutes in-flight time

- 2016: $^{13}\text{CCC}_Q(4)$ (one additional tuning) during 30 minutes in flight-time



First Interstellar Detection of the Carbon Chain Molecules ^{13}CCC and C^{13}CC toward SgrB2(M)

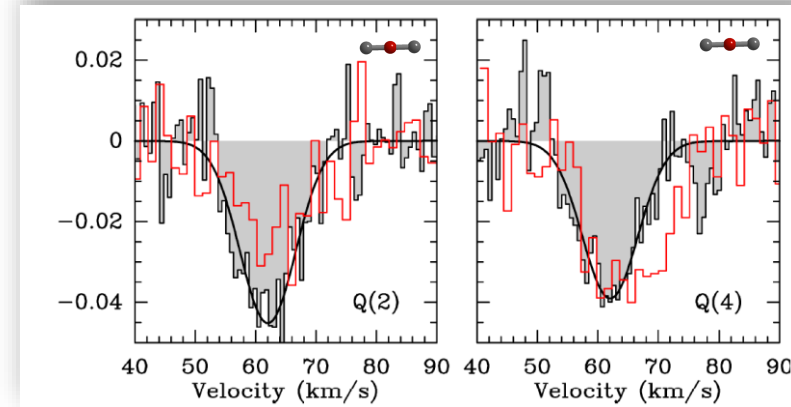
Parameters for Observed C^{13}CC and ^{13}CCC

Transition	Frequency MHz	Einstein A_{ul} s^{-1}	E_l K	g_l	g_u
C^{13}CC					
Q(2)	1819596.013	0.01309	3.7	5	5
Q(4)	1825647.312	0.01322	12.4	9	9
^{13}CCC					
Q(2)	1882638.269	0.01450	3.7	5	5
Q(4)	1888501.880	0.01463	12.4	9	9

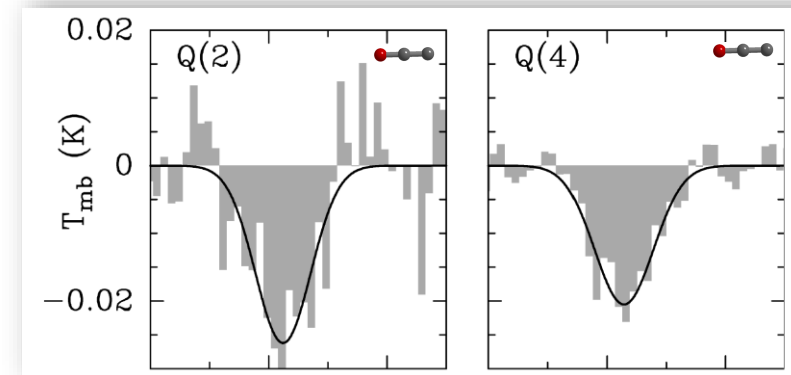
C^{13}CC and ^{13}CCC Observational Settings SOFIA/GREAT (G) and upGREAT (upG) Receivers

Transition	Rec.	Obs. Date	v_{off}^* km/s	T_{int} min	T_{rms} mK
^{13}CCC - Q(2)	G	19 July 2015	+10.0	6.8	99
^{13}CCC - Q(4)	G	19 July 2015	+0.0	27.9	62
	upG	9 June 2016	-10.0	12.5	96
C^{13}CC - Q(2)	upG	28 June 2017	0.0	5.7	112
	upG	28 June 2017	+10.0	5.1	150
C^{13}CC - Q(4)	upG	28 June 2017	0.0	8.5	56
	upG	28 June 2017	+10.0	2.8	150

* Velocity offset of LO setting relative to the v_{LSR}



Ro-vibrational spectra of C^{13}CC observed towards SgrB2(M)
H-polarization (filled spectrum), V-polarization (red).



Ro-vibrational spectra of ^{13}CCC observed towards SgrB2(M)

Ro-vibrational Transitions of the Main Isotopologue CCC

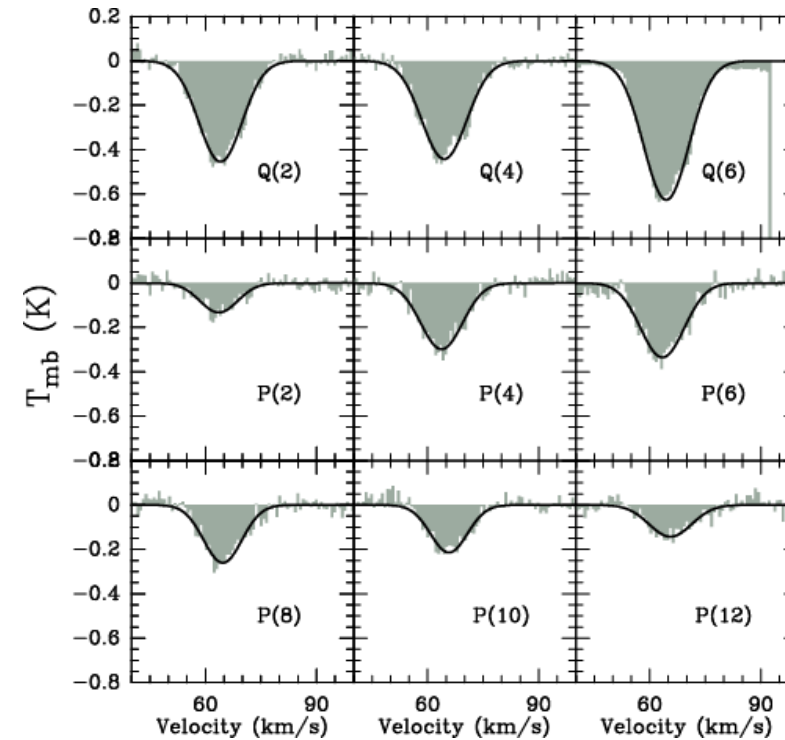
T. F. Giesen, B. Mookerjee, G. W. Fuchs, A. A. Breier, D. Witsch, R. Simon, J. Stutzki, A&A 2020, 633, A120



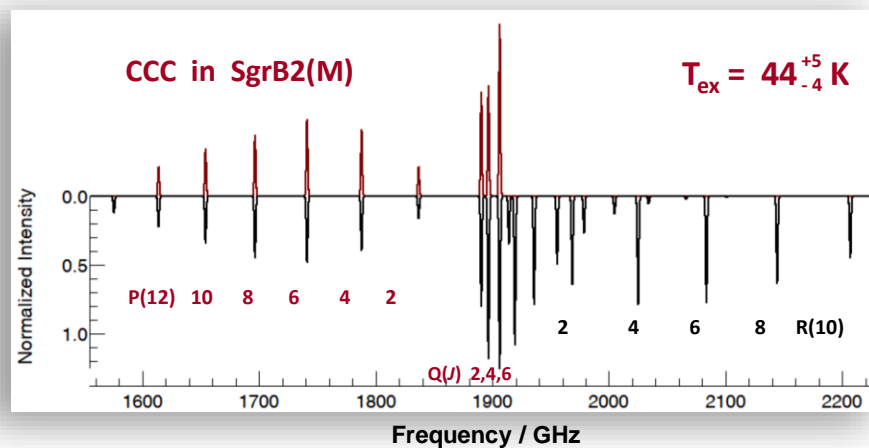
Parameters for the Observed CCC Transitions.

Transition	Frequency MHz	Einstein A_{ul} s^{-1}	E_l K	g_l	g_u
CCC					
Q(2)	1890558.188	0.01468	3.7	5	5
Q(4)	1896706.838	0.01482	12.4	9	9
Q(6)	1906337.907	0.01505	26.0	13	13
P(2)	1836823.502	0.00449	3.7	5	3
P(4)	1787890.534	0.00532	12.4	9	7
P(6)	1741122.646	0.00521	26.0	13	11
P(8)	1696525.363	0.00495	44.6	17	15
P(10)	1654087.900	0.00466	68.1	21	19
P(12)	1613805.250	0.00437	96.6	25	23

Observed CCC Transitions in SgrB2(M)



Temperature of CCC in SgrB2(M)



$$A_{ul} = \frac{16\pi^3 \nu^3}{3\epsilon_0 h c^3} \cdot |\mu_{ul}|^2 \quad \text{and} \quad |\mu_{ul}|^2 = \frac{|\mu_v|^2 L_{P/Q/R}(J)}{g_u}$$

Hönl-London factors for Lin. Mol. of $\Pi-\Sigma$ transitions:

$$L_p(J) = J-1, \quad L_Q(J) = 2J+1, \quad L_R(J) = J+2$$

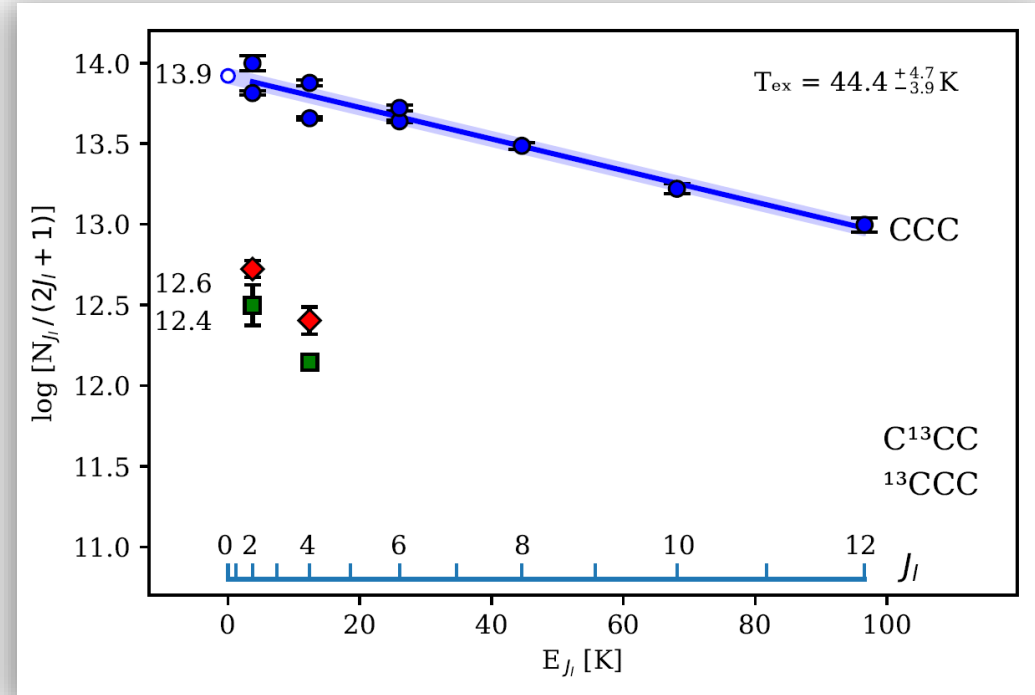
Data Analysis of C₃ CCC, ¹³CCC and C¹³CC

Table 3. Summary of observed and derived quantities for the CCC, C¹³CC, and ¹³CCC transitions.

	τ_{peak}	ν_{LSR} km s ⁻¹	$\Delta\nu$ km s ⁻¹	$\int \tau d\nu$ km s ⁻¹	$N_{J_l} \times 10^{13}$ cm ⁻²
CCC					
<i>Q</i> (2)	0.53	63.9(1)	11.6(2)	6.6(1)	32.6(5)
<i>Q</i> (4)	0.56	64.4(1)	13.7(2)	8.3(1)	40.9(5)
<i>Q</i> (6)	0.87	64.3(1)	12.5(1)	11.5(1)	56.6(5)
<i>P</i> (2)	0.15	63.5(3)	12.7(8)	2.0(1)	49.8(25)
<i>P</i> (4)	0.34	63.8(1)	12.5(3)	4.5(1)	67.8(15)
<i>P</i> (6)	0.39	63.4(1)	12.5(3)	5.2(1)	68.5(13)
<i>P</i> (8)	0.31	64.4(1)	13.0(4)	4.2(1)	52.1(12)
<i>P</i> (10)	0.24	65.7(2)	11.7(4)	2.9(1)	34.9(12)
<i>P</i> (12)	0.15	65.5(4)	13.0(8)	2.1(1)	24.8(12)
C¹³CC^(a)					
<i>Q</i> (2)	0.046	62.0(3)	10.8(7)	0.53(3)	2.64(15)
<i>Q</i> (4)	0.040	62.0(4)	10.8(10)	0.46(4)	2.28(20)
¹³CCC					
<i>Q</i> (2)	0.027	62.8(8)	11.1(15)	0.32(4)	1.58(20)
<i>Q</i> (4)	0.021	62.9(3)	11.6(6)	0.26(1)	1.26(5)

Notes. ^(a)Only H-polarization data was used.

Rotational Diagram of C₃ CCC, ¹³CCC and C¹³CC



$$N_{J_l} = \frac{8\pi\nu^3}{c^3} \frac{g_l}{A_{ul} g_u} \left[1 - \exp\left(-\frac{h\nu}{k_B T_{\text{ex}}}\right) \right]^{-1} \int \tau d\nu \quad \text{Eq (1)}$$

Data Analysis of C₃ CCC, ¹³CCC and C¹³CC

Table 3. Summary of observed and derived quantities for the CCC, C¹³CC, and ¹³CCC transitions.

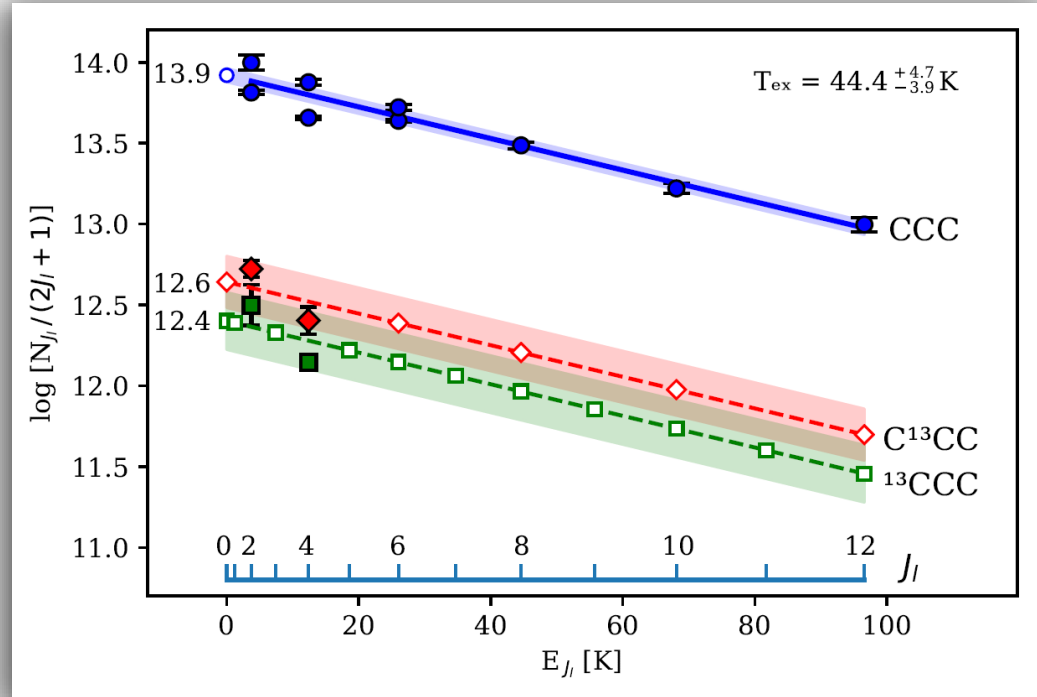
	τ_{peak}	ν_{LSR} km s ⁻¹	$\Delta\nu$ km s ⁻¹	$\int \tau d\nu$ km s ⁻¹	$N_{J_1} \times 10^{13}$ cm ⁻²
CCC					
<i>Q</i> (2)	0.53	63.9(1)	11.6(2)	6.6(1)	32.6(5)
<i>Q</i> (4)	0.56	64.4(1)	13.7(2)	8.3(1)	40.9(5)
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¹³CCC					
<i>Q</i> (2)	0.027	62.8(8)	11.1(15)	0.32(4)	1.58(20)
<i>Q</i> (4)	0.021	62.9(3)	11.6(6)	0.26(1)	1.26(5)

Notes. ^(a)Only H-polarization data was used.

$$N_{J_1} = \frac{8\pi\nu^3}{c^3} \frac{g_l}{A_{ul} g_u} \left[1 - \exp\left(-\frac{h\nu}{k_B T_{\text{ex}}}\right) \right]^{-1} \int \tau d\nu$$

Eq (1)

Rotational Diagram of C₃ CCC, ¹³CCC and C¹³CC



	T_{ex} (K)	$\log(N_0)$	Q_{rv}	$N \times 10^{14}$ cm ⁻²	rel.
CCC	$44.4^{+4.7}_{-3.9}$	13.92(04)	46.64	$38.8^{+3.9}_{-3.5}$	100%
C ¹³ CC	44.4	12.64(16)	47.66	$2.1^{+0.9}_{-0.6}$	$5 \pm 2\%$
¹³ CCC	44.4	12.40(18)	97.30	$2.4^{+1.2}_{-0.8}$	$6 \pm 3\%$

$$Q_{\text{rv}} = \sum_{J_1(v=0,1,2)} (2J_1 + 1) g_v \exp\left(\frac{-E_{\text{rv}}(J_1, v)}{k_B T_{\text{ex}}}\right)$$

Eq (2)

Results for $^{12}\text{C}/^{13}\text{C}$ Ratios

	$T_{\text{ex}}(\text{K})$	$\log(N_0)$	Q_{rv}	$N \times 10^{14}$ cm^{-2}	rel.
CCC	$44.4^{+4.7}_{-3.9}$	13.92(04)	46.64	$38.8^{+3.9}_{-3.5}$	100%
C^{13}CC	44.4	12.64(16)	47.66	$2.1^{+0.9}_{-0.6}$	$5 \pm 2\%$
^{13}CCC	44.4	12.40(18)	97.30	$2.4^{+1.2}_{-0.8}$	$6 \pm 3\%$

$$N = Q_{\text{rv}} \frac{N_{J_1}}{2J_1 + 1} \exp\left(\frac{E_{\text{rot}}(J_1)}{k_B T_{\text{ex}}}\right) \quad \text{Eq (3)}$$

MEASURED RATIO		
$N(\text{CCC})$	$N(^{13}\text{CCC})$	$N(\text{C}^{13}\text{CC})$
18.6	1.2	1.0

EXPECTED RATIO		
$N(\text{CCC})$	$N(^{13}\text{CCC})$	$N(\text{C}^{13}\text{CC})$
20	2.0	1.0

MEASURED RATIO	
$\frac{^{12}\text{C}}{^{13}\text{C}}$	$= \frac{3N(\text{CCC})}{N(^{13}\text{CCC}) + N(\text{C}^{13}\text{CC})} = 25.8$

EXPECTED RATIO	
$\frac{^{12}\text{C}}{^{13}\text{C}}$	$= 20$

Total vs. Line by Line Ratio of Column Densities

N^a/N^b	$\frac{N(\text{CCC})}{N(\text{C}^{13}\text{CC})}$	$\frac{N(\text{CCC})}{N(^{13}\text{CCC})}$	$\frac{N(^{13}\text{CCC})}{N(\text{C}^{13}\text{CC})}$	$\frac{^{12}\text{C}}{^{13}\text{C}}$
TOTAL	18.6(7.0)	16.0(6.7)	1.2(0.6)	25.8(7.5)
$Q(2)$	12.1(0.7)	9.9(1.3)	1.2(0.2)	16.3(1.2)
$Q(4)$	17.6(1.6)	15.6(0.6)	1.1(0.1)	24.7(1.2)
AV	14.8(2.7) ^(a)	12.7(2.8) ^(a)	1.2(0.1) ^(a)	20.5(4.2) ^(a)
EXPEC	20	10	2	20

Line by Line Ratio of Column Density

$$\frac{N(a)}{N(b)} = \frac{Q_{\text{rv}}^a N_{J_1}^a (2J_1^b + 1)}{Q_{\text{rv}}^b N_{J_1}^b (2J_1^a + 1)} \cdot \exp\left(\frac{\Delta E_{\text{rot}}}{k_B T_{\text{ex}}}\right) \quad \text{Eq (4)}$$

Exchange and Rearrangement Reactions



Equilibrium Reaction Rate

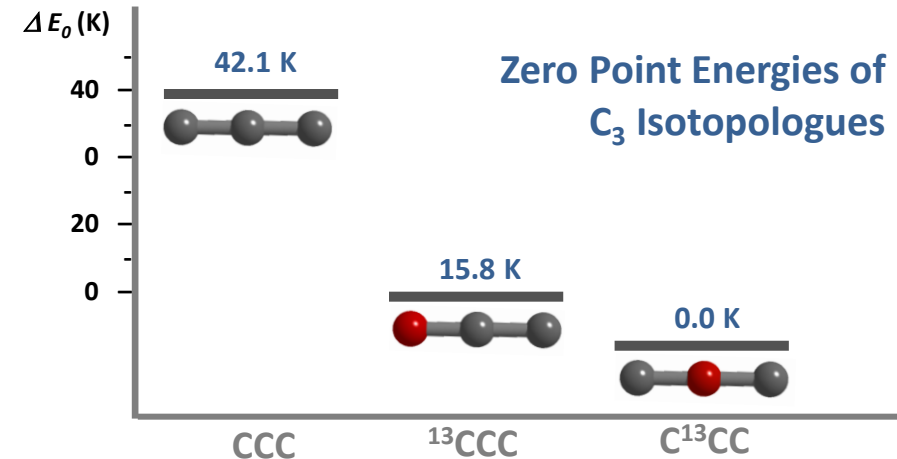
$$k_p = \exp\left(\frac{-\Delta G}{k_B T_{\text{ex}}}\right) = \frac{Q^{^{13}\text{CCC}}}{Q^{\text{C}^{13}\text{CC}}} \exp\left(\frac{-\Delta E_0}{k_B T_{\text{ex}}}\right) \quad \text{Eq (6)}$$

ΔG = Difference in Gibbs energy

ΔE_0 = Difference in zero point energy

Low Temperature Equilibrium at 44 K

$$\frac{N(^{13}\text{CCC})}{N(\text{C}^{13}\text{CC})} = \frac{1.4}{1} \quad \text{measured: } \frac{1.2}{1}$$



- At low temperatures of 44 K the $N(^{13}\text{CCC})/N(\text{C}^{13}\text{CC})$ ratio shifts from 2.0 to 1.4, which is in much better agreement with the measured ratio of 1.2(1)
- The $^{12}\text{C}/^{13}\text{C}$ ratio in SgrB2(M) derived from C_3 isotopologues is 20.5 +/- 4.2, which is in good agreement with the result of 20 derived from earlier measurements
- Need of further $^{13}\text{C}_3$ measurements with SOFIA in other astronomical sources, including higher J -transitions